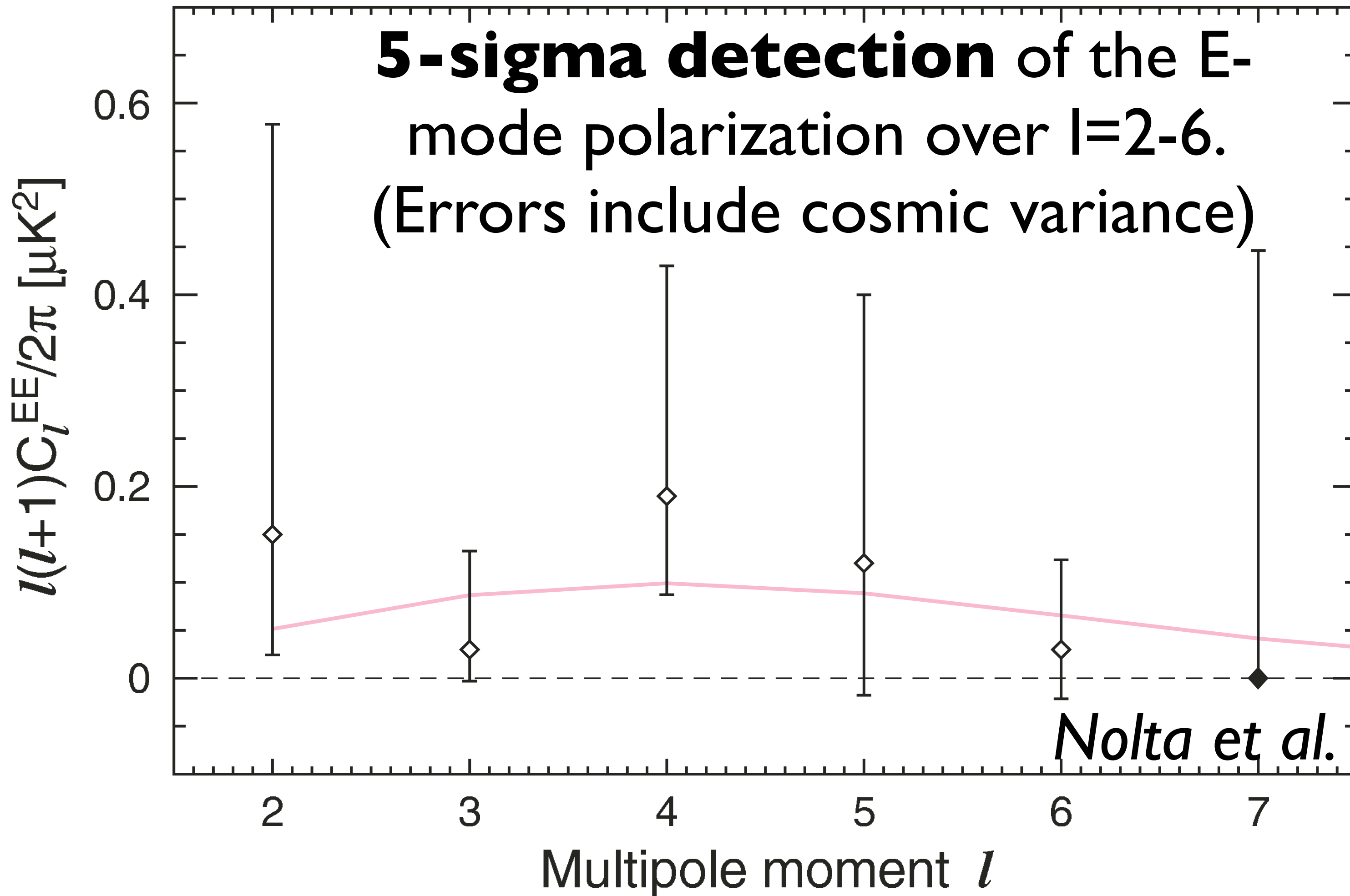


Reionization and Cosmological Parameters

Eiichiro Komatsu (Univ. of Texas at Austin)
CMBPol Workshop, June 24, 2008

