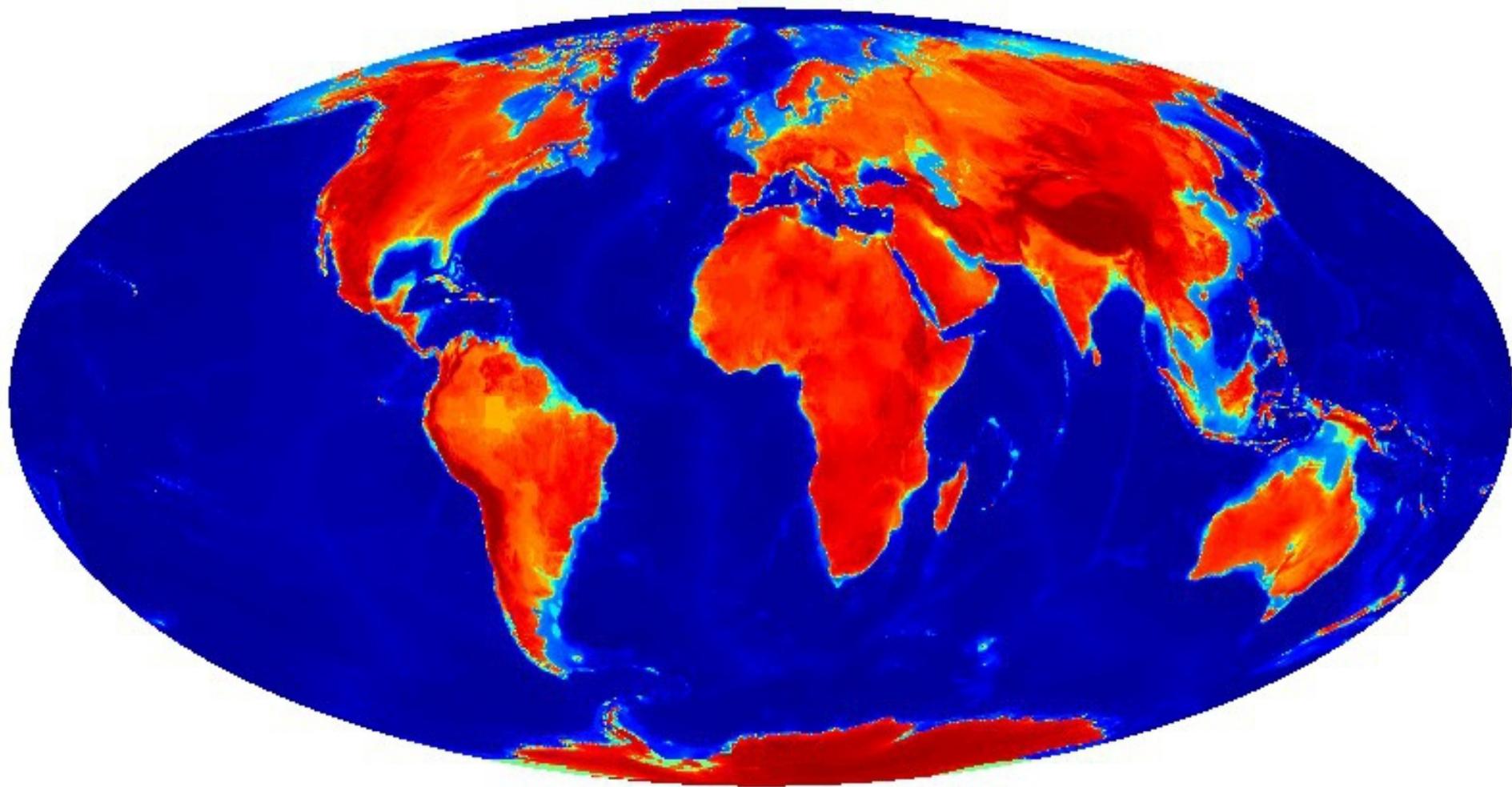


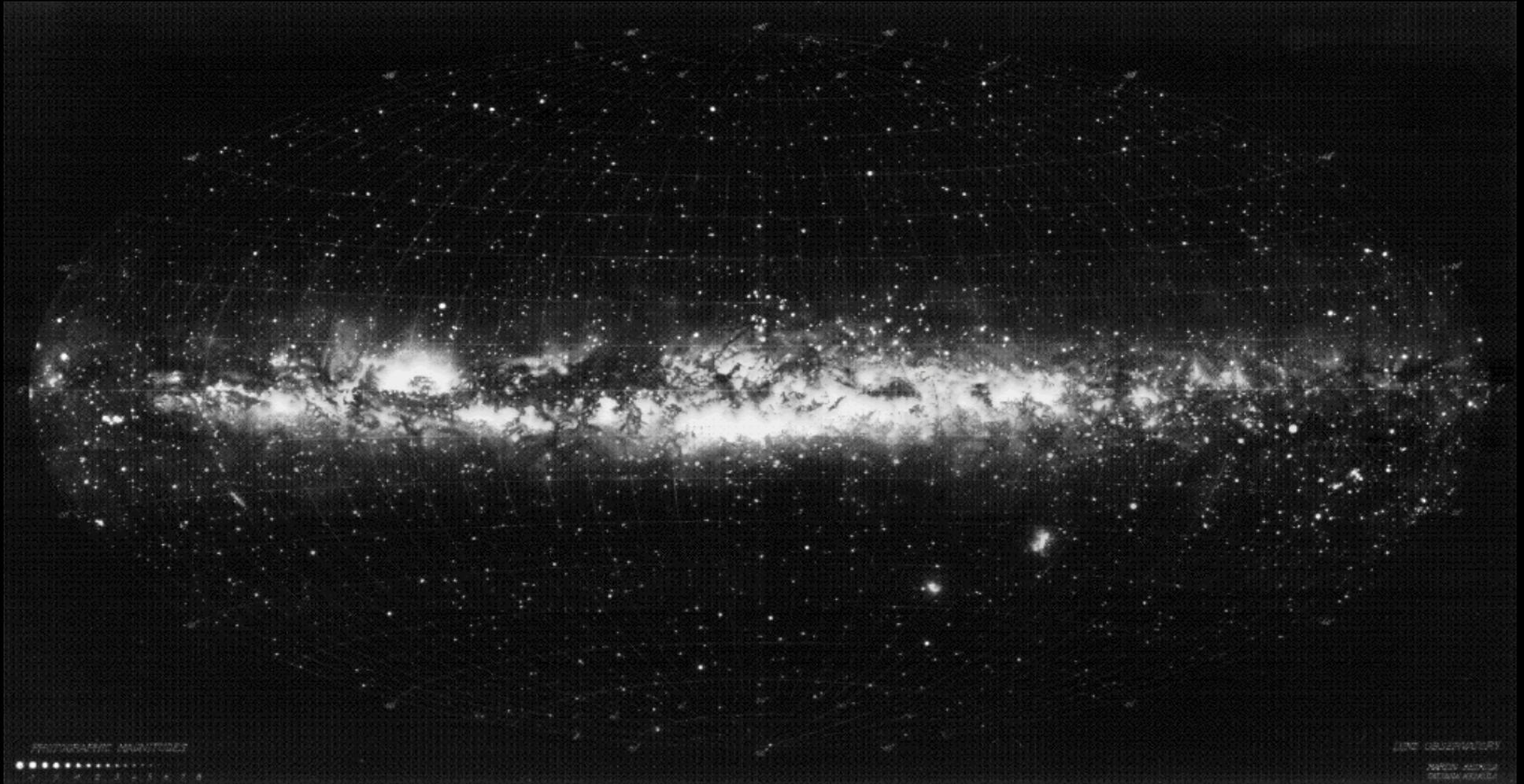
# All from Nothing: the structuring of our Universe

*Simon White*

*Max Planck Institute for Astrophysics*

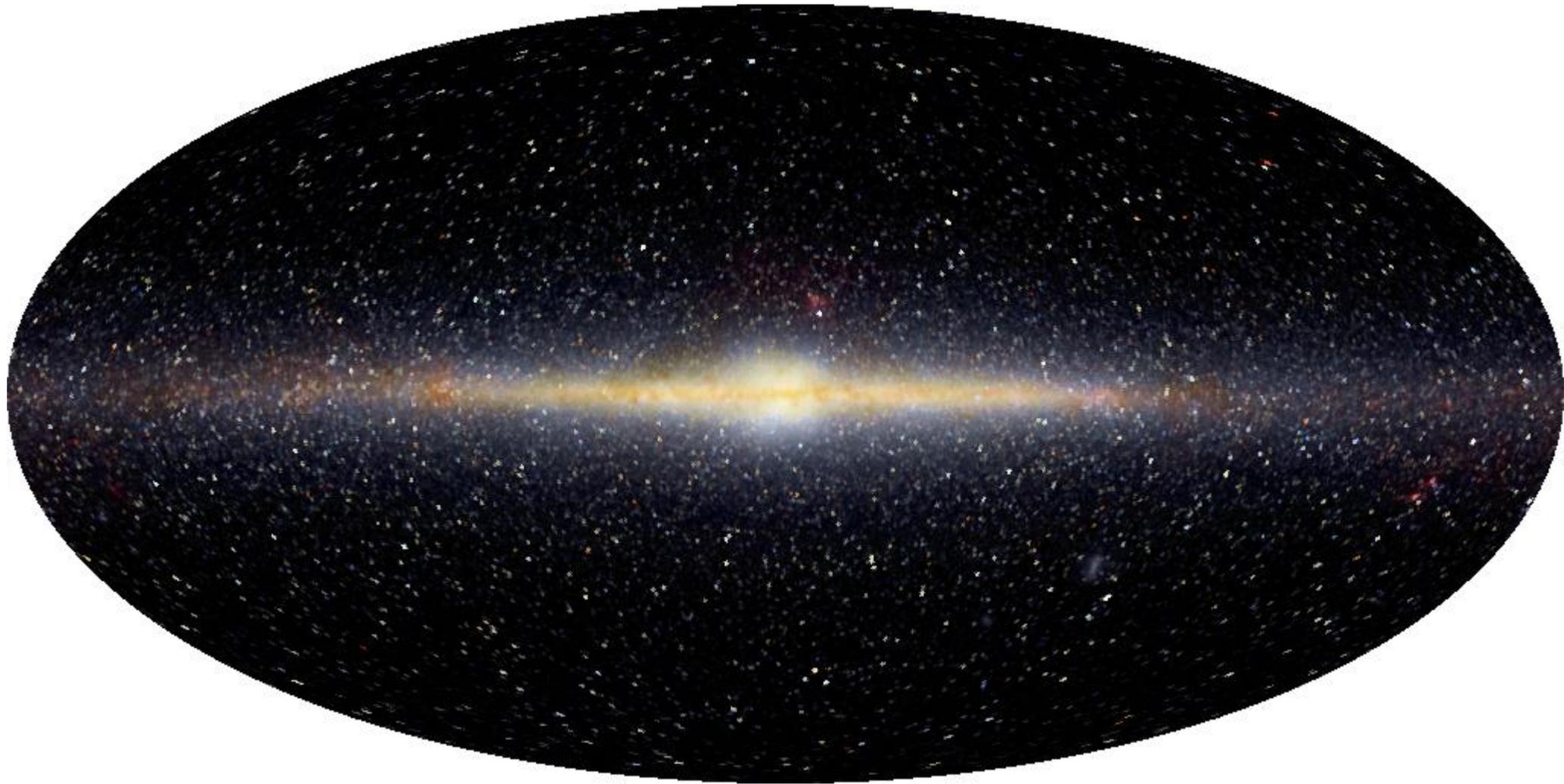


# Star map of the whole sky



..out to 10,000 light-years

# COBE's infrared map of the whole sky



..out to 30,000 light-years

# The Andromeda Nebula: our biggest neighbour



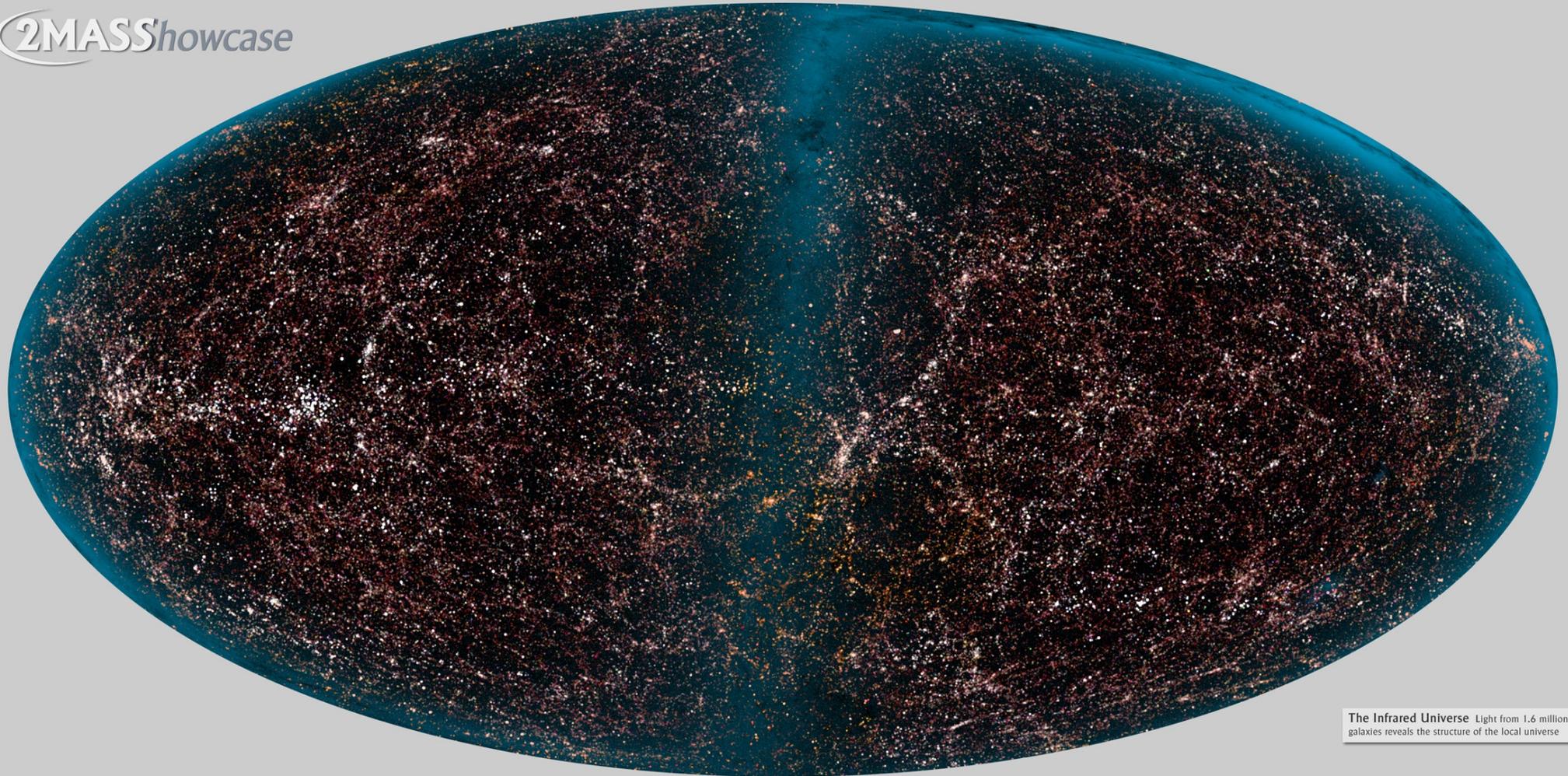
..out to 2,000,000 light-years

**NGC 4414 -- another  
Galaxy like our own**



# A galaxy map of the whole sky

2MASS Showcase



The Infrared Universe Light from 1.6 million galaxies reveals the structure of the local universe

