# KiDS survey: first lensing results

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Konrad Kuijken Leiden Observatory

300-Mpix CCD mosaic camera OmegaCAM@VST

#### KiD5 "The other surveys"

- Dark Energy Survey 5000 sqdeg, i<sub>AB</sub>=24
- Kilo-Degree Survey 1500 sqdeg, r<sub>AB</sub>=24.9
- HyperSuprimeCam 1375 sqdeg, i<sub>AB</sub>=26
- soon: PAU, J-PASS 8500 sqdeg narrow-band mAB~22.5

- Each targets large-scale structure/cosmology
- KiDS focus is on high-fidelity lensing / tomography



# **KiDS Lensing team**



### KiDS KiDS in a nutshell

- 9-band survey u-K with VST/ OmegaCAM and VISTA
- 1500 square degrees
- 2mag deeper than SDSS, 2x sharper
- weak lensing + photoz
  optimized: DARK M&E
- started Oct 15 2011
- 500 sqdeg observed



#### 2.6m VST ugri: KiDS

4.2m VISTA ZYJHK: VIKING – (PI. Alastair Edge)

- KiDS High-fidelity
  - homogeneous PSF width, low anisotropy, constant plate scale, median seeing for lensing data 0.7"



9-band photometry for photo-z ugriZYJHK

#### KiDS Data quality (first 148 sq.deg.)



Jelte de Jong



~500 out of 1500 square degrees observed (ugri). Seeing (r<g<u) + moon (i) dictate schedule. Observations go to full depth (5 dithers) at once. No variability



#### **GAMA: Galaxy And Mass Assembly**

Driver++ 2012, Liske ++2015

- 250,000 redshifts from AAT
- flux-limited (r<19.8)
- complete in dense regions
- deep, so efficient lenses
- Group catalogue
  Robotham 2011



KiDS Data processing

- Two parallel pipelines. Significant CFHTLenS heritage
  - Shapes: THELI + *lens*fit (r band)
    - Thomas Erben, Lance Miller, Catherine Heymans
    - single-exposure model fitting, no pixel resampling
    - Bayesian treatment of size and ellipticity priors
    - calibrated with simulations
    - catalogues blinded analyse 3 sabotaged data sets + the real one.
    - Photometric redshifts: GAaP + BPZ (u,g,r,i band)

KK, Hendrik Hildebrandt

- stacked images
- PSF Gaussianization + matched-aperture photometry
- BPZ (Benitez 2000)

#### KiDS Photometric redshifts (ugri)



#### KiDS Photometric redshifts (ugri)





## KiDS First KiDS papers!

#### ArXiv 2015.00740±14

#### Dark matter halo properties of GAMA galaxy groups from 100 square degrees of KiDS weak lensing data

M. Viola<sup>1\*</sup>, M. Cacciato<sup>1</sup>, M. Brouwer<sup>1</sup>, K. Kuijken<sup>1</sup>, H. Hoekstra<sup>1</sup>, P. Norberg<sup>2</sup>,

Gravitational Lensing Analysis of the Kilo Degree Survey

Konrad Kuijken<sup>1\*</sup>, Catherine Heymans<sup>2</sup>, Hendrik Hildebrandt<sup>3</sup>, Reiko Nakajima<sup>3</sup>, Thomas Erben<sup>3</sup>, Jelte T.A. de Jong<sup>1</sup>, Massimo Viola<sup>1</sup>, Ami Choi<sup>2</sup> Henk Hoekstra<sup>1</sup>

Lance Miller<sup>4</sup>, Edo van Uitert<sup>3,5</sup>, Alexandra Amon<sup>2</sup>, Chris B Axel Buddendiek<sup>3</sup>, Ian Fenech Conti<sup>7,8</sup>, Martin Eriksen<sup>1</sup>, An Kilo Degree Survey

C. Tortora<sup>1\*</sup>, F. La Barbera<sup>1</sup>, N.R. Napolitano<sup>1</sup>, N. Roy<sup>1,2</sup>, M. Radovich<sup>3</sup>,

First discoveries of  $z \sim 6$  quasars with the Kilo Degree Survey and VISTA Kilo-Degree Infrared Galaxy survey<sup>\*</sup>

B. P. Venemans,<sup>1</sup><sup>†</sup> G. A. Verdoes Kleijn,<sup>2</sup> J. Mwebaze,<sup>2</sup> E. A. Va <u>E Bañados <sup>1</sup> R Decarli <sup>1</sup> I. T. A. de Iong <sup>3</sup> I. R. Findlav <sup>4</sup> K. H</u> halo of the Milky Way

Machine Learning based photometric redshifts for the KiDS ESO DR2 galaxies

S. Cavuoti<sup>1\*</sup>, M. Brescia<sup>1</sup>, C. Tortora<sup>1</sup>, G. Longo<sup>2</sup>, N. R. Napolitano<sup>1</sup>, M. Radovich<sup>3</sup>, F. L. Parbaral, M. Caracciali<sup>4</sup>, L. da Larab, E. Catarant, A. Cardal, M. Barkilla<sup>2</sup>

The masses of satellites in GAMA galaxy groups from 100 square degrees of KiDS weak lensing data

Cristóbal Sifón<sup>1\*</sup>, Marcello Cacciato<sup>1</sup>, Henk Edo van Uitert<sup>2,3</sup>, Massimo Viola<sup>1</sup>, Ivan Bald Ami Choi<sup>7</sup>, Simon P. Driver<sup>8,9</sup>, Thomas Erber