5-Year E-Mode Polarization Power Spectrum at Low l

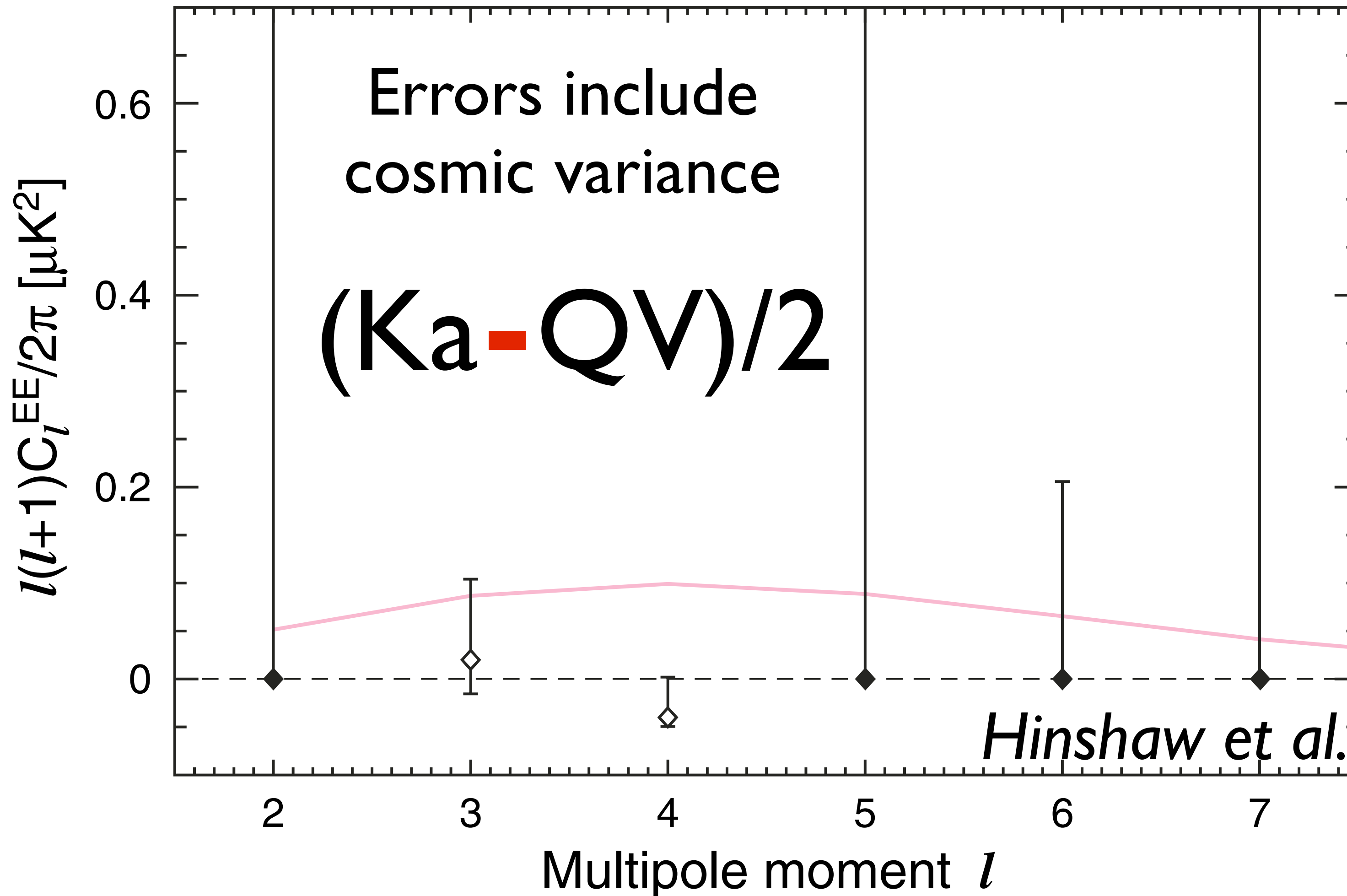
E-Mode Angular Power Spectrum



Black Symbols are upper limits

The signal is on the sky, baby...