Two Micron All Sky Survey Image Mosaic: Infrared Processing and Analysis Center/Caltech & University of Massachusetts

..out 1,000,000,000 light-years



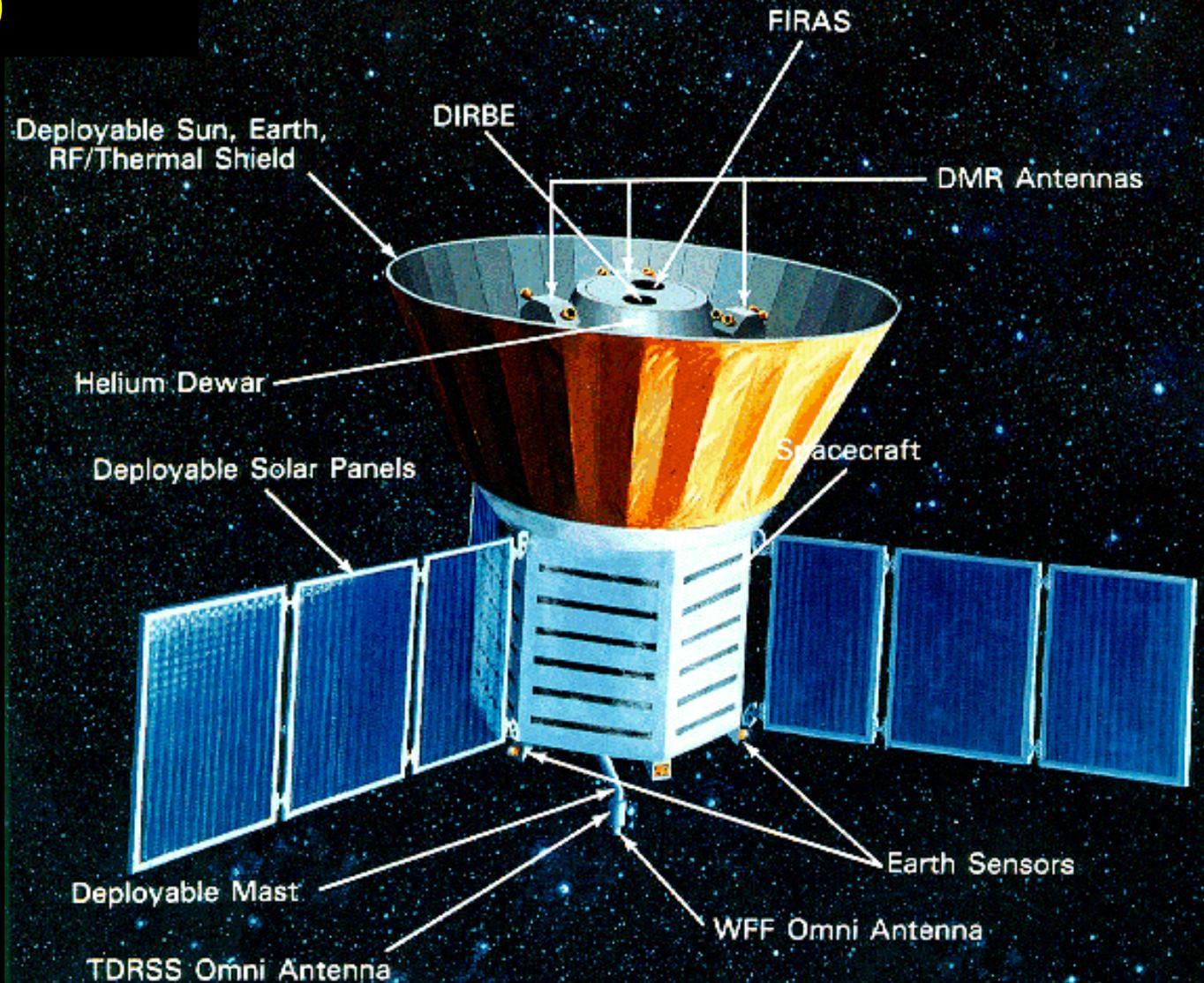
The deepest  
optical image  
ever made

A 300 hour  
exposure with  
the Hubble  
Space Telescope

..out to more than 30,000,000,000 light-years

# The COBE Satellite (1989 - 1993)

- Two Instruments mapped the whole sky at microwave and infrared wavelengths
- One instrument took a precise spectrum of the sky at microwave wavelengths



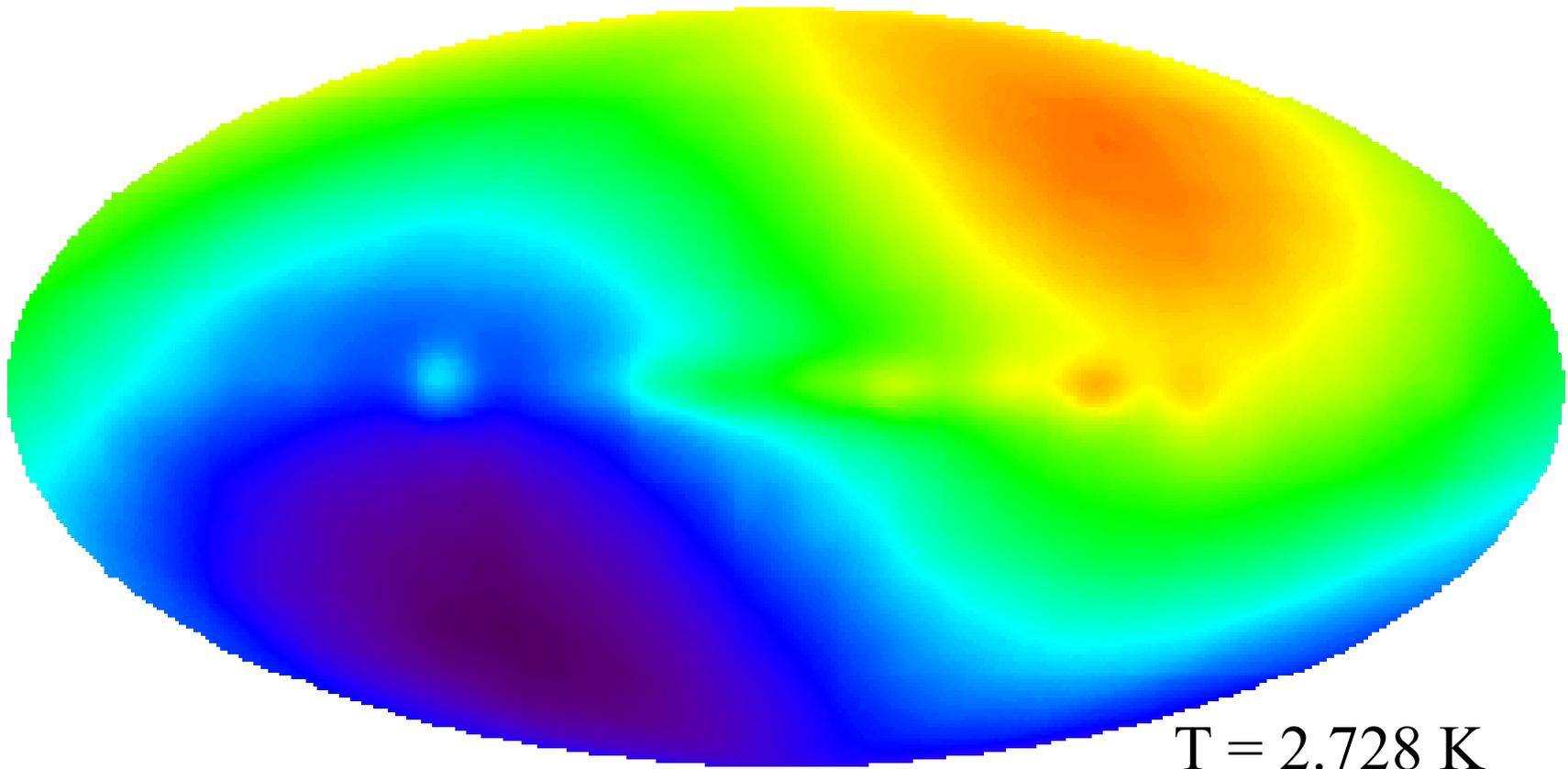
Nobel Prize 2006

# COBE's temperature map of the whole sky



$T = 2.728 \text{ K}$   
 $\Delta T = 0.1 \text{ K}$

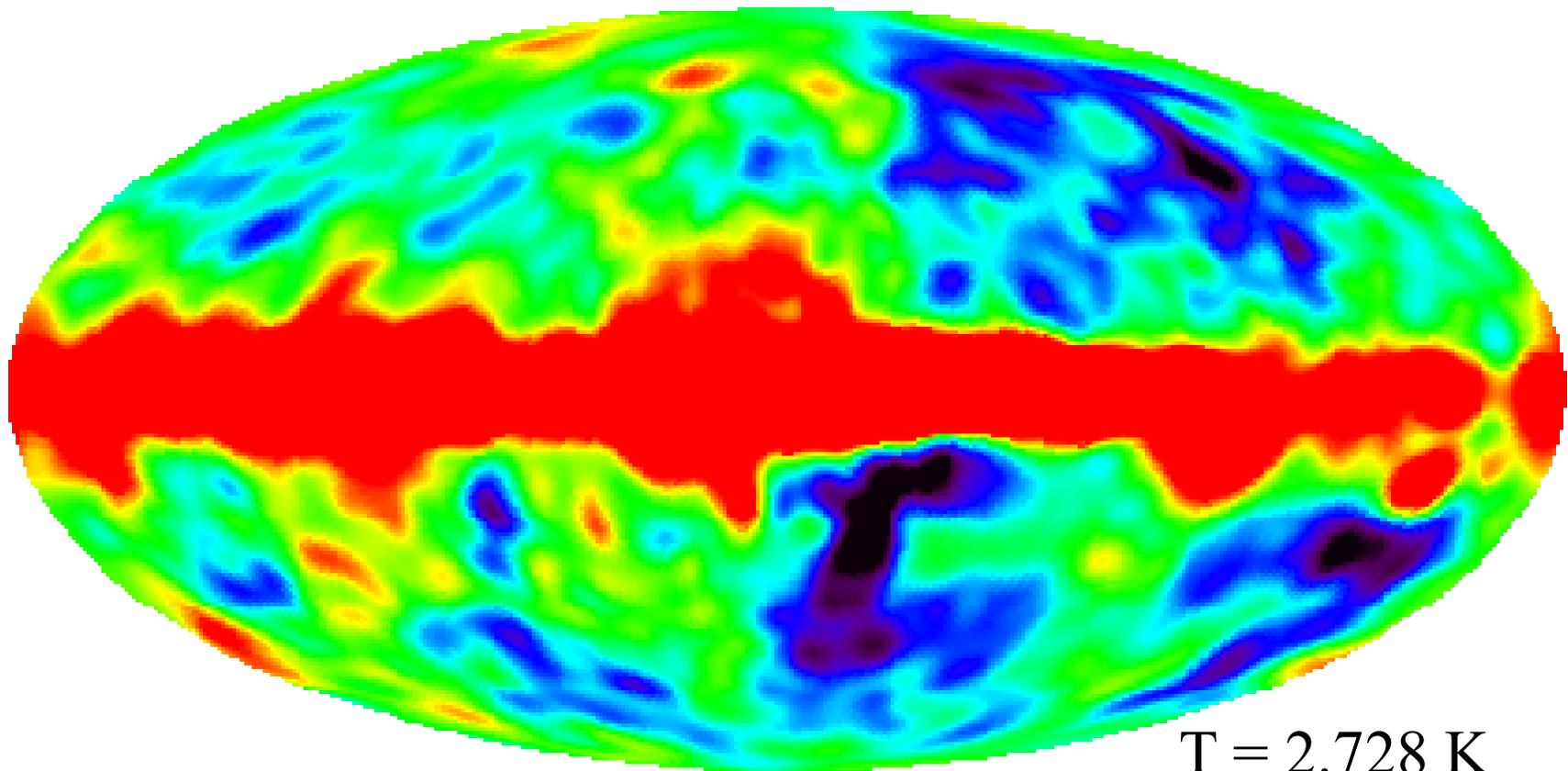
# COBE's temperature map of the whole sky



$T = 2.728 \text{ K}$

$\Delta T = 0.0034 \text{ K}$

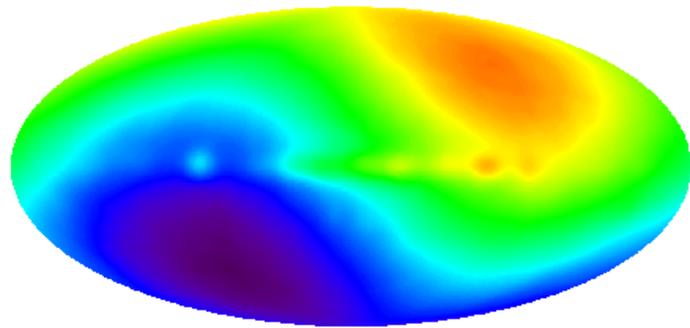
# COBE's temperature map of the whole sky



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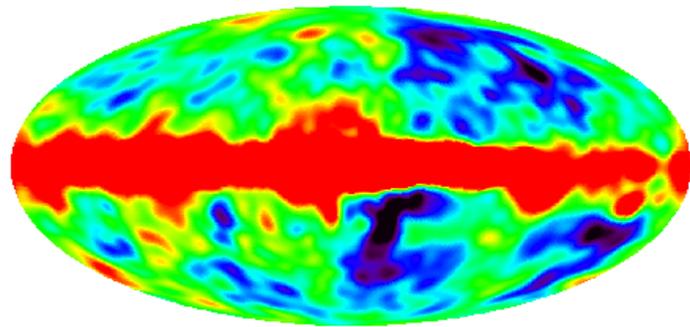
$\Delta T = 0.00002 \text{ K}$

# Structure in the COBE Map

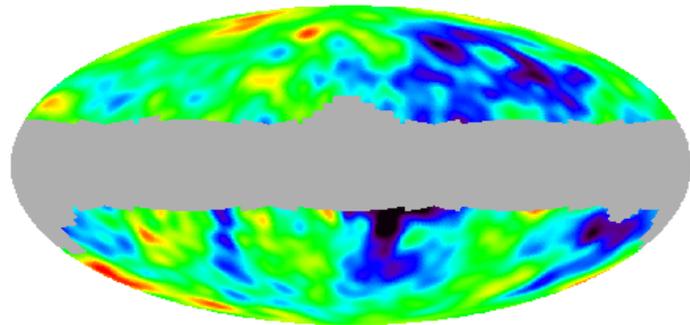


- One side of the sky is cold, the other hot  
our Galaxy's motion through the cosmos

→  $V_{\text{Milchstrasse}} = 600 \text{ km/s}$



- Radiation from dust and gas in our own Galaxy



- Structure in the microwave background itself

# Structure in the microwave background

Where is the structure?

