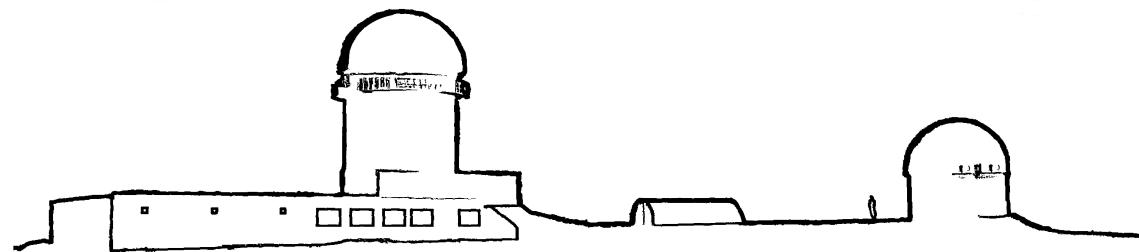


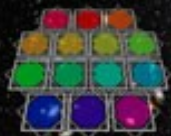
Javalambre  
Physics of the Accelerating Universe  
Astrophysical  
Survey



# The J-PAS Survey

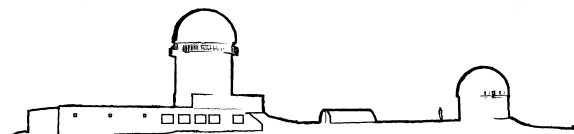
Silvia Bonoli

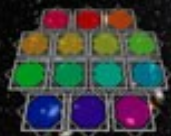




# The Javalambre-PAU Astrophysical Survey

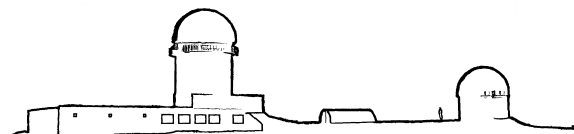
**A Spanish-Brazilian collaboration,**  
the J-PAS survey will scan  
~8500 deg<sup>2</sup> of the northern sky  
with 54 narrow-band filters covering  
the whole optical range  
from the dedicated 2.5m telescope in  
the *sierra of Javalambre*

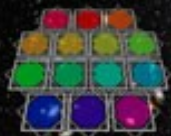




# The Javalambre-PAU Astrophysical Survey

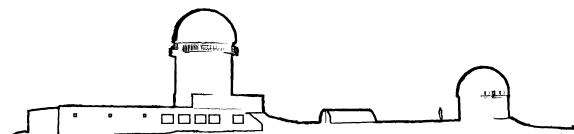
A Spanish-Brazilian collaboration,  
the J-PAS survey will scan  
 $\sim 8500 \text{ deg}^2$  of the northern sky  
with 54 narrow-band filters covering  
the whole optical range  
from the dedicated 2.5m telescope in  
the *sierra of Javalambre*





# The Javalambre-PAU Astrophysical Survey

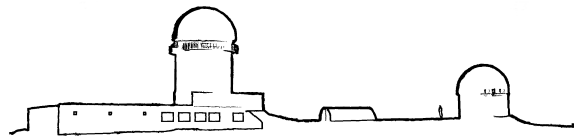
A Spanish-Brazilian collaboration,  
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with 54 narrow-band filters covering  
the whole optical range  
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the *sierra of Javalambre*

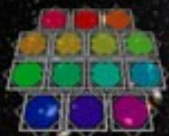




# The Javalambre-PAU Astrophysical Survey

A Spanish-Brazilian collaboration,  
the J-PAS survey will scan  
 $\sim 8500 \text{ deg}^2$  of the northern sky  
with 54 narrow-band filters covering  
the whole optical range  
from the dedicated 2.5m telescope in  
the *sierra of Javalambre*