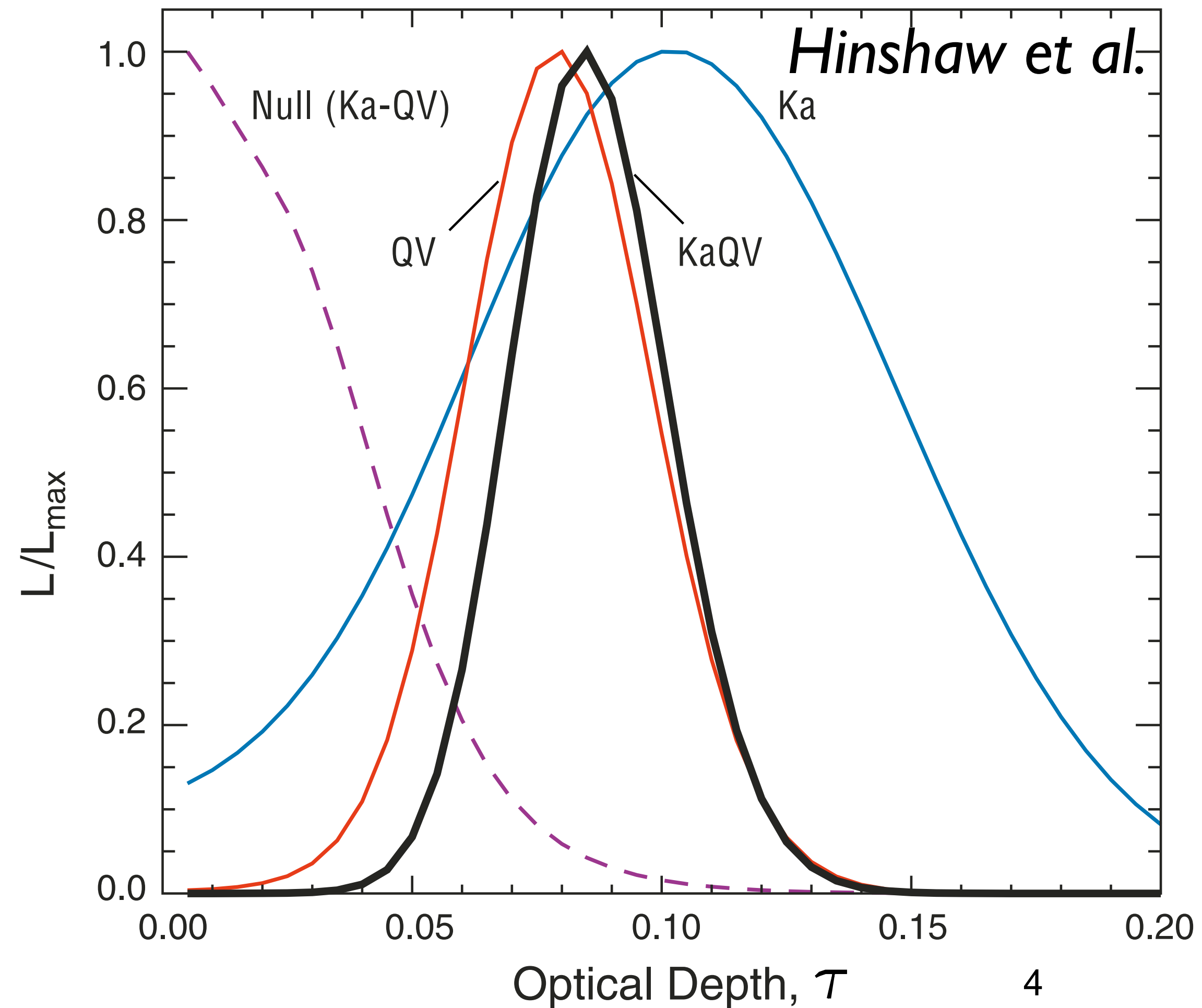
E-Mode Angular Power Spectrum



Black Symbols are upper limits

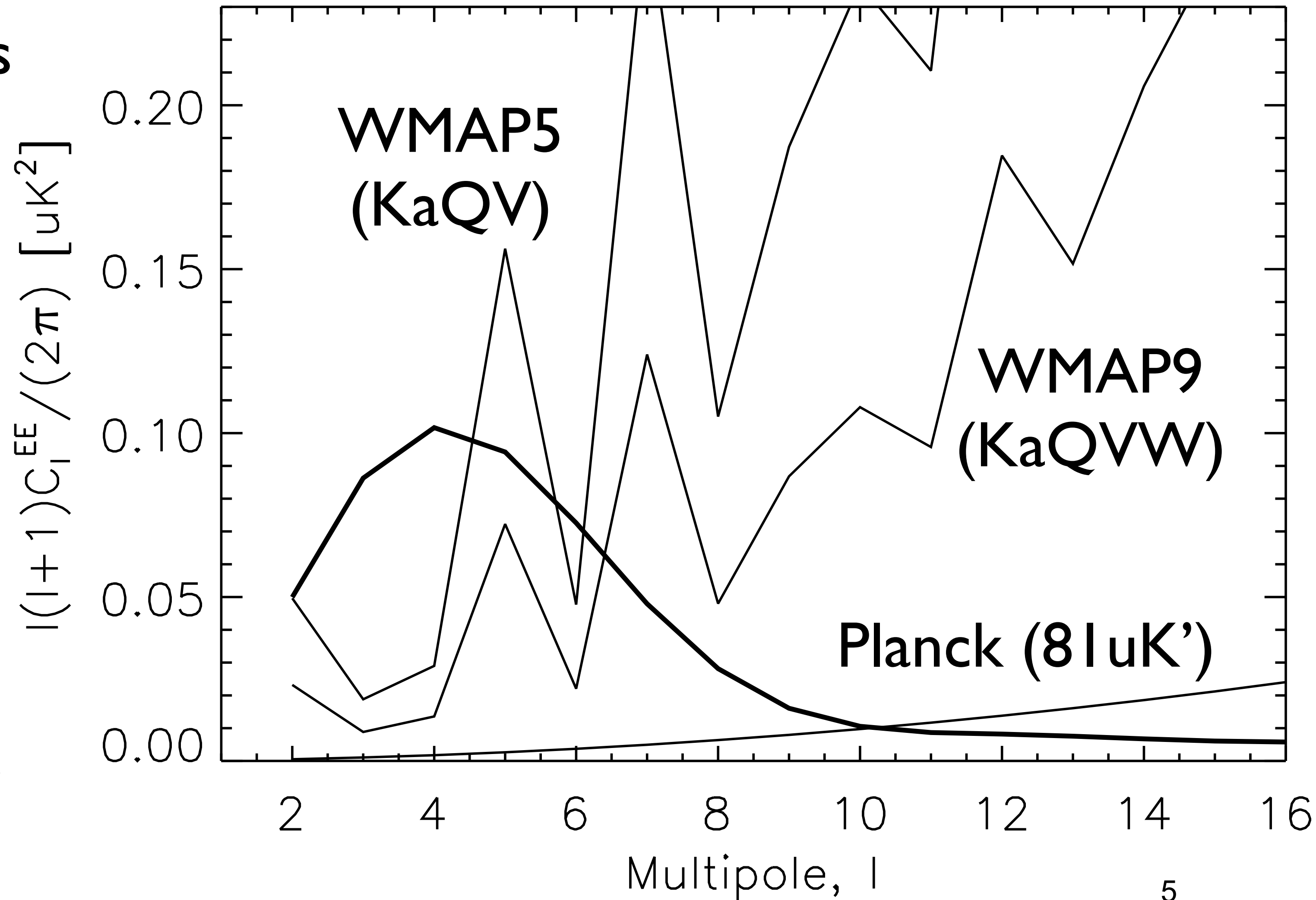
Measuring The Optical Depth of the Universe

- Optical Depth measured from the E-mode power spectrum:
- $\tau(5\text{yr}) = 0.087 \pm 0.017$
- $\tau(3\text{yr}) = 0.089 \pm 0.030$
(Page et al.; QV only)
- 3-sigma improved to 5-sigma!
- **What is WMAP actually measuring?**