# Structure in the microwave background

Where is the structure?

In cosmic clouds at the far edge of the visible Universe

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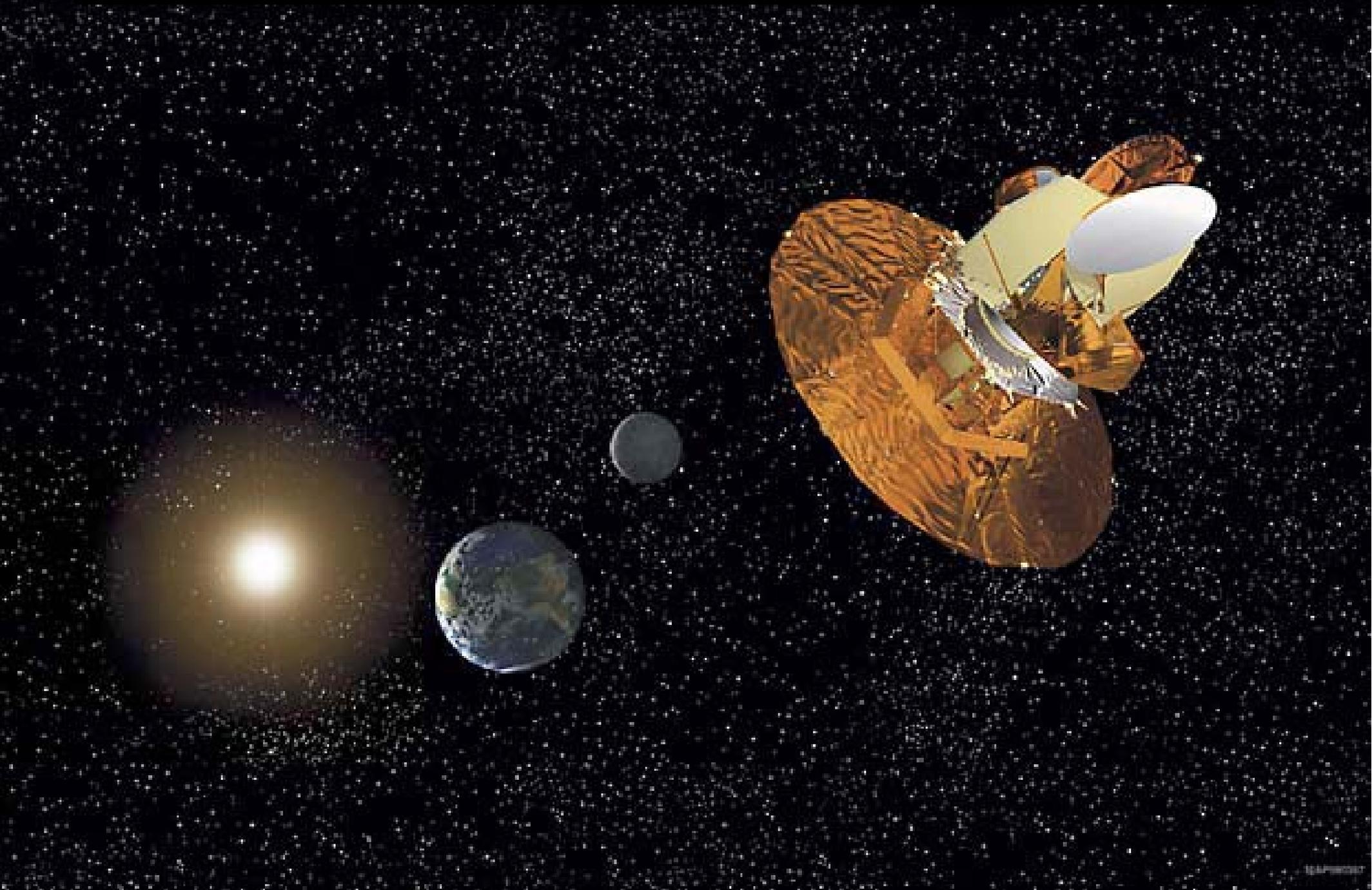
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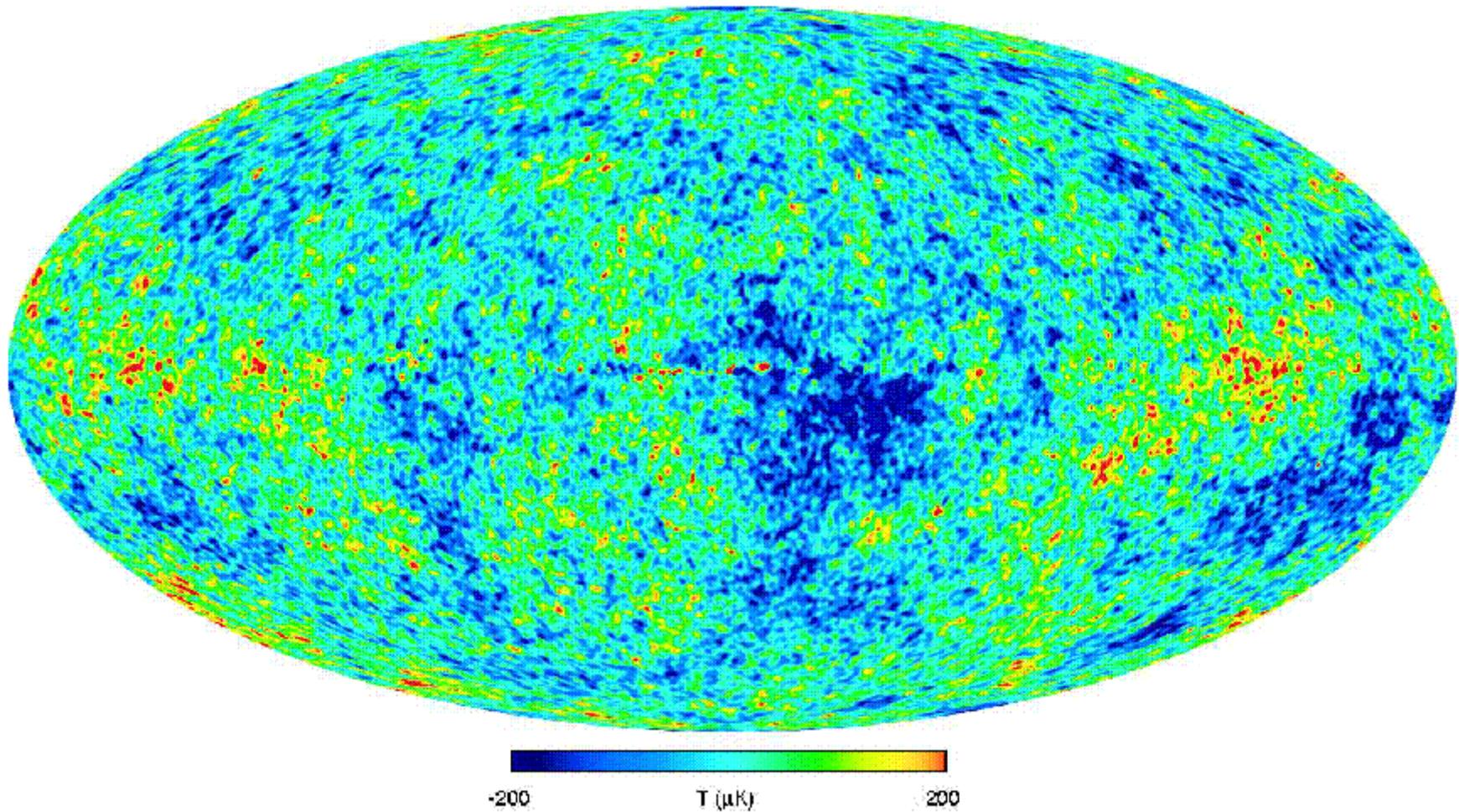
What has this structure become?

Everything we see around us (galaxies, stars, planets, people...)

# The *WMAP* satellite at Lagrange point L2



# *WMAP's* map of the whole sky



Bennett et al 2003

# What do we learn from these structures?

The pattern of structure is influenced by three things:

--the geometry of the Universe

--the content of the Universe

--the process which created the structure

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Dark (non-baryonic) Matter

Dark Energy

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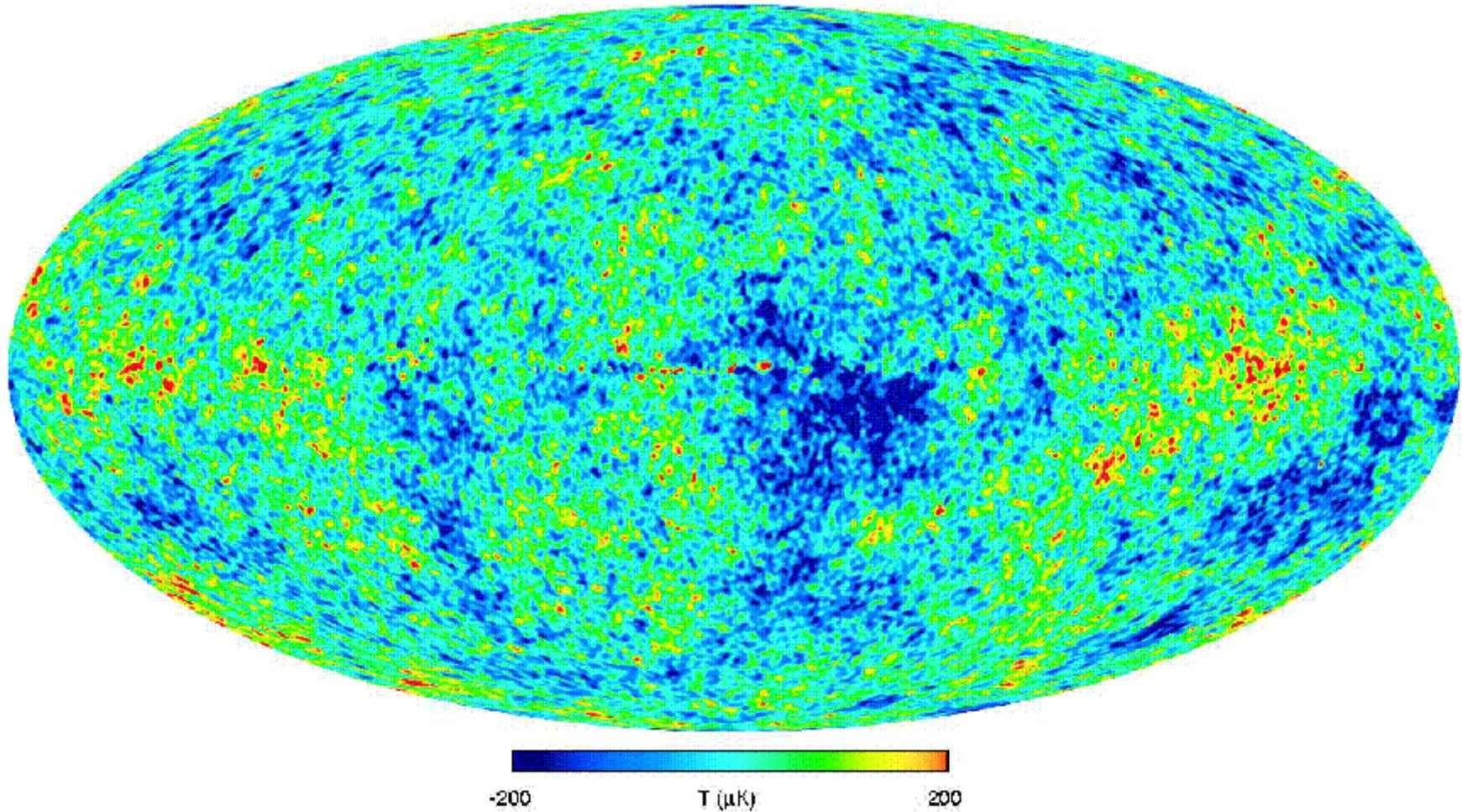
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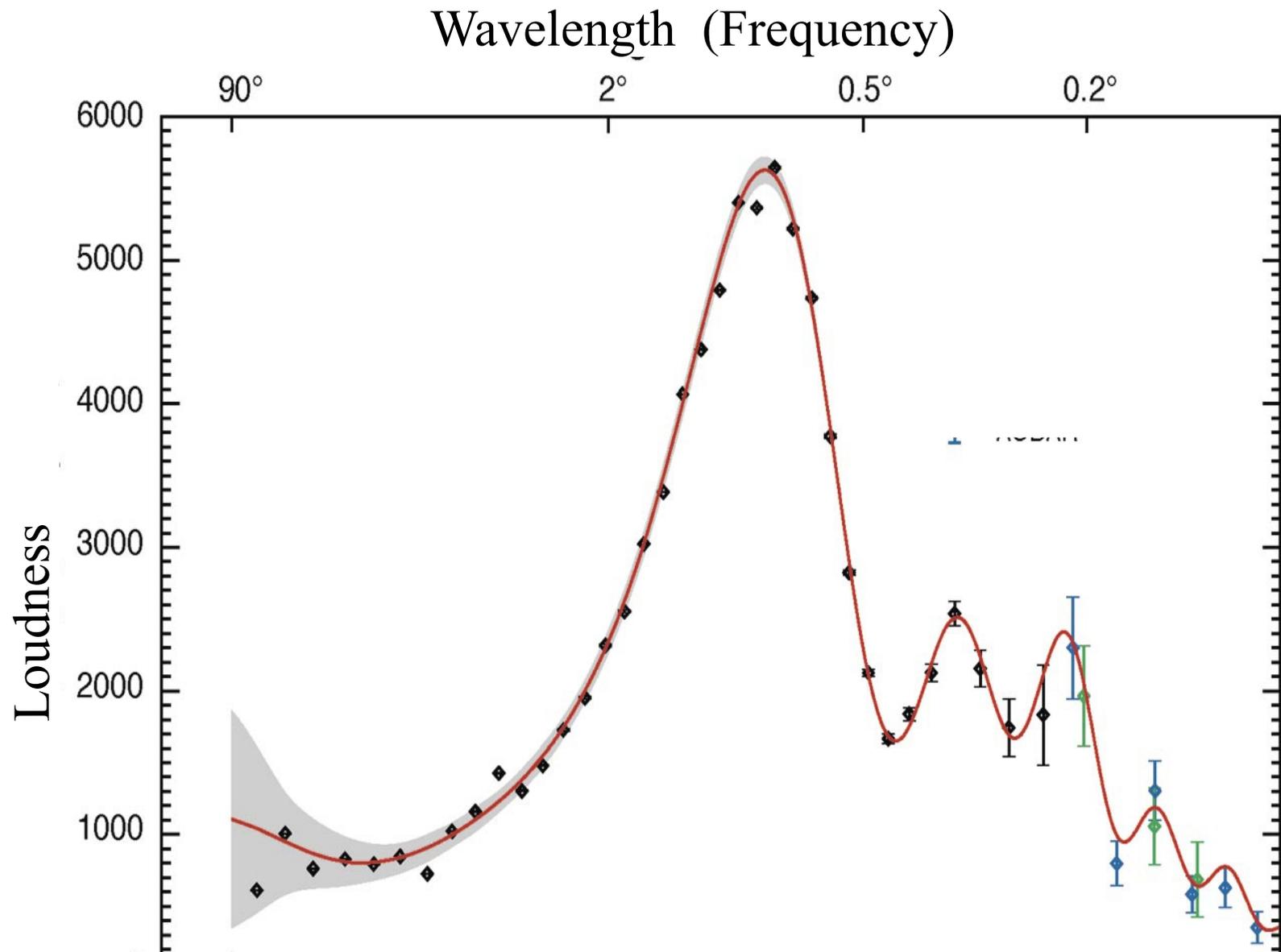
--the process which created the structure