# The filter system

## - 54 NB filters

(FWHM~145Å;  $\Delta\lambda\sim 10\text{nm}$   
From 3785Å to 9100Å

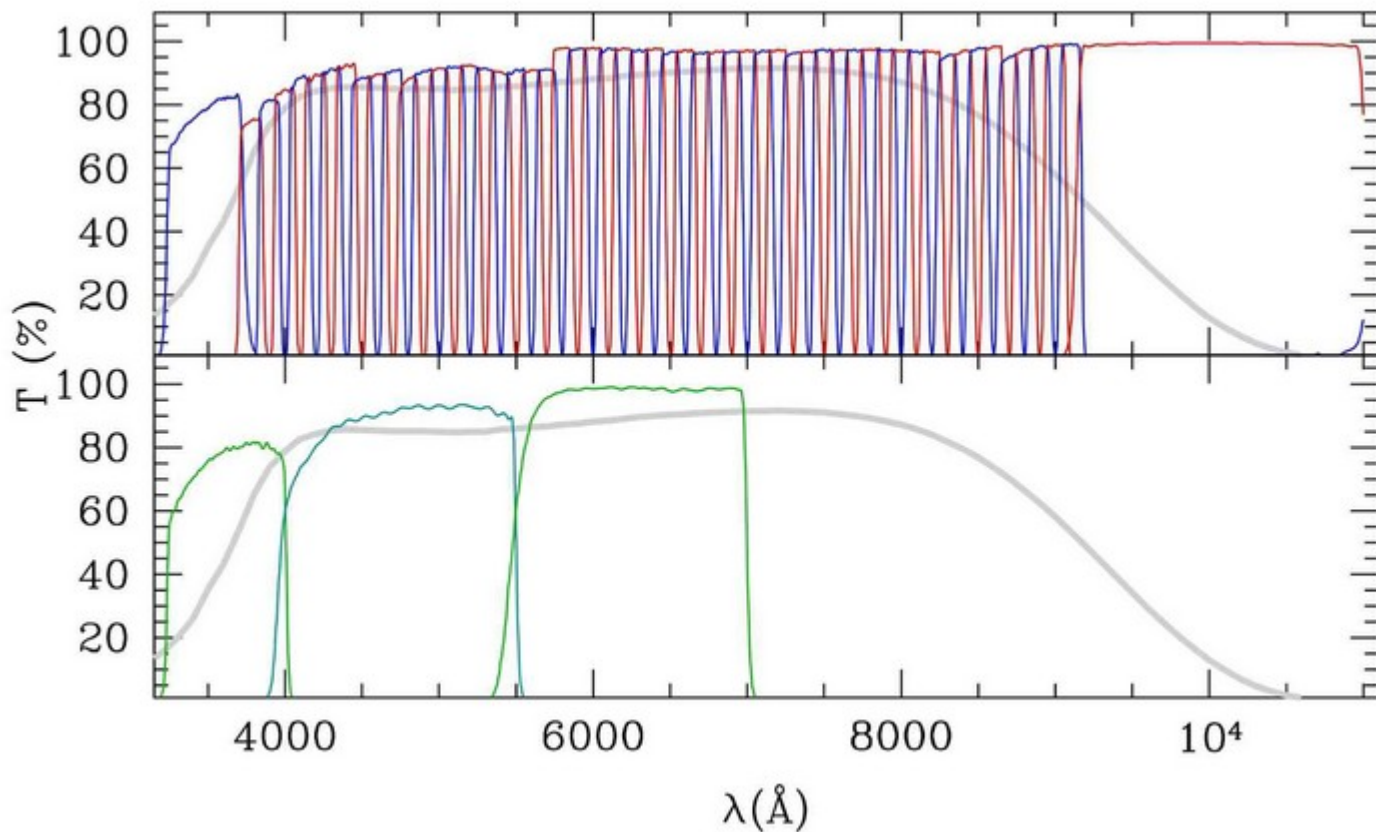
## - 1 Blue MB filter

(FWHM~260Å;  $\lambda_c\sim 3600\text{Å}$