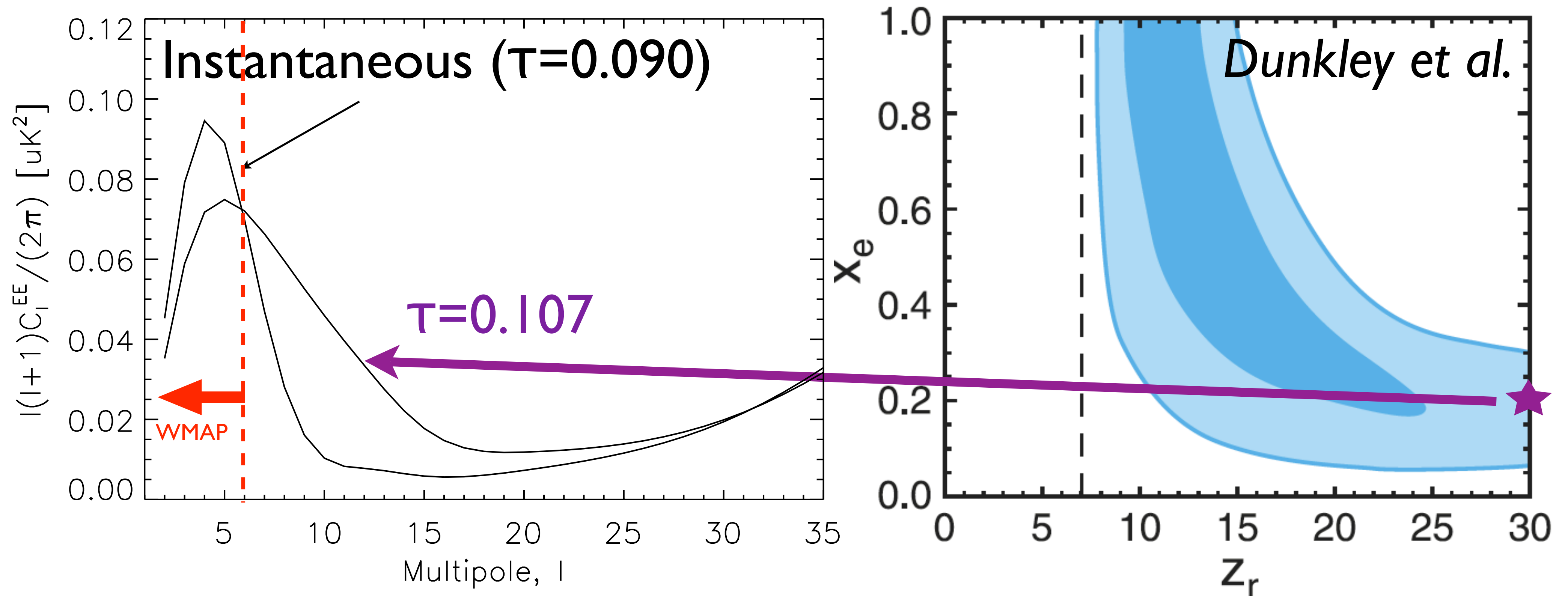


Signal vs Noise C_l (unbinned)

- Signal: instantaneous reionization with $\tau=0.090$
- WMAP9: CV dominated at $l < 6$
- Planck: CV dominated at $l < 10$, if noise is white
- $1/f$ noise increases the noise further

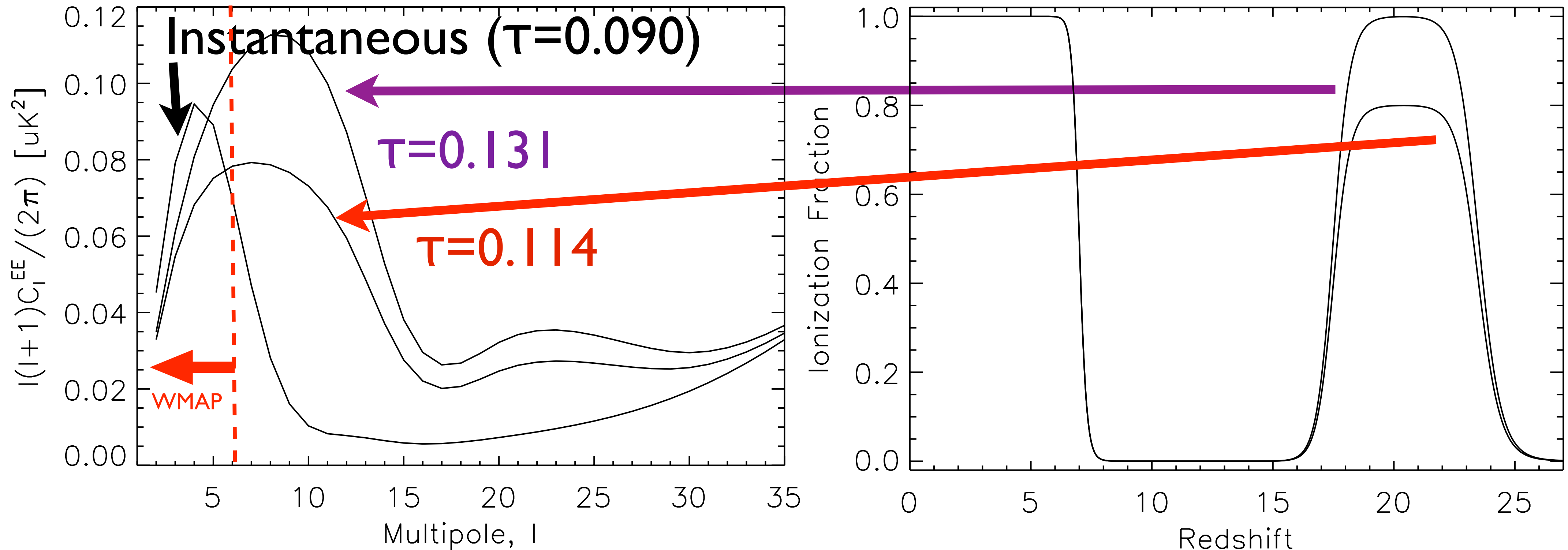


Two-step Reionization



- WMAP5 cannot constrain the EE power spectrum at $l > 7$.
- One has a freedom to choose reionization histories at $z > \sim 20$.
- You can hide τ at $z > \sim 20$!