The earliest instants of creation

# *WMAP's* map of the whole sky



Bennett et al 2003



The harmonic content of sound waves in the cosmic clouds:  
*WMAP* measurements compared to a theoretical prediction

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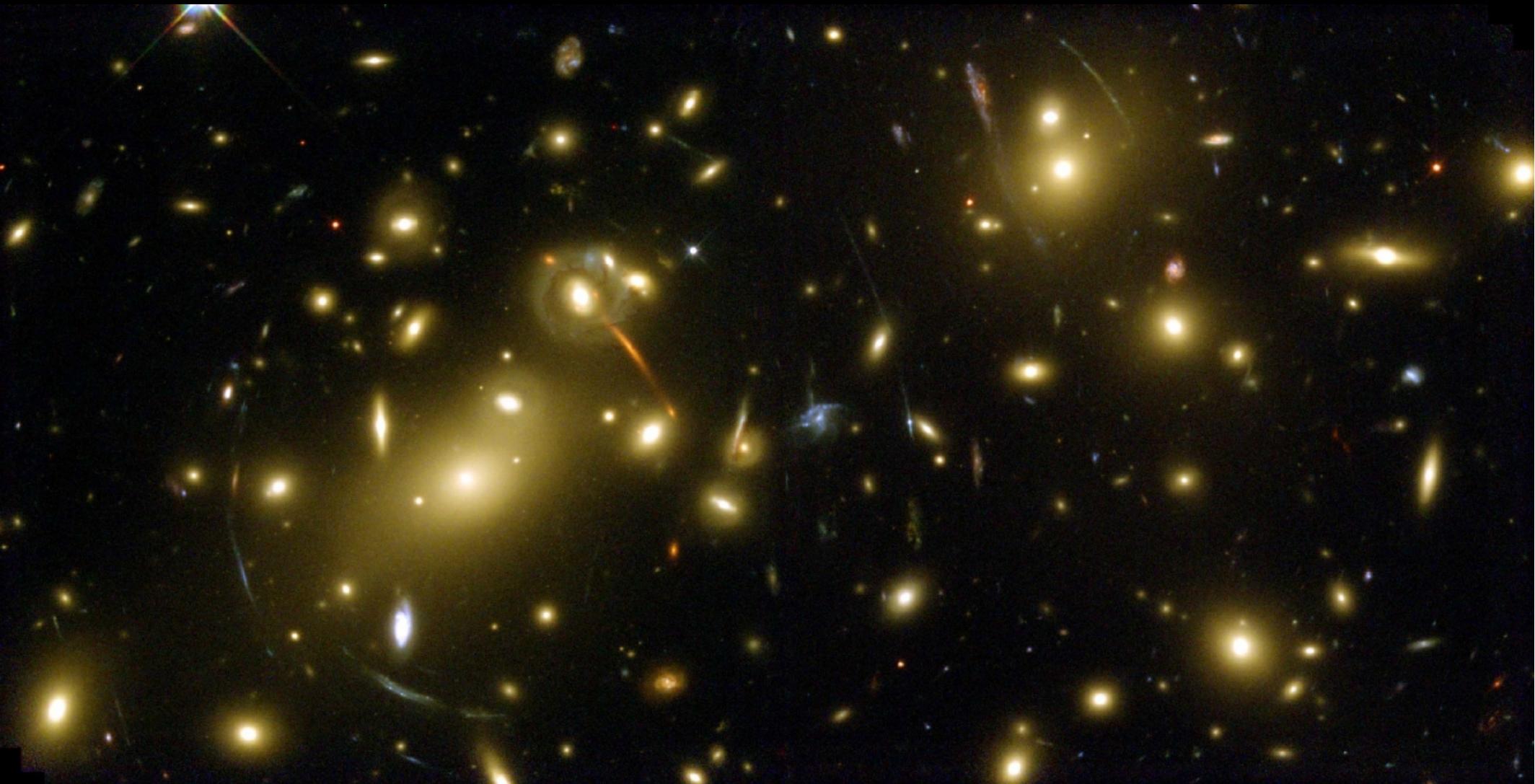
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(already “seen” through gravitational lensing effects)

# A galaxy cluster acting as a gravitational lens

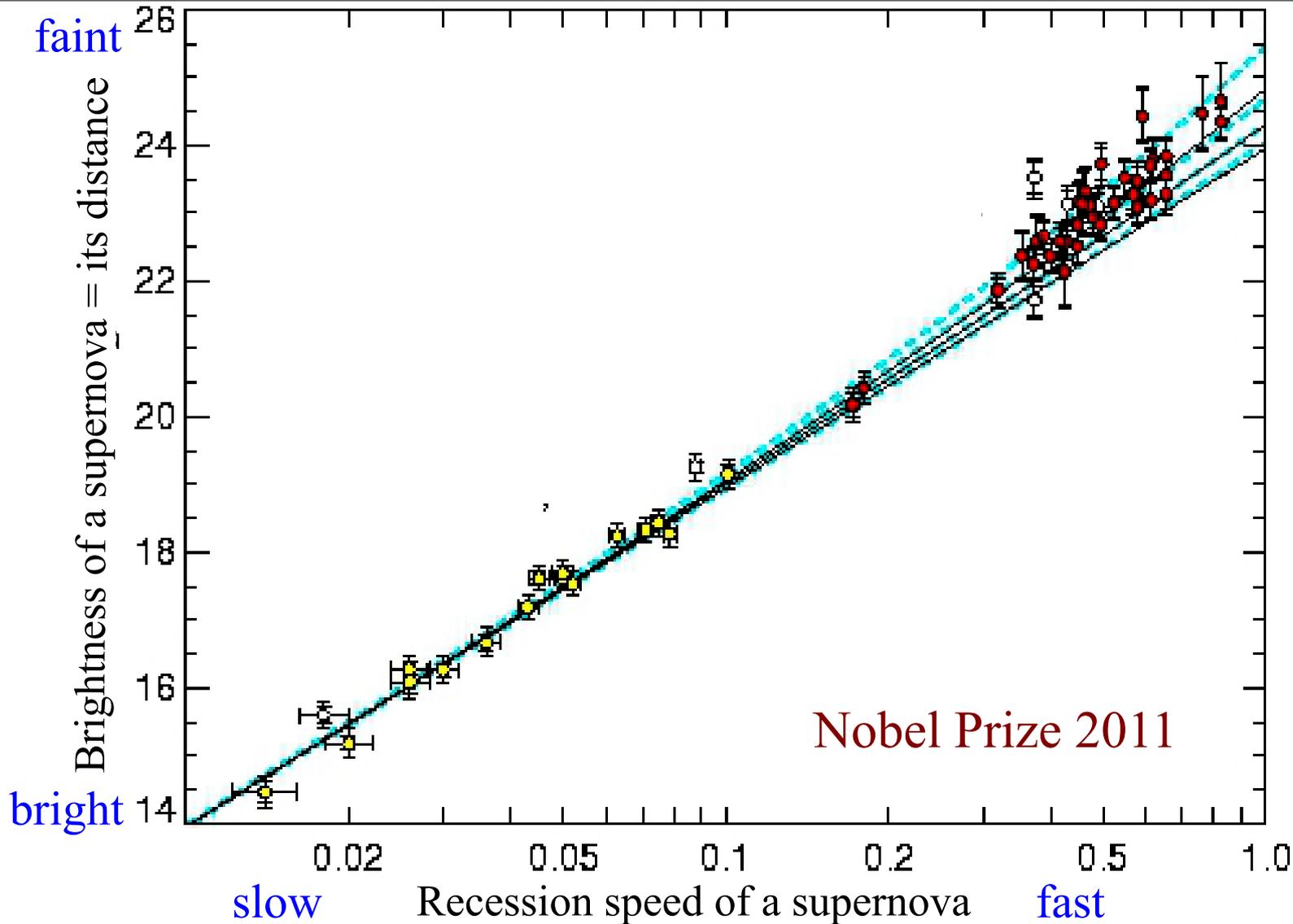
Abell 2218



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# Dark Energy drives cosmic acceleration



Distant supernovae, seen when the Universe was younger, are fainter than expected —→ the Universe expands **faster** today than in the past

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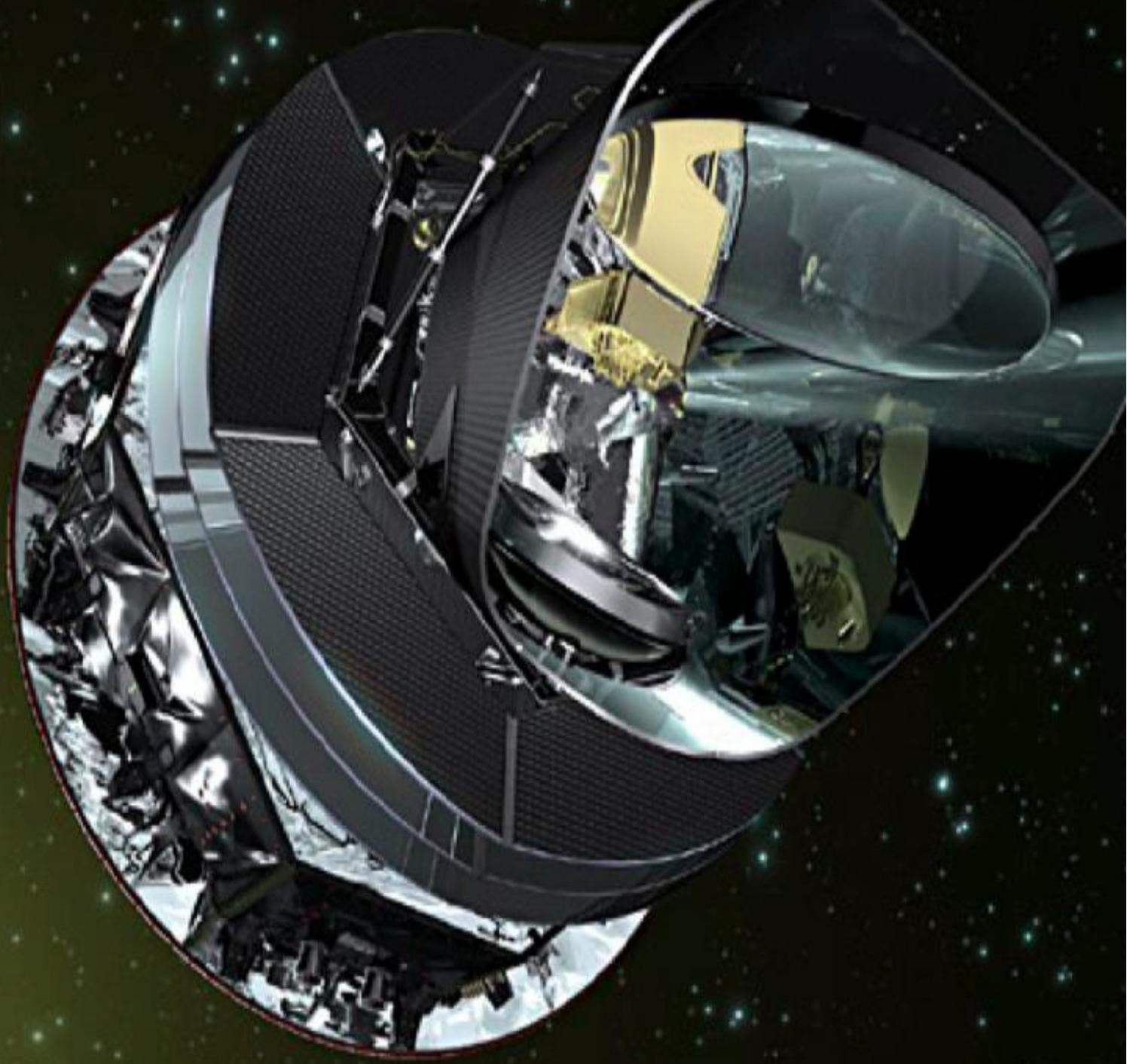
Everything formed from the Vacuum!