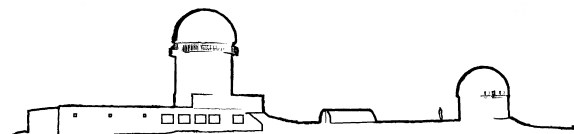
## - 1 Red BB filter

(FWHM~620Å;  $\lambda_c\sim 9500\text{Å}$

## - Sloan u, g, r



Pseudo-spectrum in every  
pixel of the sky down to  
 $M_{AB} \sim 22.5-24$



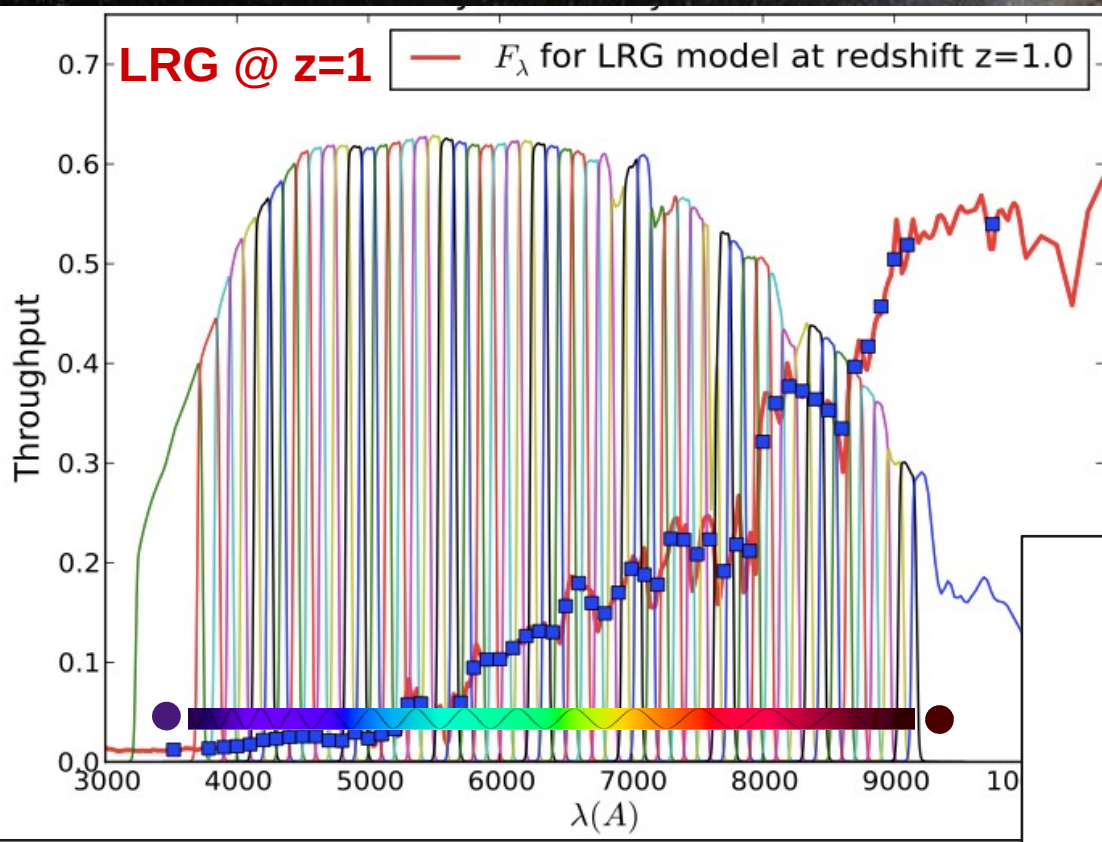
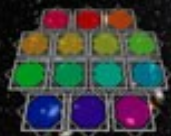
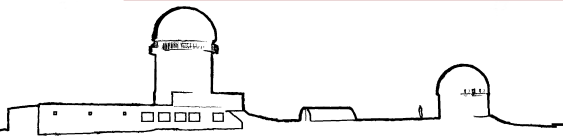
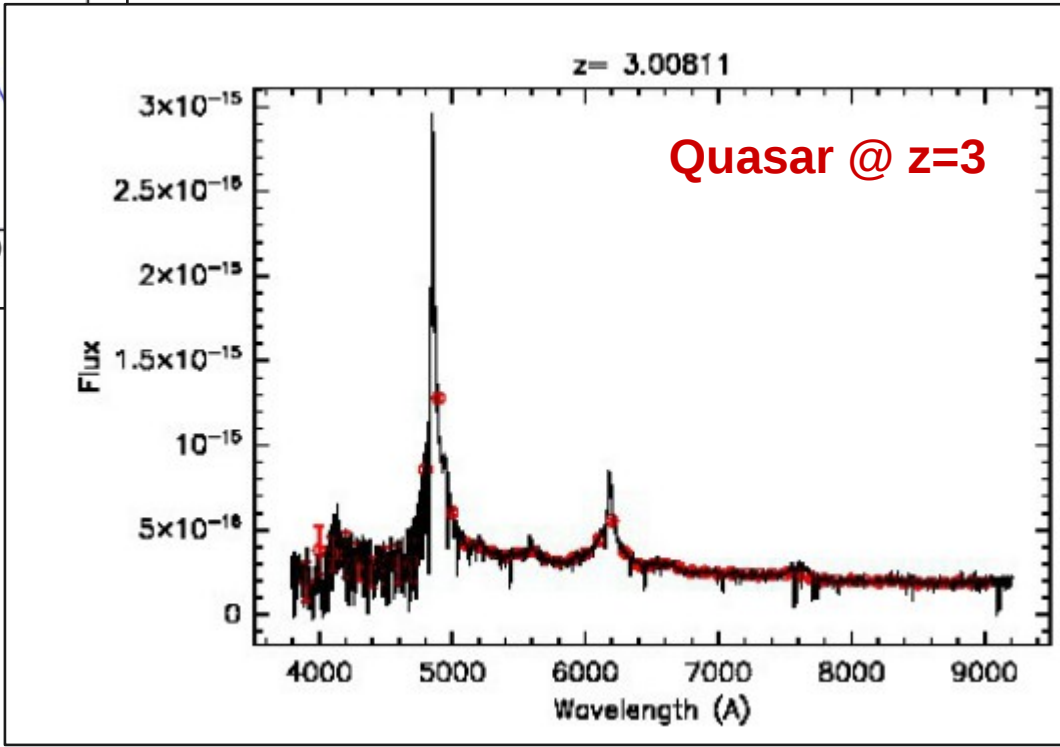