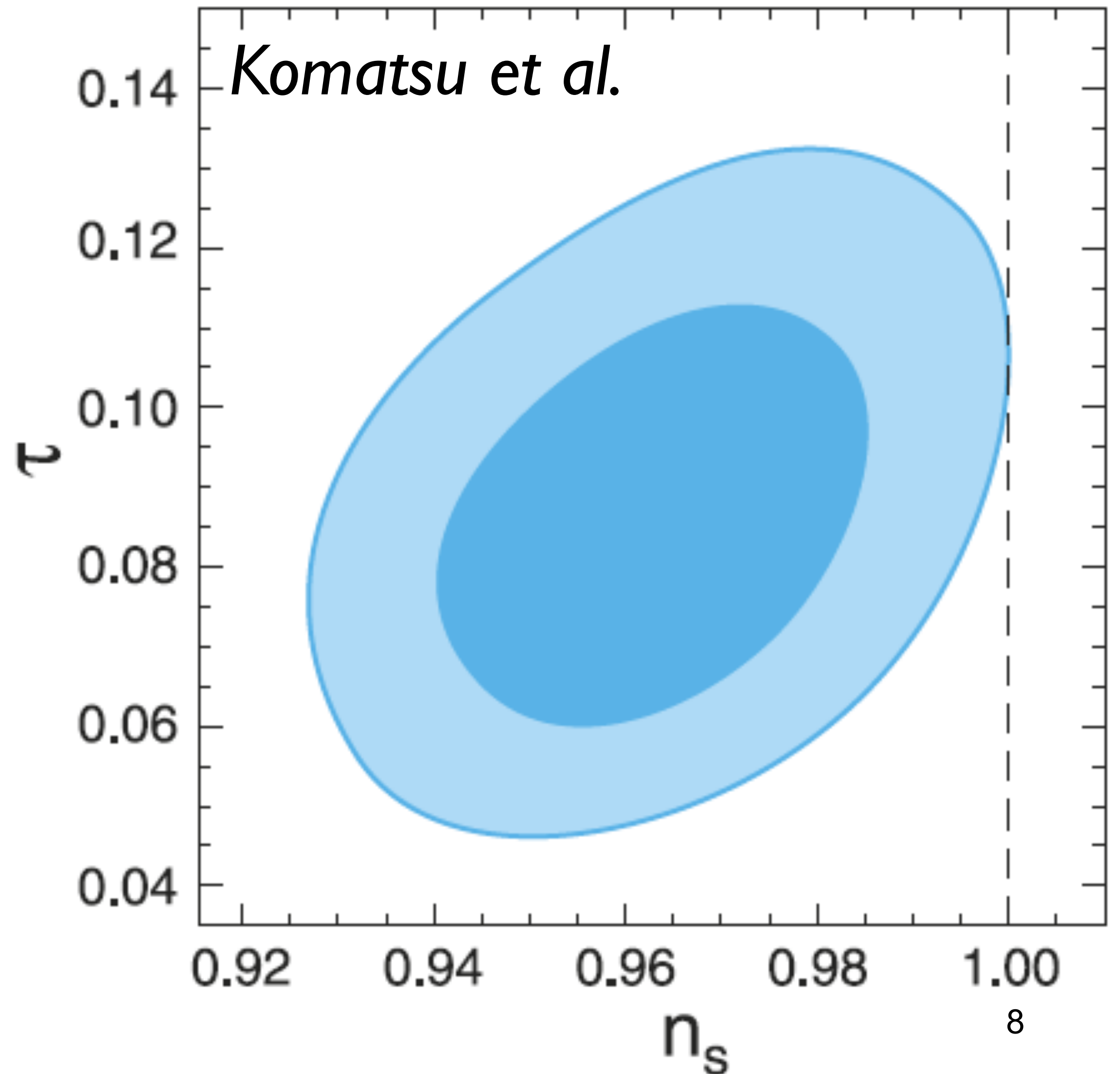
Double Reionization



- Even more τ can be hidden in the WMAP 5-year data for double reionization (i.e., non-monotonic reionization)