# Launch of the *Planck* satellite



Kourou, French Guyana: May 14, 2009

*Planck at L2*



# The nine *Planck* maps

30 GHz

44 GHz

70 GHz

100 GHz

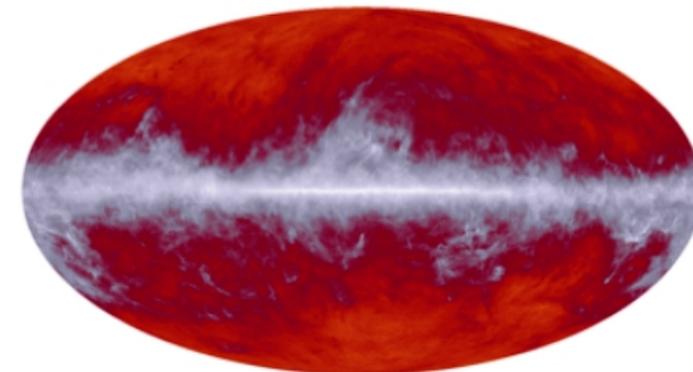
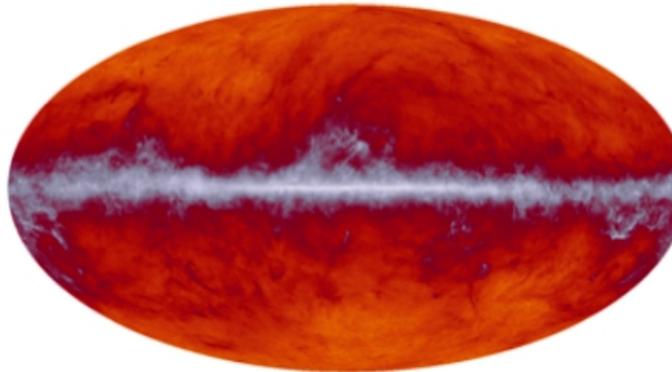
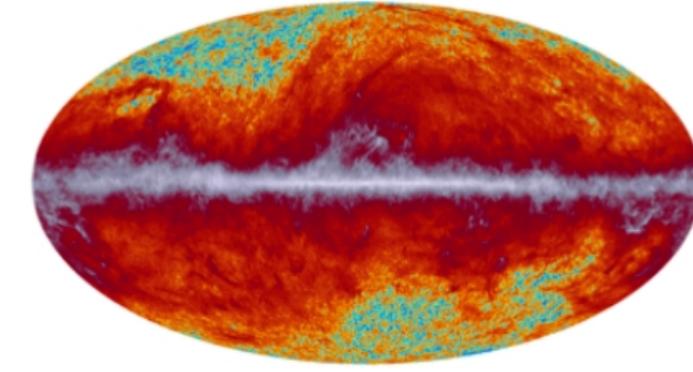
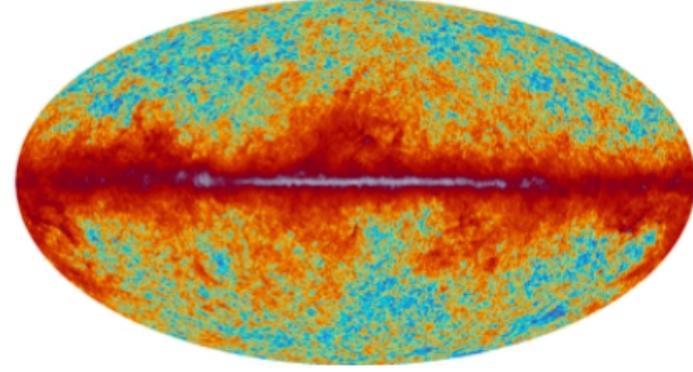
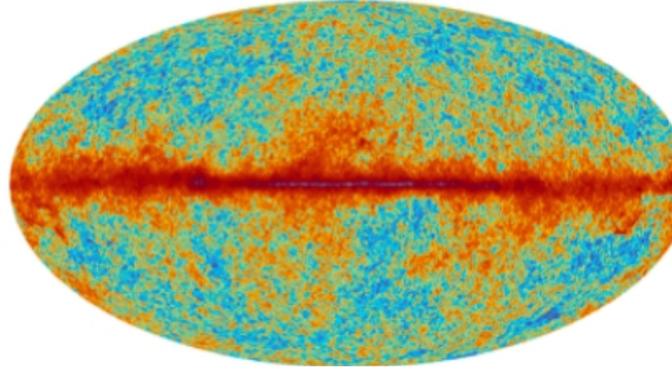
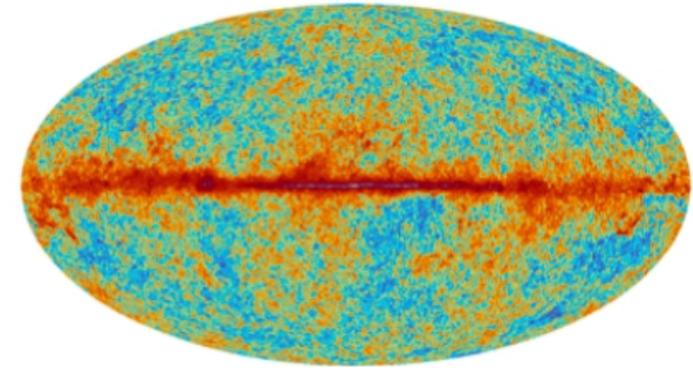
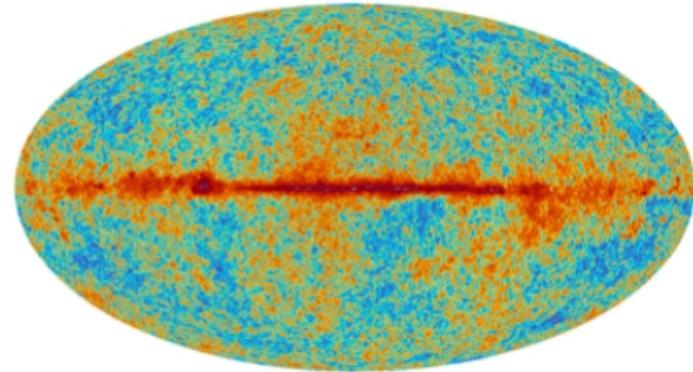
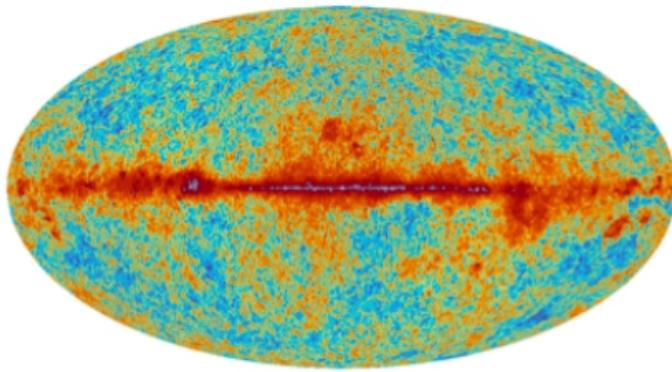
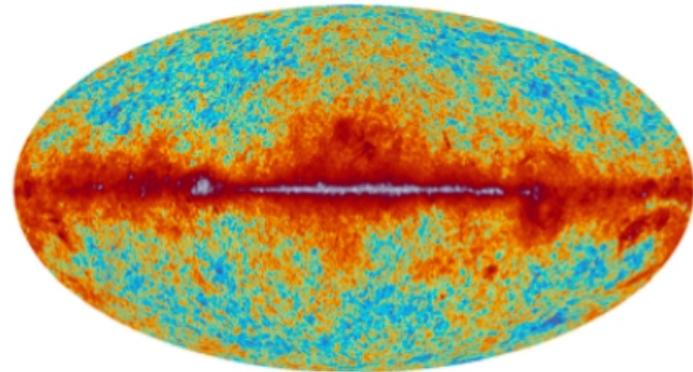
143 GHz

217 GHz

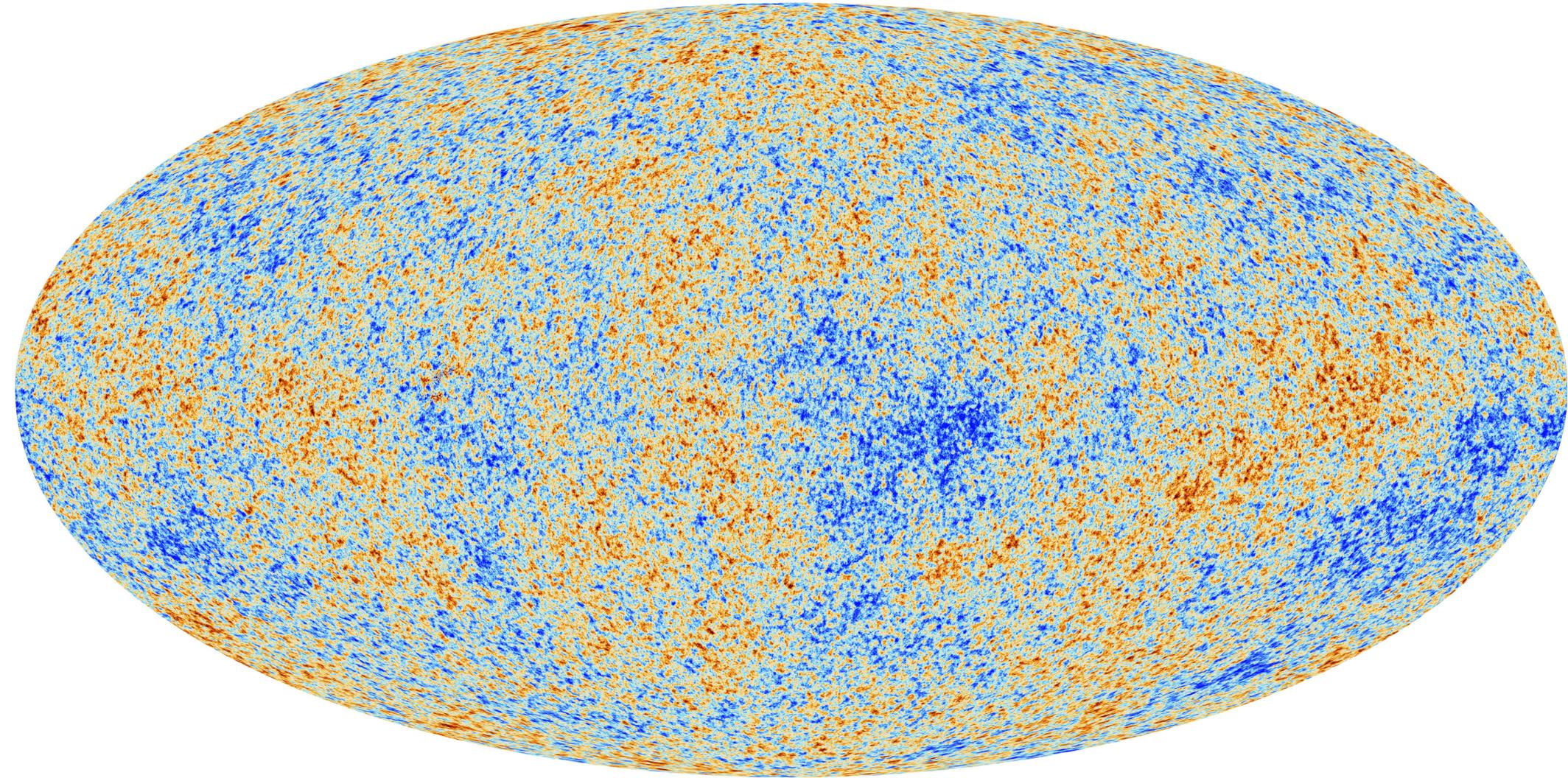
353 GHz

545 GHz

857 GHz

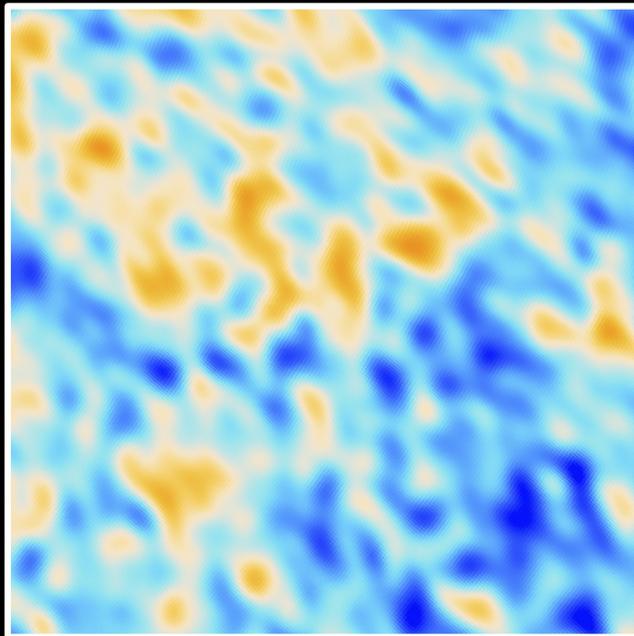
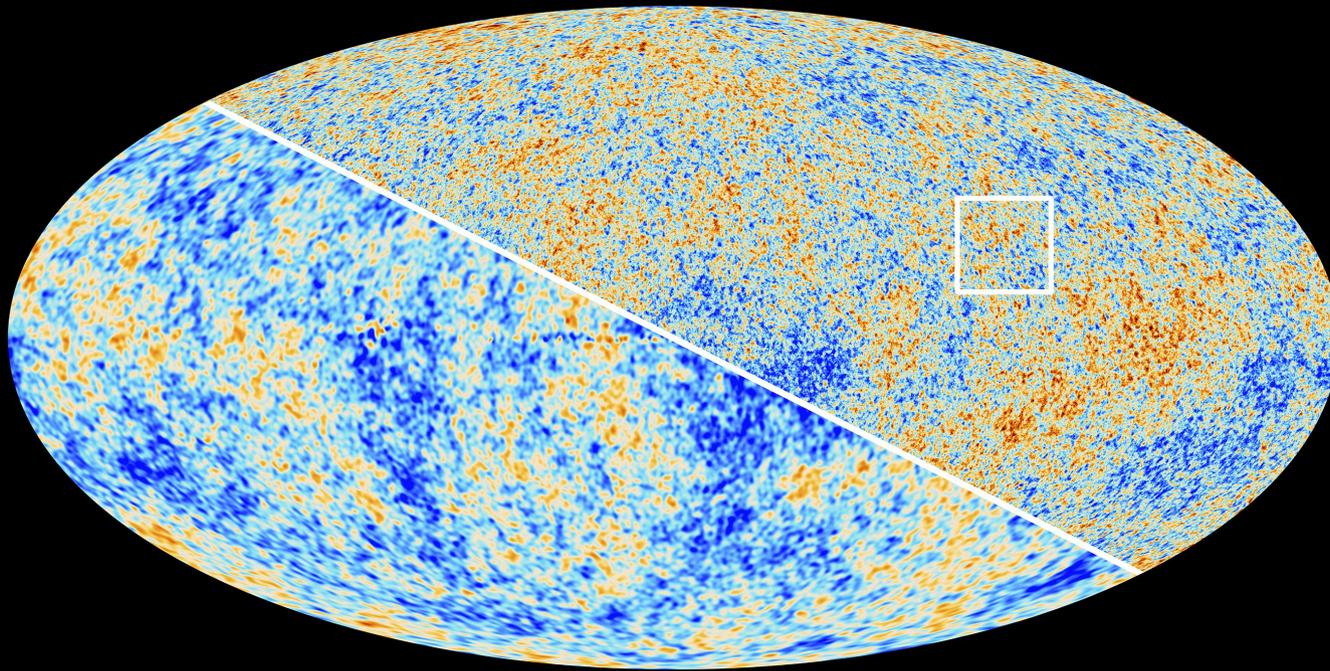