Photo-z precision as good as  $0.003(1+z)$

200M galaxies  
90M LRG+ELG  
Ms of quasars





# The camera JPCam

1.2 Giga pixels  
(14 CCD of 9200x9200)  
0.22 arcsec/pixel  
4.5 deg<sup>2</sup>

Actuator System  
**NTE SENER**

JPCam interface  
with T250

Filter and Shutter Unit

- Filter Unit
- Shutter Unit
- Anti-cond. Sys.
- e-Box

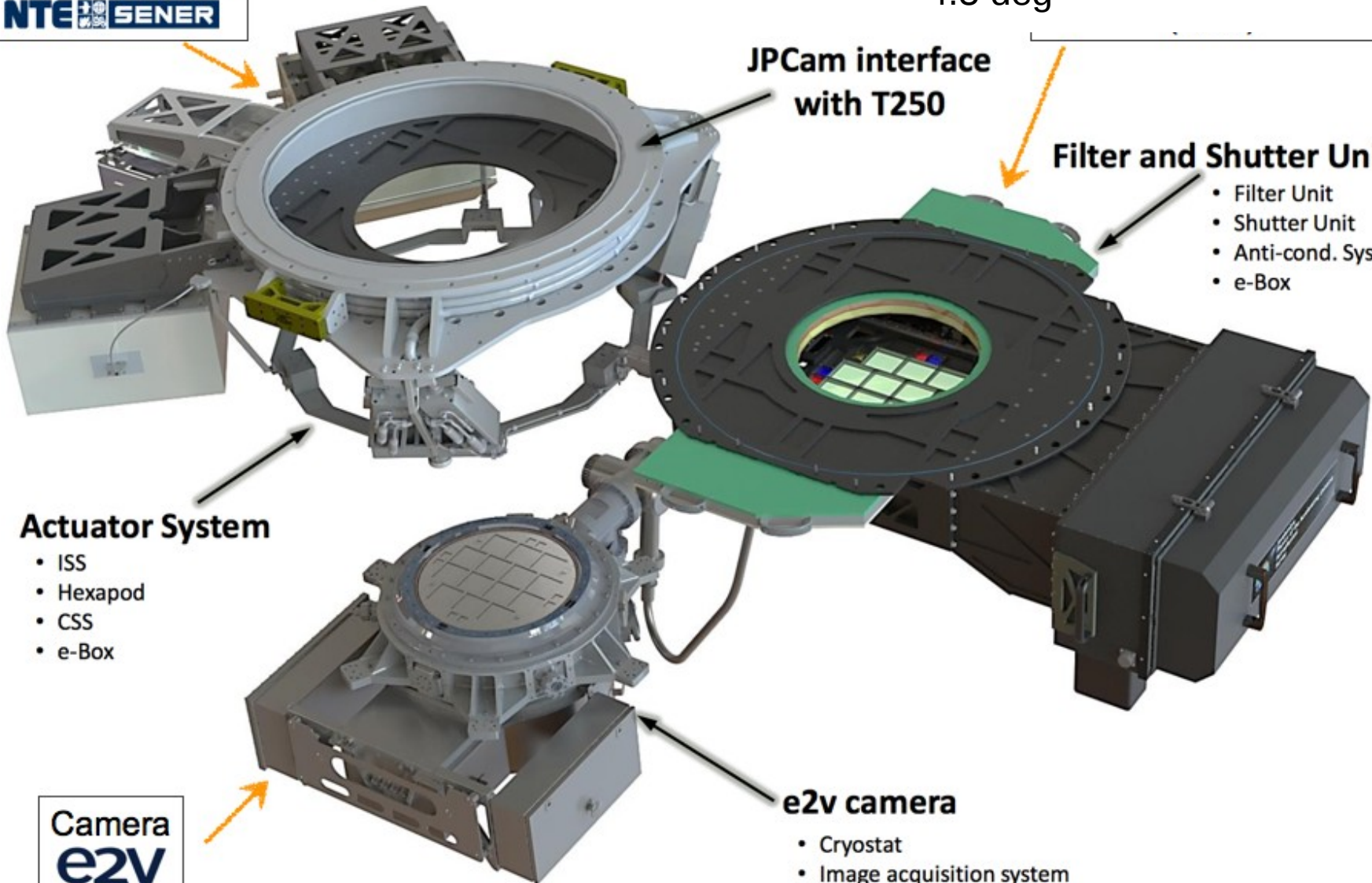
Actuator System

- ISS
- Hexapod
- CSS
- e-Box

Camera  
**e2v**

e2v camera

- Cryostat
- Image acquisition system



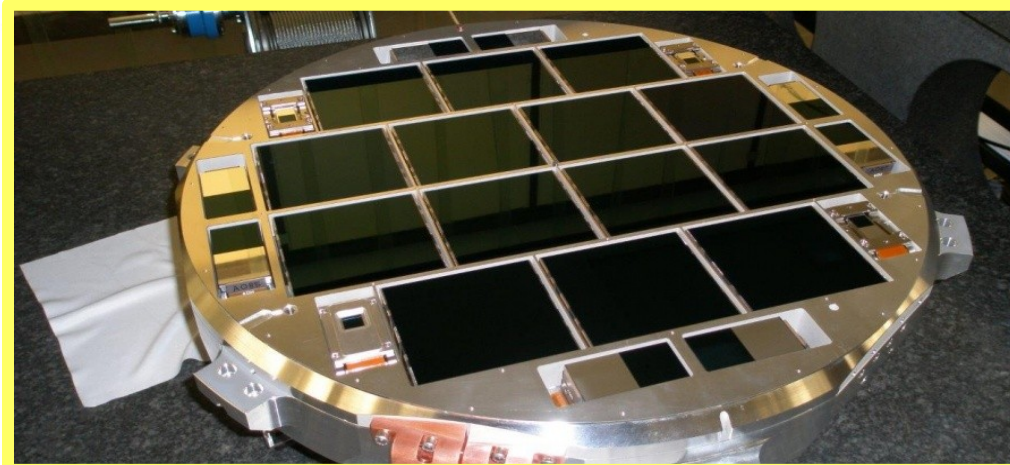
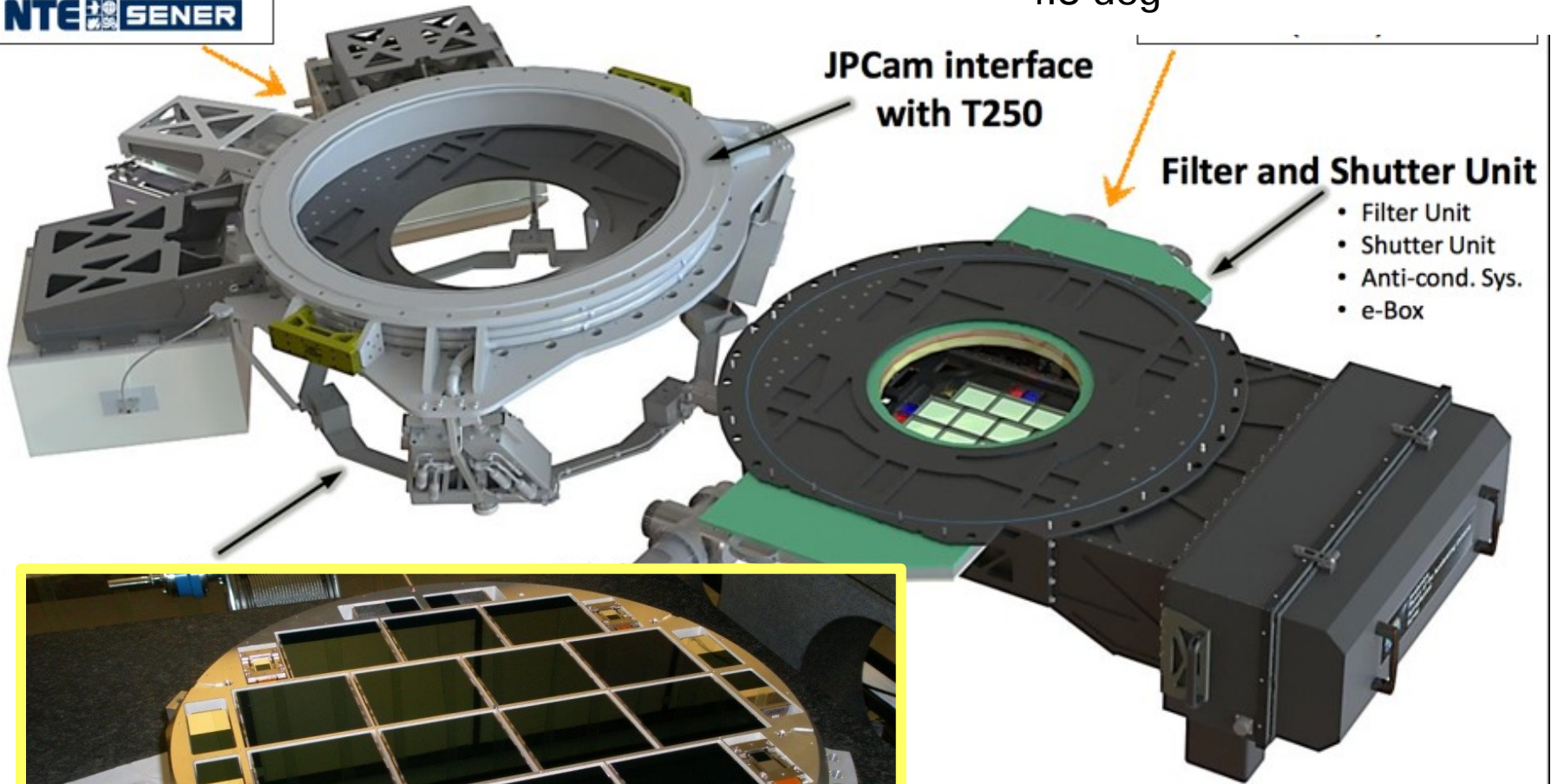