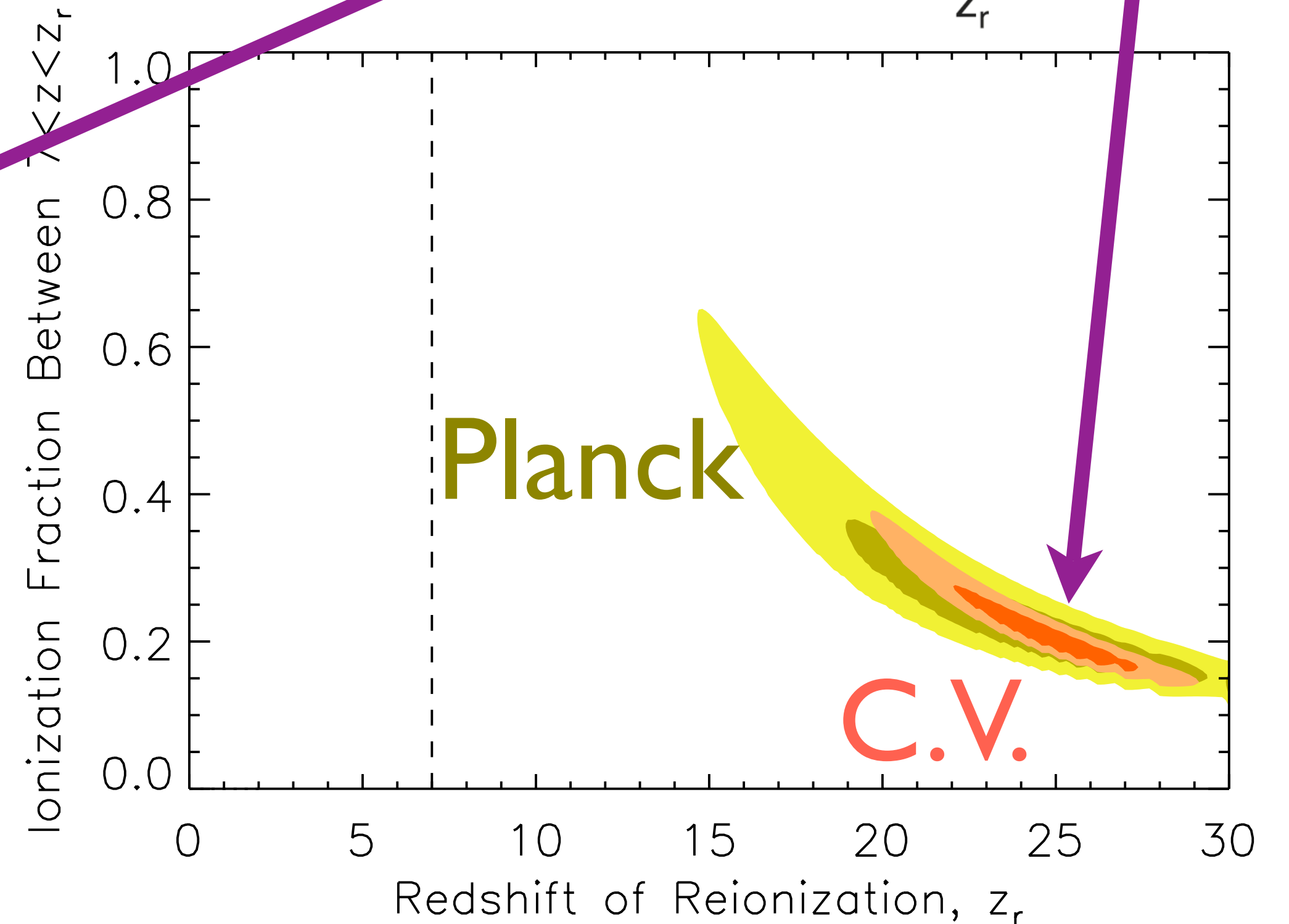
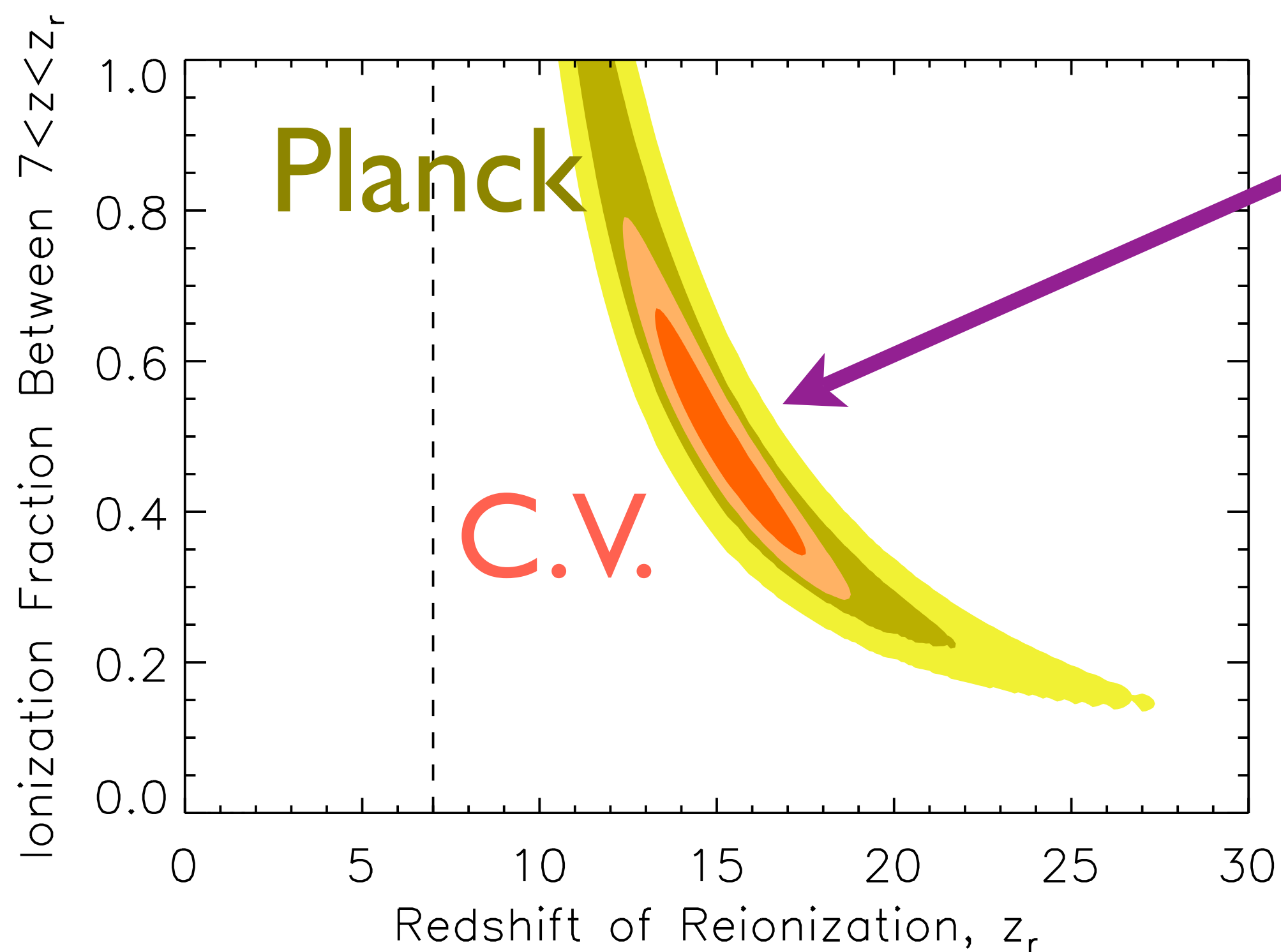
n_s - τ correlation

- n_s is closer to unity, if more τ is hidden in the WMAP data.
- Note that this plot assumes the instantaneous reionization.
 - The constraint would be relaxed for double reionization models (i.e., non-monotonic reionization)



Planck vs C.V. Limited

($f_{\text{sky}}=0.8$; $l_{\text{max}}=30$)



- The cosmic variance limited experiment constrains both x_e and z_{reion} MUCH better than Planck.

Conclusion

- Planck's EE measurement is cosmic variance limited (l-by-l) only up to $l \sim 10$, assuming white noise.
- Contribution of $1/f$ noise degrades the sensitivity further
- Note that WMAP9's EE would be cosmic variance limited (l-by-l) up to $l \sim 6$
- Will the cosmic variance limited experiment improve the limits on reionization histories? *Yes, significantly!*
- Implications for the cosmological parameters?
 - Accurate determination of n_s is possible only if we can find ALL of τ out there.

What I learned yesterday from Elena Pierpaoli

- n_s and τ are totally de-correlated for Planck and CMBPol
- These experiments won't need τ for measuring n_s .
- So, implications of Planck's and CMBPol's reionization measurement for the other cosmological parameters may be a lot less than those for WMAP.
- In fact, there may be no implication at all.
- Good news!