# The *Planck* map of the microwave background

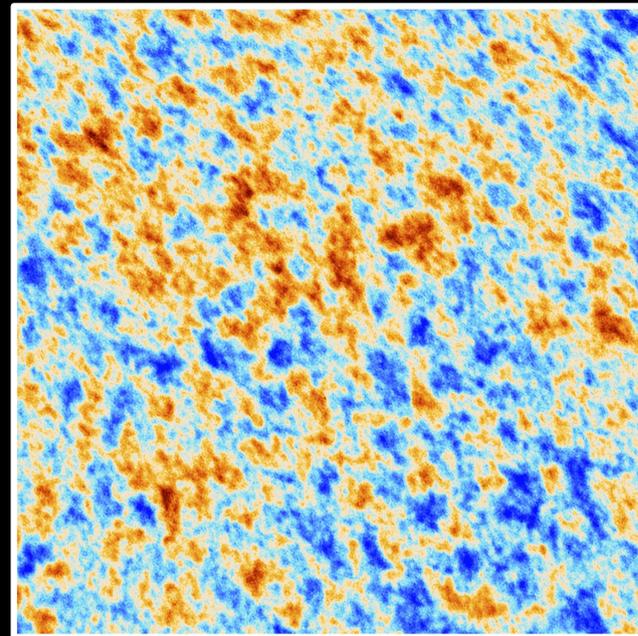


An image of the boundary of the observable Universe

# *Planck and WMAP in comparison*

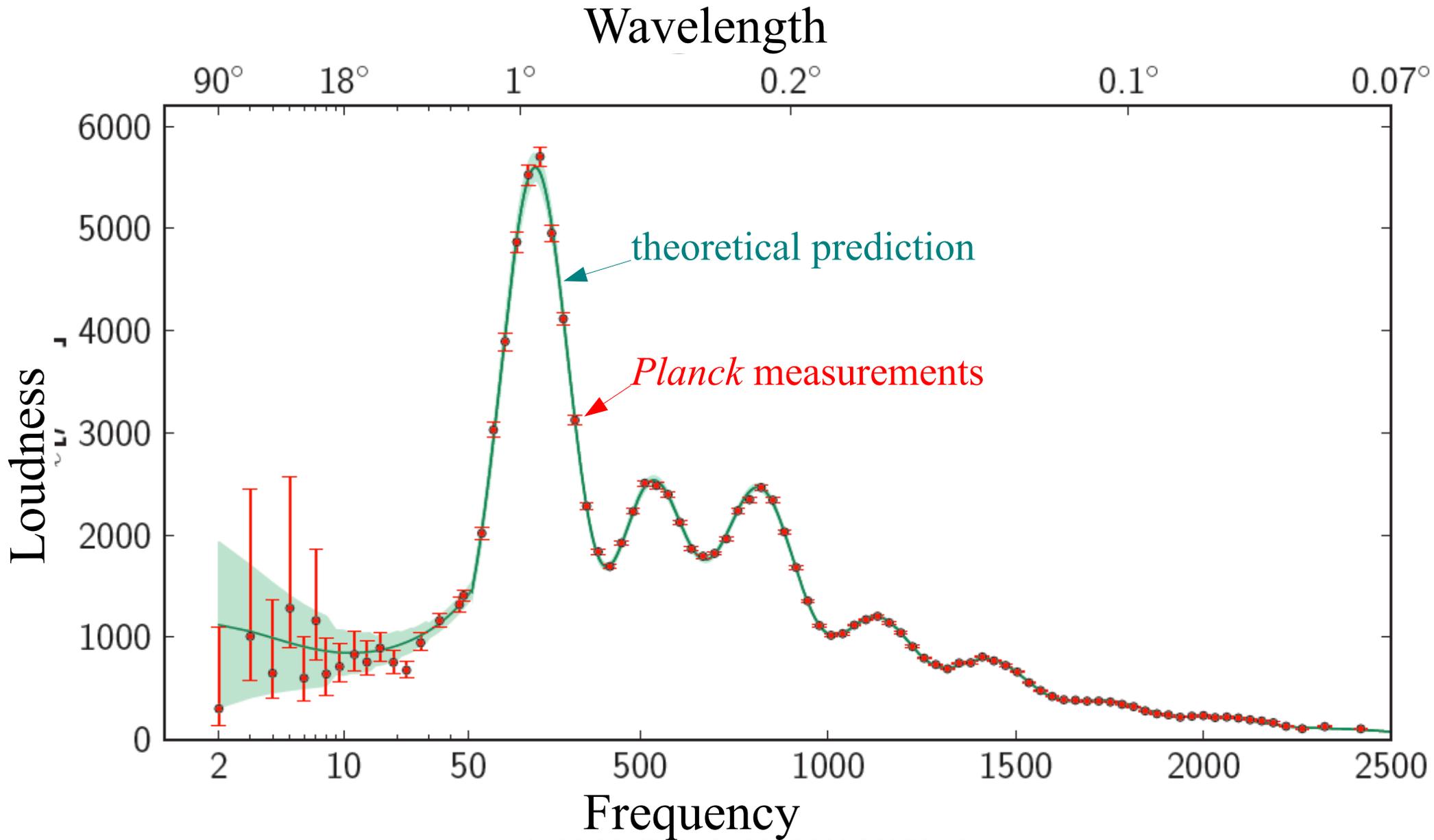


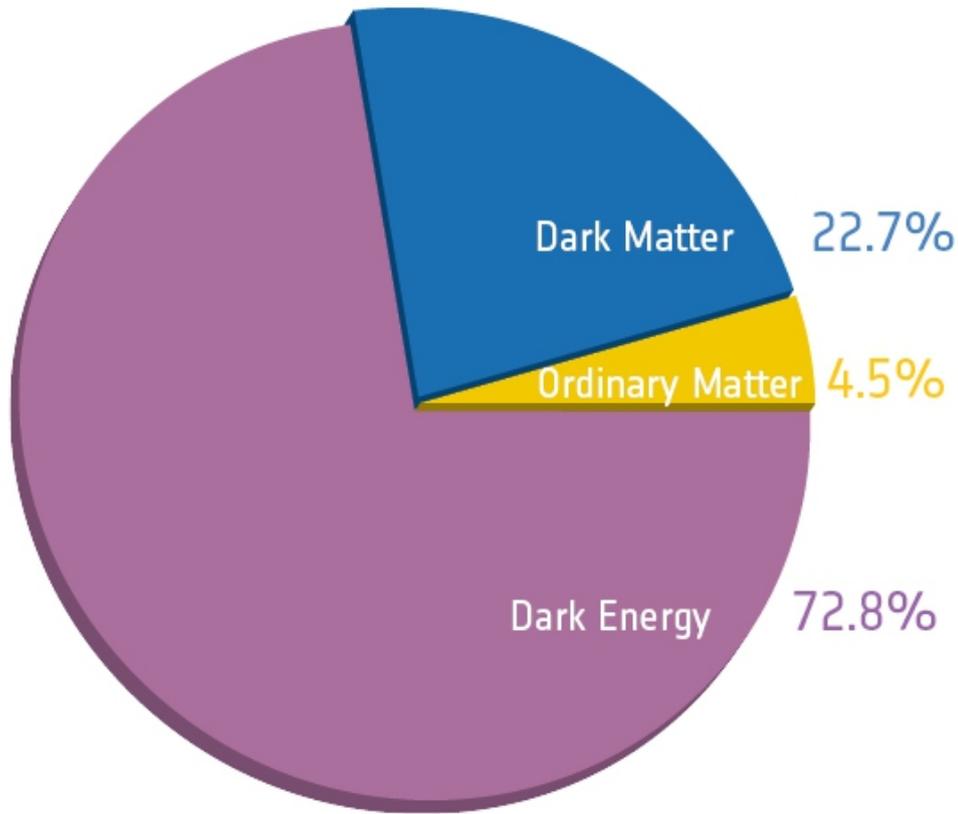
*WMAP*



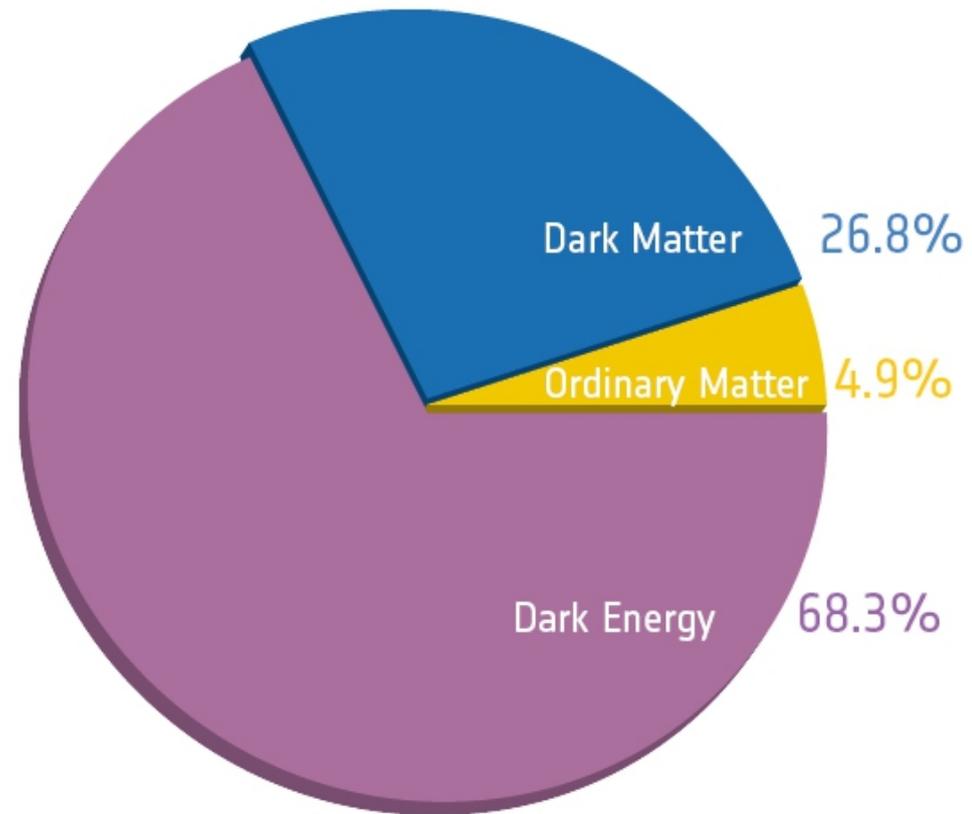
*Planck*

# Sound content of the cosmic clouds according to *Planck*





before *Planck*



after *Planck*

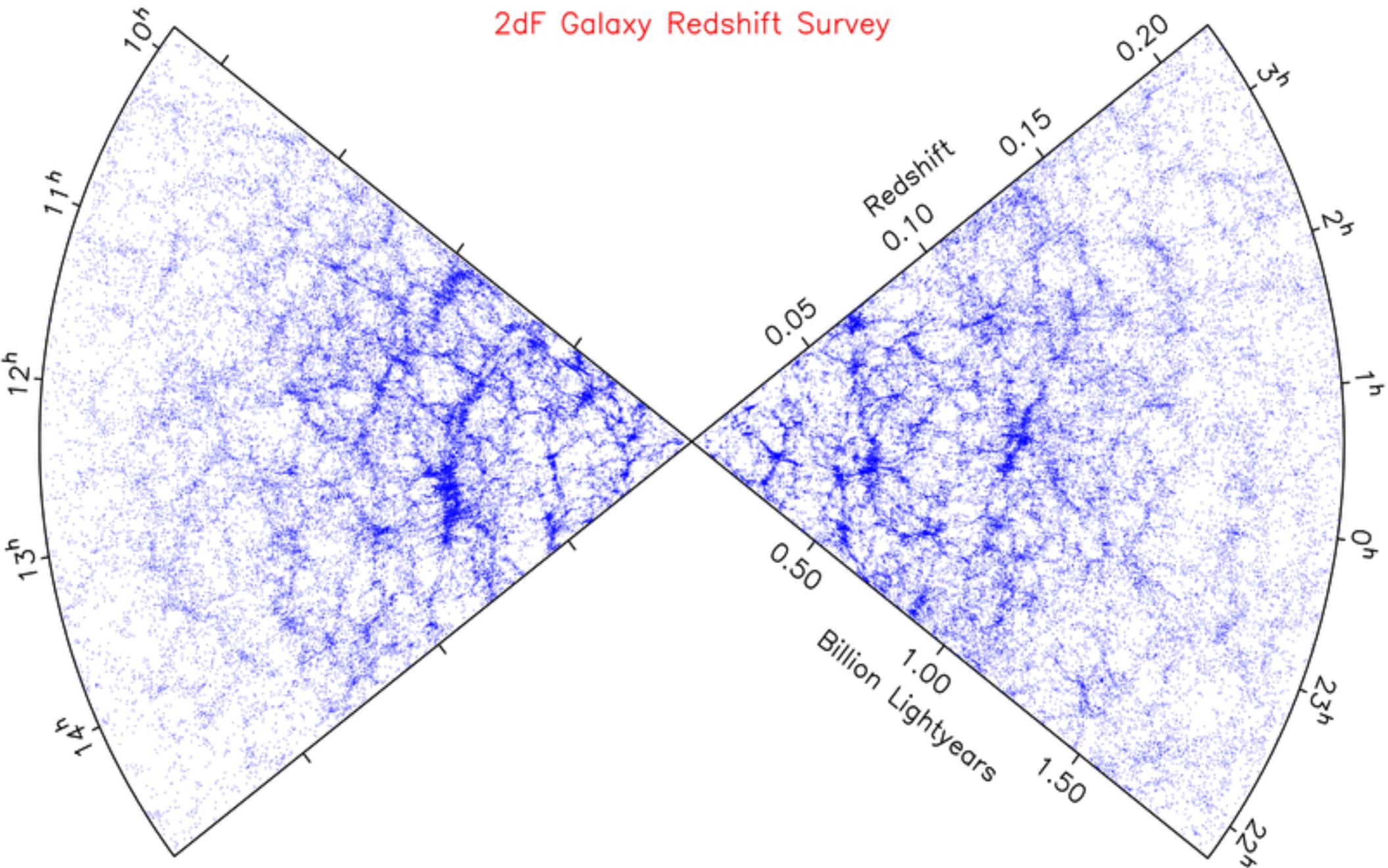
The Universe is still flat – now to better than 0.5% precision

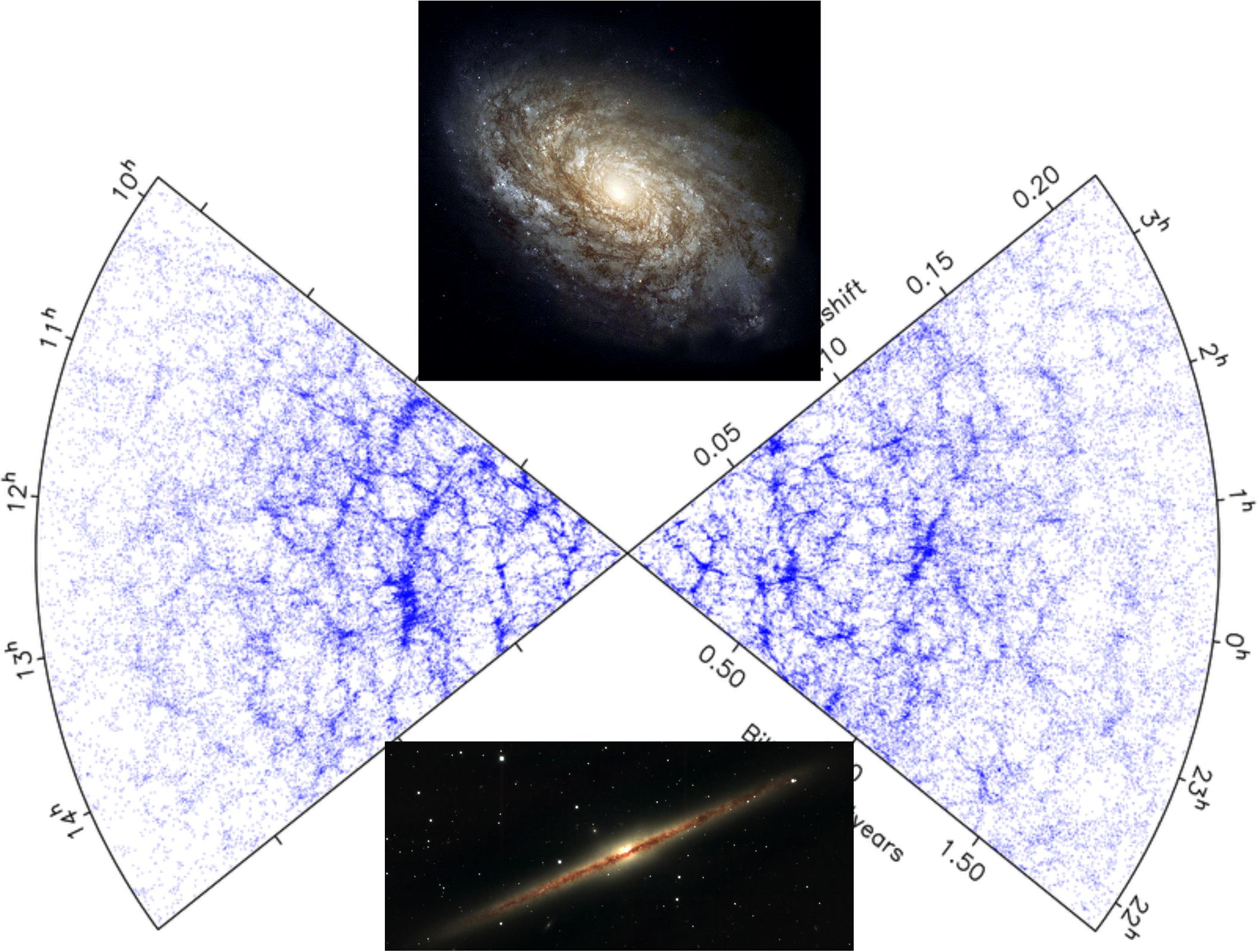
Its expansion rate is 7% slower than previously thought, so its age has increased by 80,000,000 years!