#### The first and second data releases of the Kilo-Degree Survey

Jelte T. A. de Jong<sup>1</sup>, Gijs A. Verdoes Kleijn<sup>2</sup>, Danny R. Boxhoorn<sup>2</sup>, Hugo Buddelmeijer<sup>2</sup>, Massimo Capaccioli<sup>3</sup>, Fedor Getman<sup>3</sup>, Aniello Grado<sup>3</sup>, Ewout Helmich<sup>1</sup>, Zhuoyi Huang<sup>3</sup>, Nancy Irisarri<sup>1</sup>, Konrad Kuijken<sup>1</sup>, Francesco LaBarbera<sup>3</sup>, John P. McFarland<sup>2</sup>, Nicola R. Napolitano<sup>3</sup>, Mario Radovich<sup>4</sup>, Gert Sikkema<sup>2</sup>, Edwin A. Valentijn<sup>2</sup>, Kor G. Begeman<sup>2</sup>, Massimo Brescia<sup>3</sup>, Stefano Cavuoti<sup>3</sup>, Ami Choi<sup>5</sup>, Oliver-Mark Cordes<sup>6</sup>, Giovanni Covone<sup>7</sup>, Massimo Dall'Ora<sup>3</sup>,

itert<sup>4,5</sup>, M. Alpaslan<sup>12</sup>, I.K. Baldry<sup>6</sup>, A. Choi<sup>7</sup>,

go<sup>2</sup> F Getman<sup>1</sup> M Capaccioli<sup>2</sup> L Grado<sup>1</sup>

Proefschrift

Summary of first-lensing data

- 109 square degrees of lensing + photo-z data
  - data observed before Oct 2013 that overlap with GAMA
  - 75 sq.deg. after conservative masking



8.9 galaxies arcmin<sup>-2</sup> with shape measurement

 $--> n_{eff}=4.5 \text{ arcmin}^{-2}$ 

Reiko Nakajima

--> shear inverse variance density = 69 arcmin<sup>-2</sup>.

median redshift 0.63, —> 0.53 with lens weight.

# KiD5 Lensing data quality

- Good seeing helps precision
  - $\sigma(e) \sim [1 + (r_{PSF}^2/r_{gal}^2)]^{1/2}$  at given flux, background
- so does a benign PSF pattern
- Various systematic tests for *accuracy*:
  - print-through of PSF ellipticity pattern
  - scaling of galaxy lensing with redshift
  - dependence of lensing signal on signal-to-noise ratio
  - clustering of photo-z bins
  - 2-pt correlation function B-modes

# **Test PSF print-through in shears**

- Use sub-exposures for each field to disentangle PSF and galaxy contributions to fitted ellipticities (cf CFHTLenS: Heymans et al 2012).
- Some chance alignment expected - check with mocks
- Consistent!
- (CFHTLenS had to reject 25% of fields)



### KiDS Galaxy lensing source redshift scaling



 $z_{B}$ 

- Spectroscopic lens samples from GAMA
- Photo-z sources in bins of best-fit photo-z
- Use full p(z) posteriors to predict signal in each bin

#### Galaxy lensing at different SNR



- galaxy-galaxy lensing by lenses with 20<r<21</li>
- compare CFHTLenS vs KiDS, each corrected for own noise bias

#### KiDS Clustering of photo-z bins

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#### KiDS Clustering of photo-z bins

0.005<z\_<0.1

Kids  $\times$  Kids

#### Ami Choi





#### KiDS E-B mode in 2-pt shear correlation



No sign of B mode, as expected when no systematic residuals remain

#### KiDS Yes! We have no bananas

 E-mode 2pt correlation function, compared to Planck-EE, Planck-TT, CFHTLenS





- Focussed on synergy with GAMA survey
  - 4 patches of 60 sq.deg., >95% complete redshifts to r=19.8
  - Group catalogue (Robotham et al 2011), zm~0.2
  - Weigh groups and galaxies with lensing

#### KiDS GAMA groups (Viola, Cacciato, Brouwer, KK et al 2015)





#### GAMA groups (Viola, Cacciato, Brouwer, KK et al 2015)



#### KiDS Satellites in groups (Sifón, Cacciato, Hoekstra, et al 2015)



KiDS Satellites in groups (Sifón, Cacciato, Hoekstra, et al 2015)





- Galaxy morphology studies (2m deeper than SDSS)
- Milky Way halo (star /gal separation to faint levels)
- Cluster searches
- z~6 QSO search (with VIKING)
- Asteroids (dithered observations)

#### KiDS Follow-up / external data sets

- GAMA fantastic resource
  Edo van Uitert
- 2dfLenS @ AAT 100k redshifts in KiDS BOSS footprint
- (ATLAS photometry for target selection)
   Dominik Klaes

  WAVES @ 4MOST
  - (GAMA-like survey over KiDS-S footprint)

Simon Driver, Joe Liske

Chris Blake

• ACTpol overlap with KiDS-N

## KiDS Summary

- KiDS delivering lensing science (1507.00735, 37, 38, 42)
  - Iensing/photo-z catalogues available <u>kids.strw.leidenuniv.nl</u>
- depth ~ DES; seeing 0.7"; PSF benign; ugri (+ ZYJHK)
- systematic/null tests check out
- lensfit/THELI shapes, BPZ/GAaP photometric redshifts
- 500 sq.deg. observed; 1500 target
- First results `weighing' GAMA structures
- Ready for cosmology!





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# Thanks.