# The camera JPCam

1.2 Giga pixels  
(14 CCD of 9200x9200)  
0.22 arcsec/pixel  
4.5 deg<sup>2</sup>

Actuator System  
**NTE SENER**



**v camera**

Cryostat  
Image acquisition system

# The telescope

T250

M1 ( $\varnothing$ ) = 2.55 m

FoV ( $\varnothing$ ) = 3 deg = 476 mm at FP

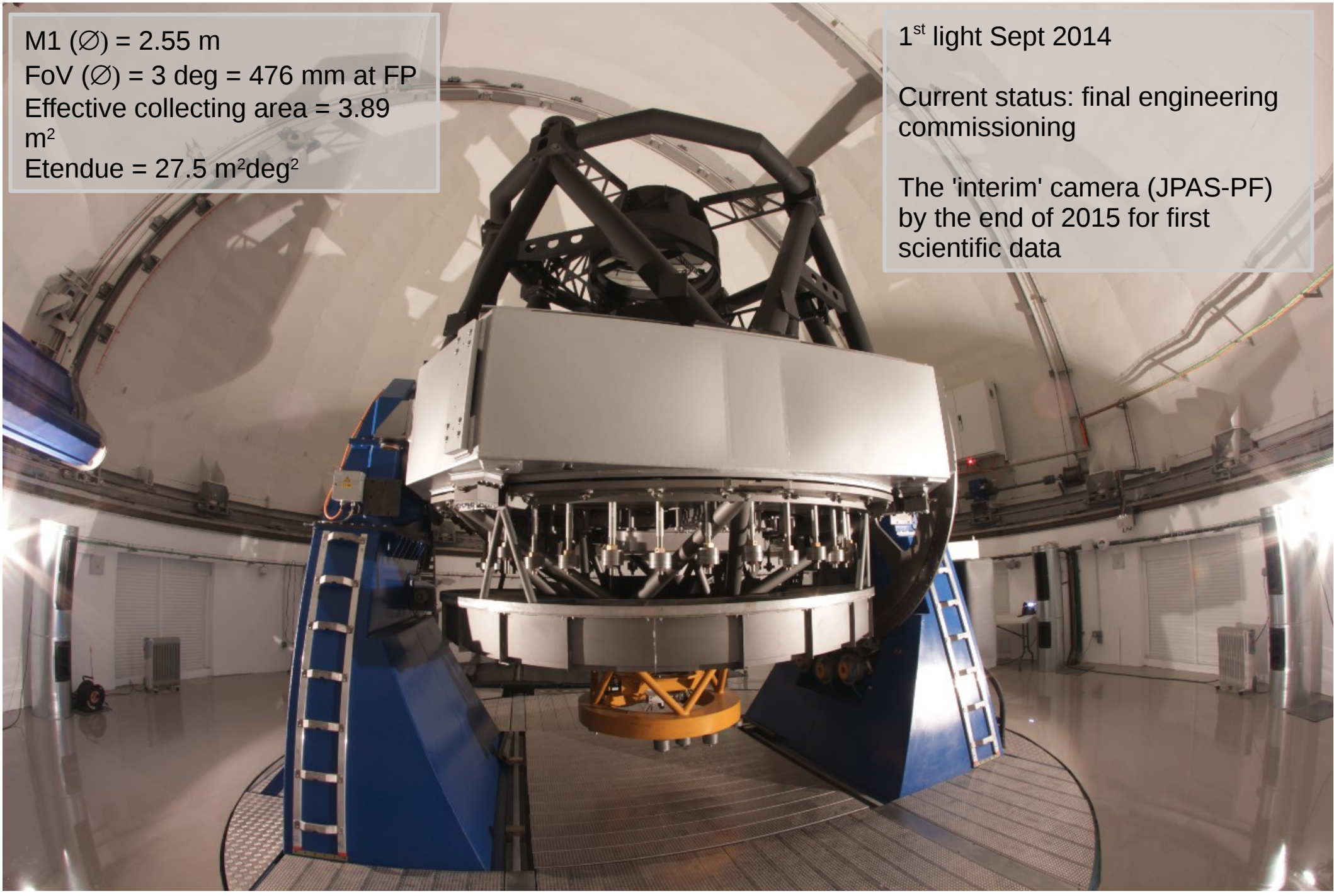
Effective collecting area = 3.89 m<sup>2</sup>