*Planck* results strongly support the idea that all structure originated from quantum zero-point fluctuations of the very early vacuum

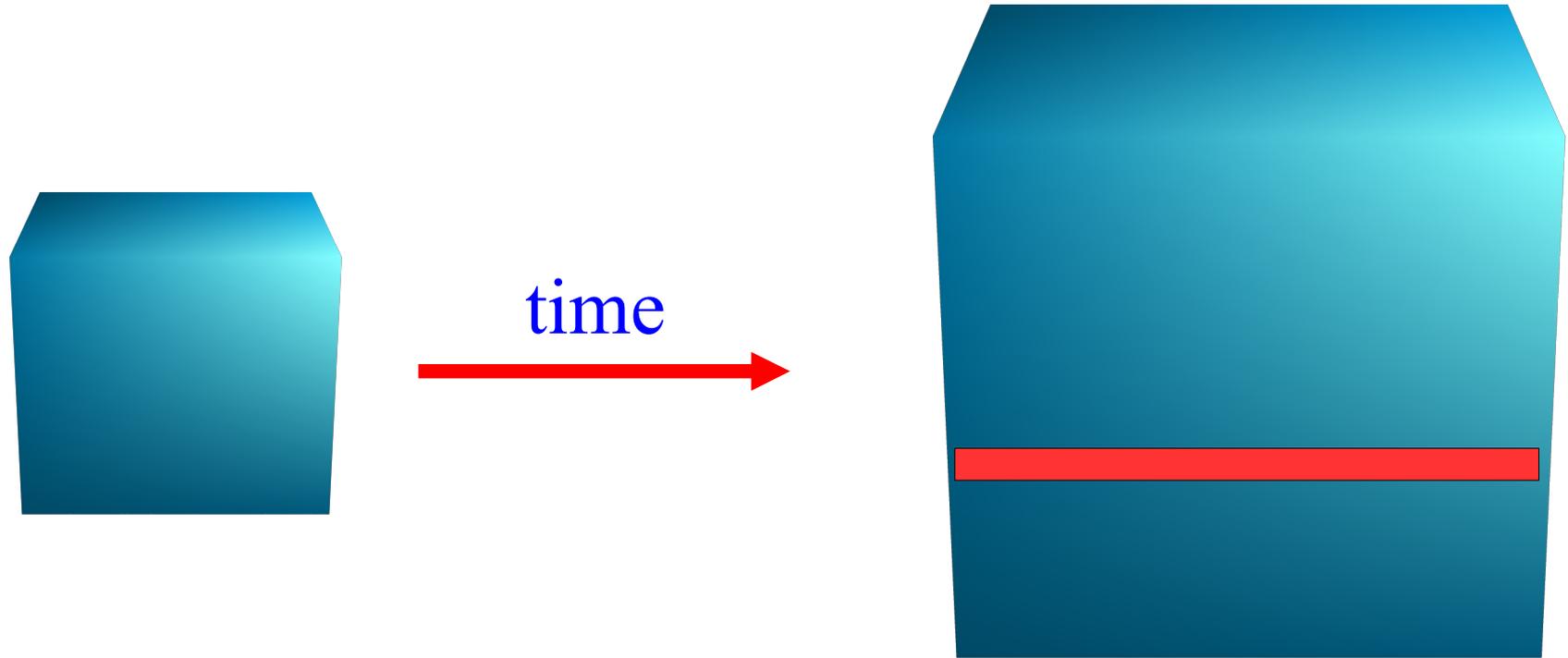
# Large-scale structure in the present-day Universe

2dF Galaxy Redshift Survey





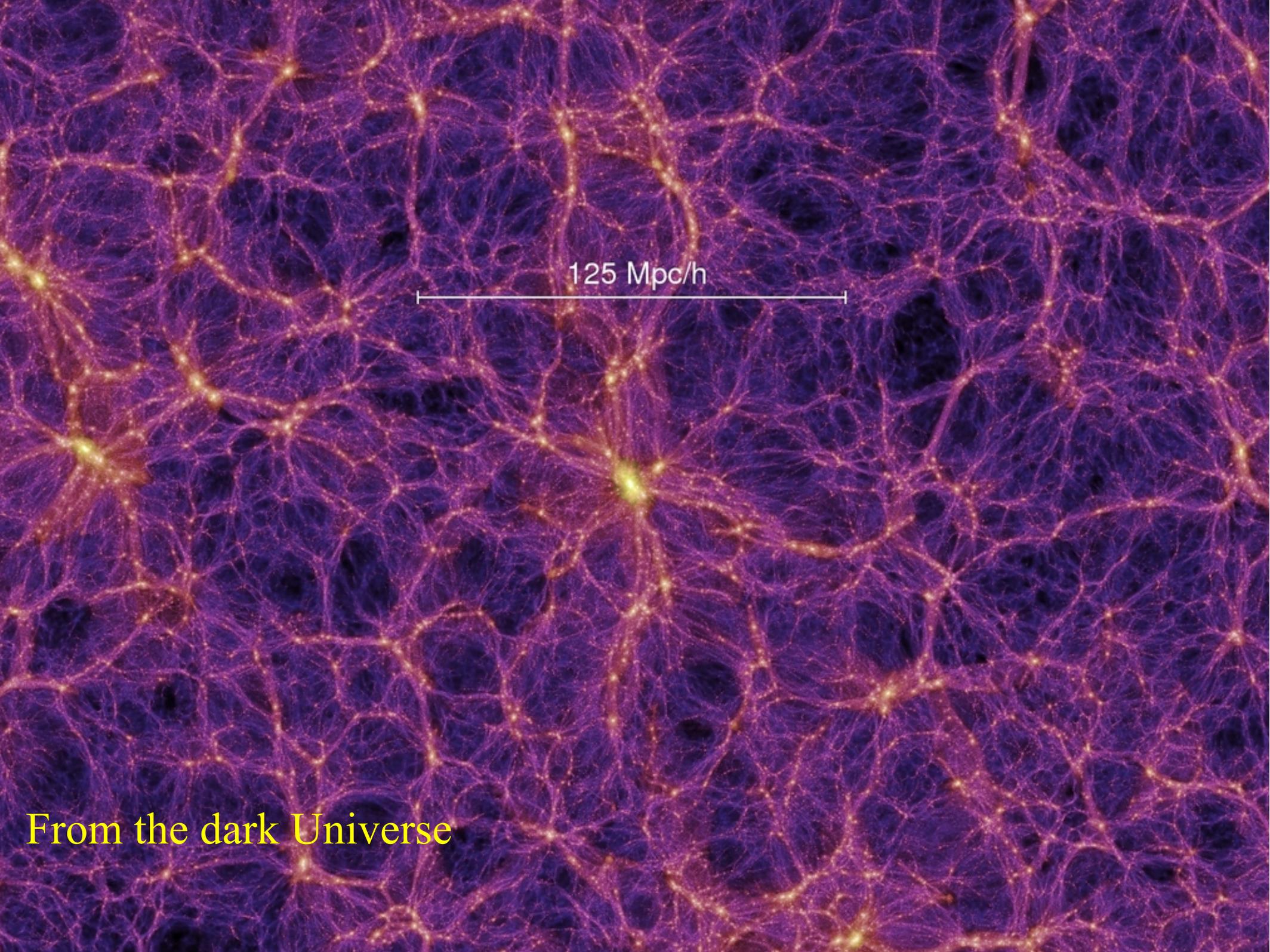
# How to follow cosmic evolution in a supercomputer



- Follow the material in a cube which expands with the Universe
- Start 400,000 years after the Big Bang
- Set initial conditions to match microwave background structure
- Calculate the evolution forwards to the present day

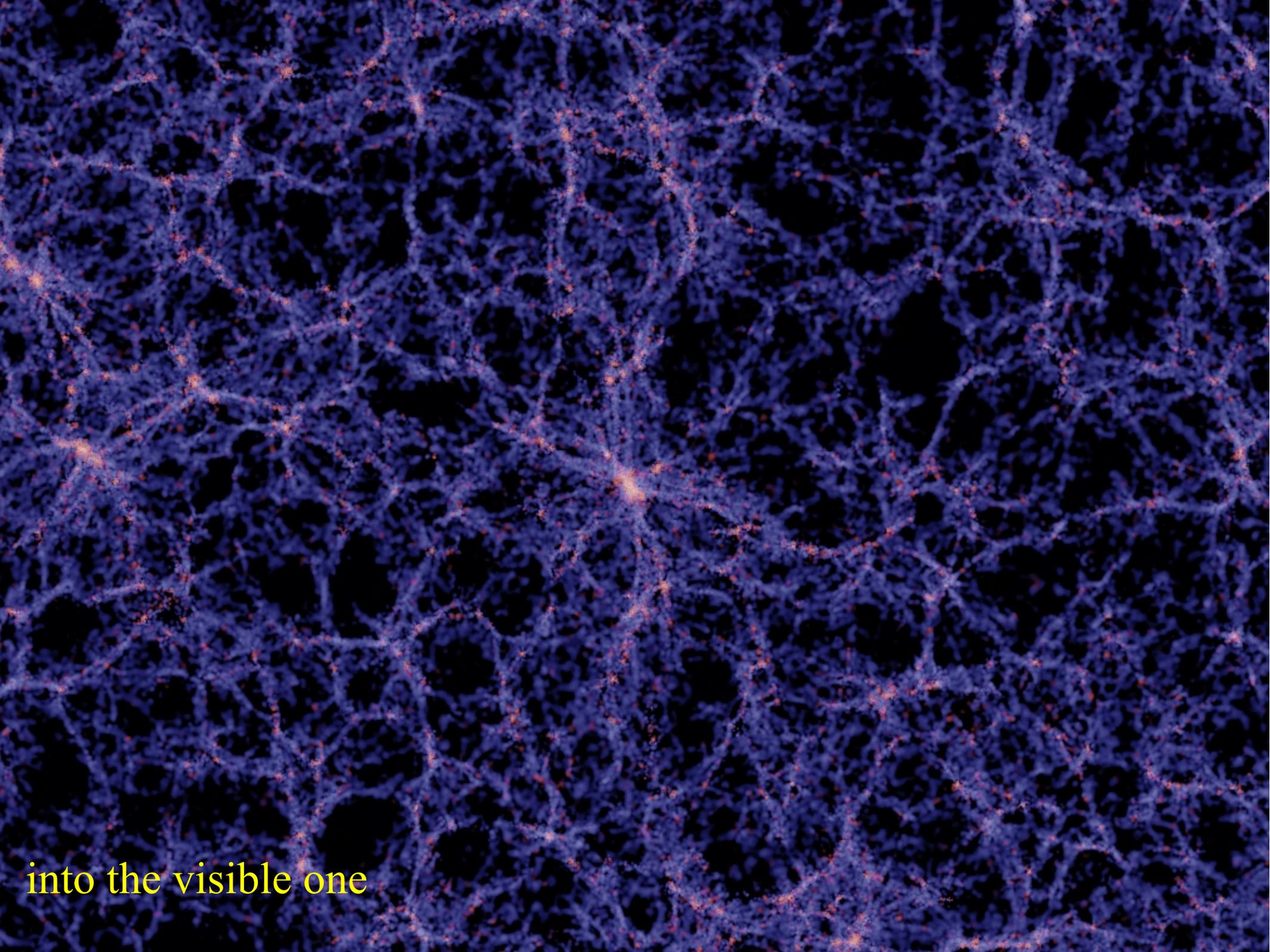
# Images of the Dark Matter distribution in a virtual Universe

- Evolution of structure in a thin, expanding slice
- A zoom from the entire visible Universe into a galaxy cluster
- A flight through the dark Universe

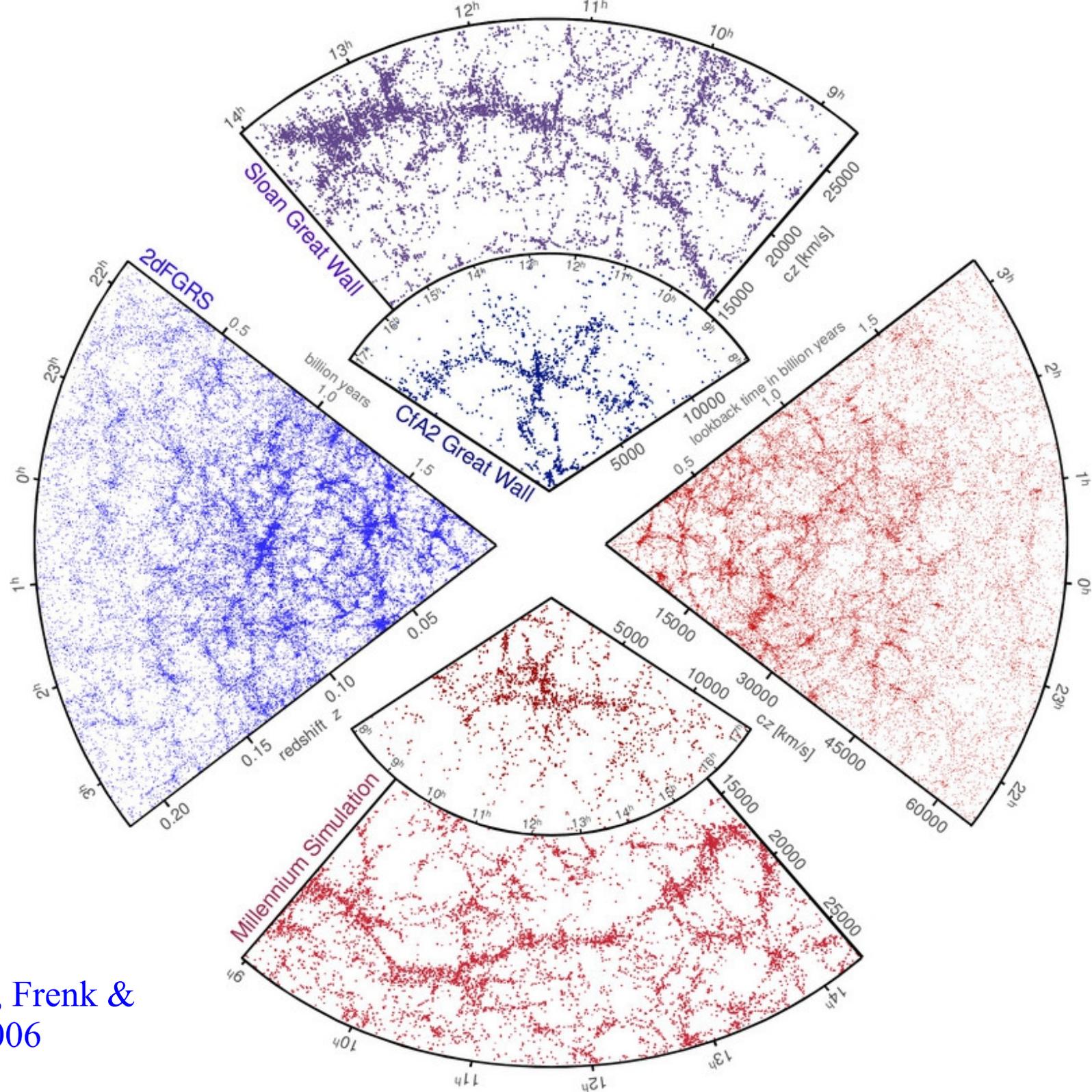


125 Mpc/h

From the dark Universe

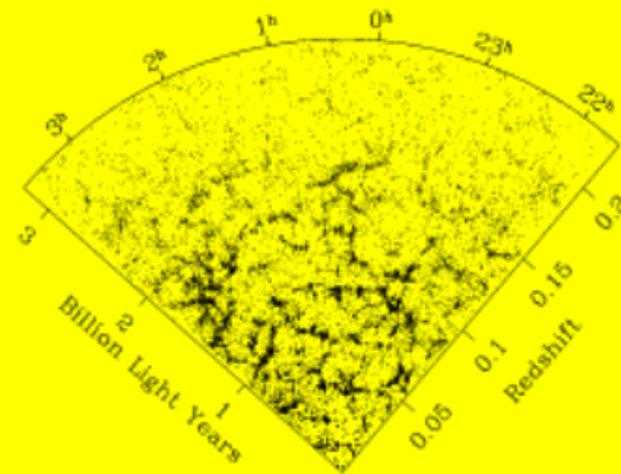
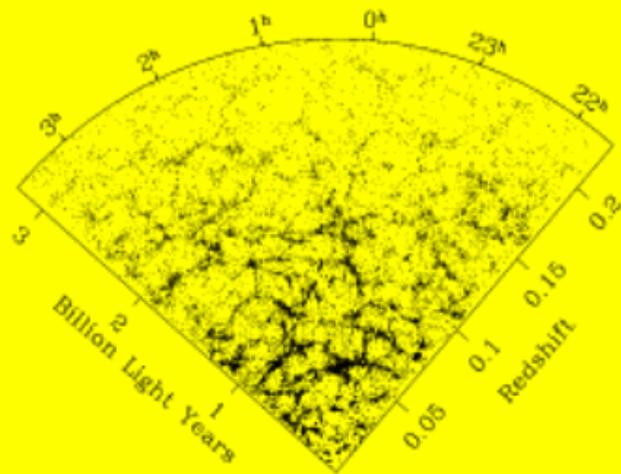
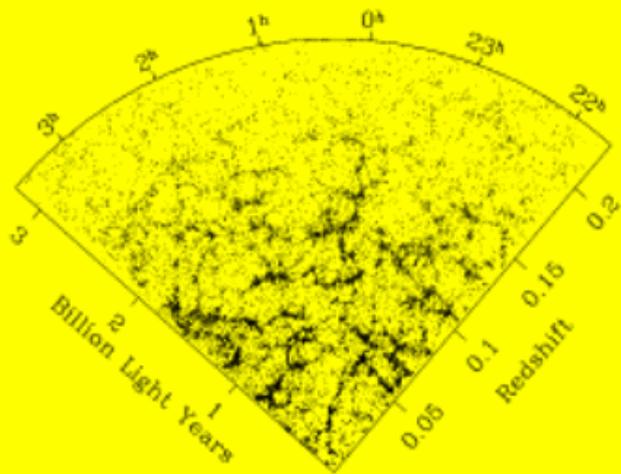
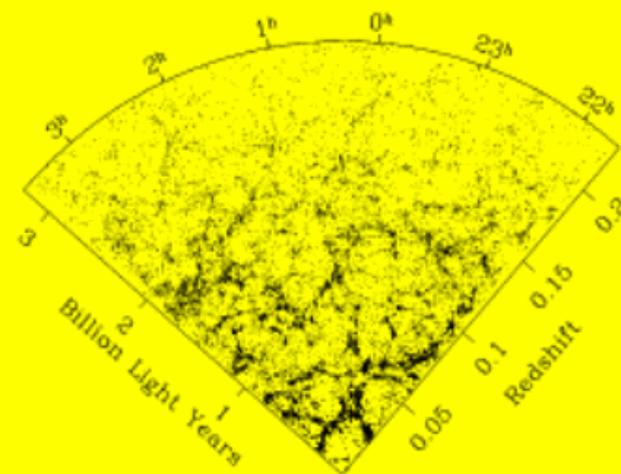
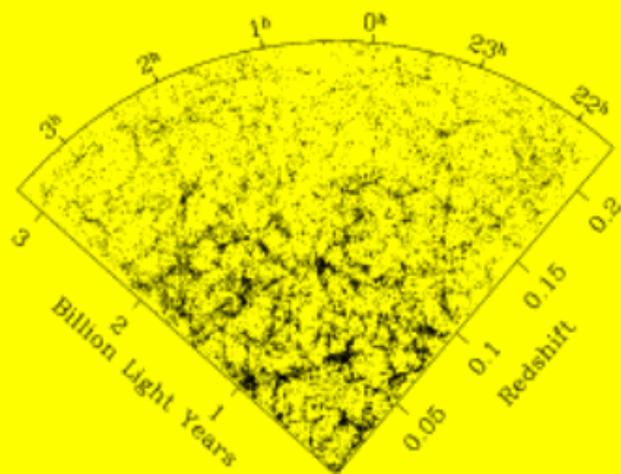
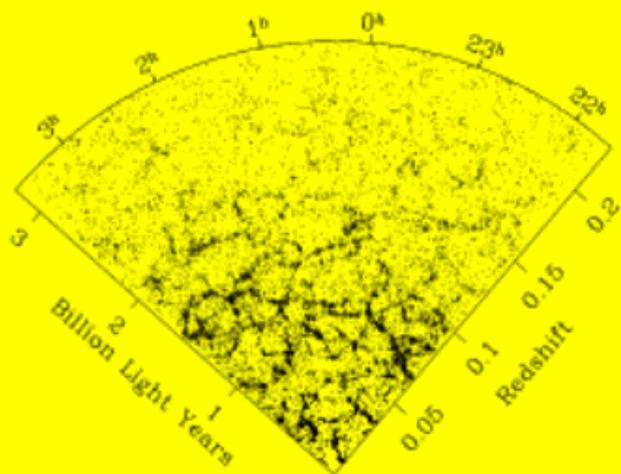


into the visible one



Springel, Frenk &  
White 2006

# Virtual versus real Universes



## To conclude....

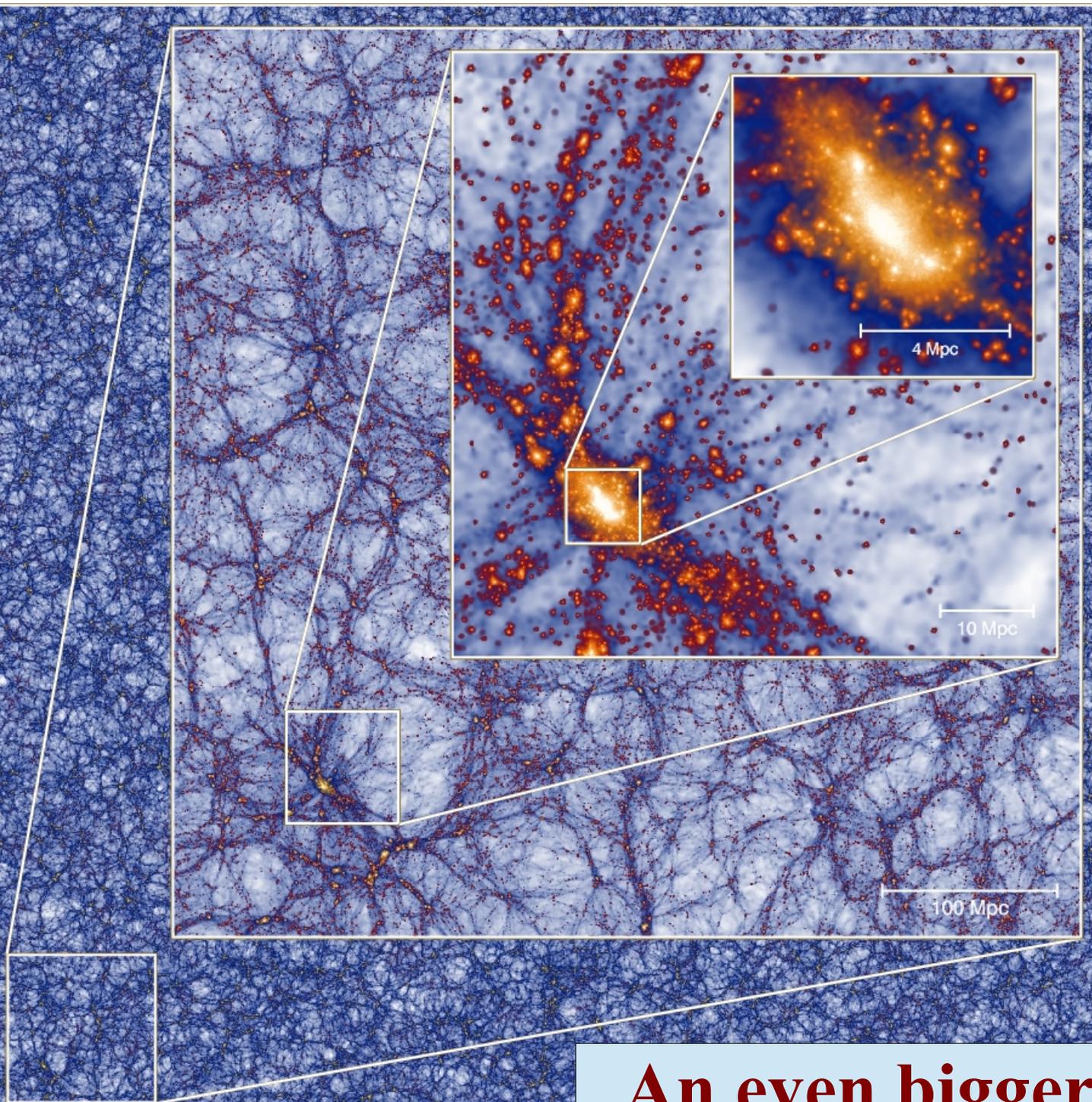
- Our Universe was born 13.7 billion years ago in a hot and almost uniform explosion -- the Big Bang
- All structure grew from quantum fluctuations of the early vacuum -- Everything has formed from “nothing”!
- Only 5% of today's Universe is made of ordinary matter
- About 27% is made of as yet unidentified elementary particles -- the Dark Matter
- About 68% consists of a new form of energy which accelerates the expansion of today's Universe -- the Dark Energy
- Galaxies and galaxy clusters, stars and planets formed out of sound waves in the primordial gas through the effects of Gravity



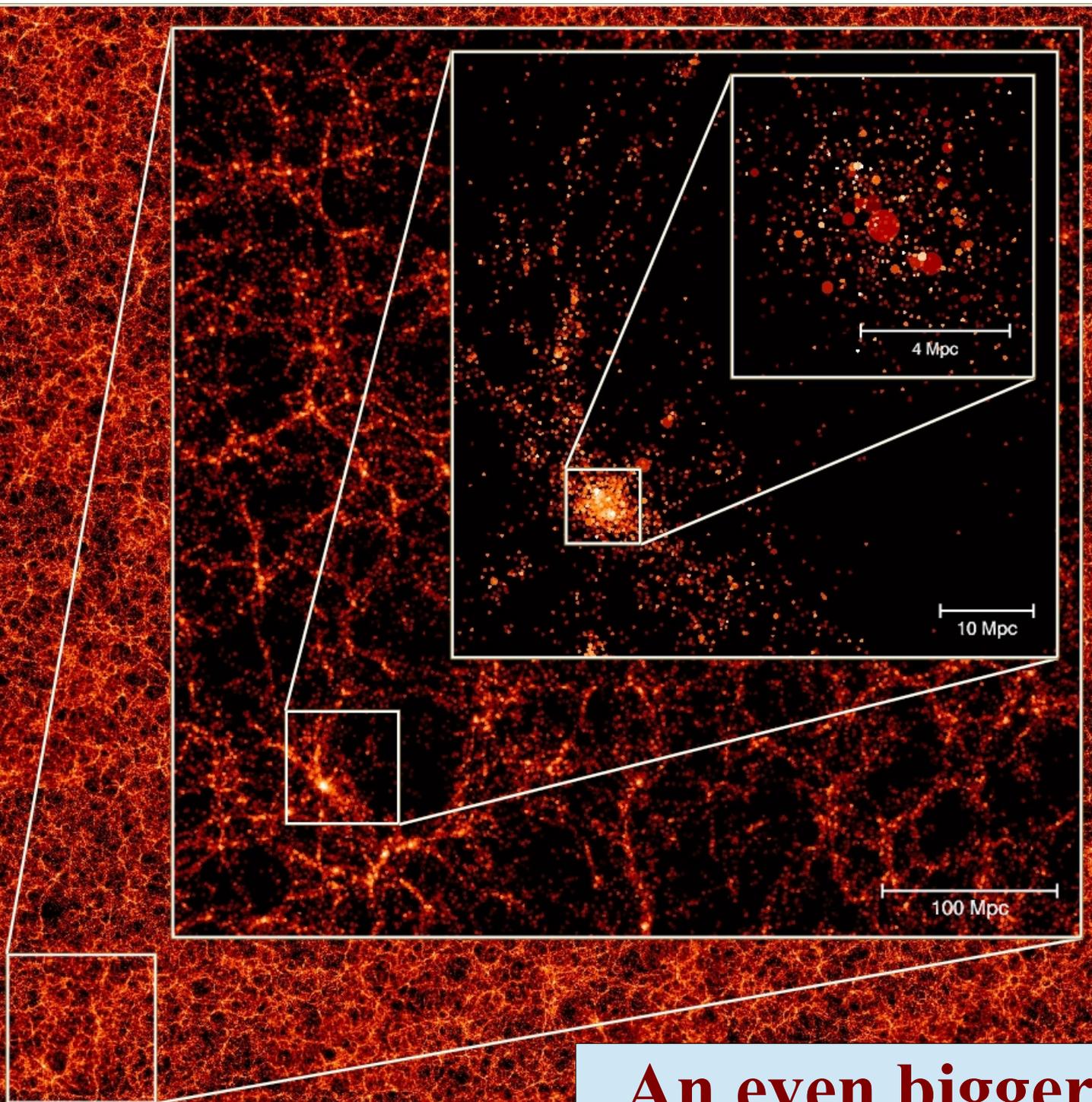
# Can we calculate into the future?

- What will happen to our Milky Way when it runs into the Andromeda nebula in about 3.5 billion years from now?
- What will become of the Earth?
- What will happen to humankind?





**An even bigger simulation:  
The Millennium XXL**



**An even bigger simulation:  
The Millennium XXL**