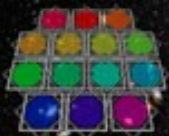
Etendue = 27.5 m<sup>2</sup>deg<sup>2</sup>

1<sup>st</sup> light Sept 2014

Current status: final engineering commissioning

The 'interim' camera (JPAS-PF) by the end of 2015 for first scientific data





Javalambre  
Physics of the Accelerating Universe  
Astrophysical  
Survey

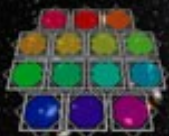


# The Javalambre Observatory

In the “Sierra de Javalambre” @1960m

now officially a Spanish “scientific and technical facility” (soon available for 20% open-time)



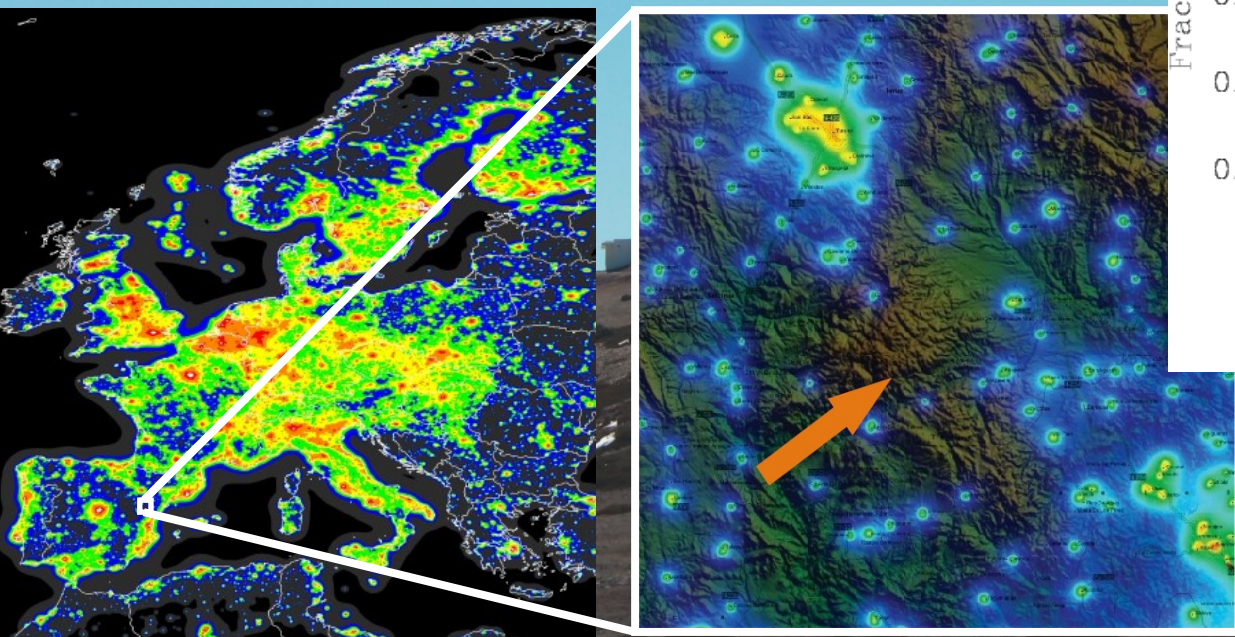
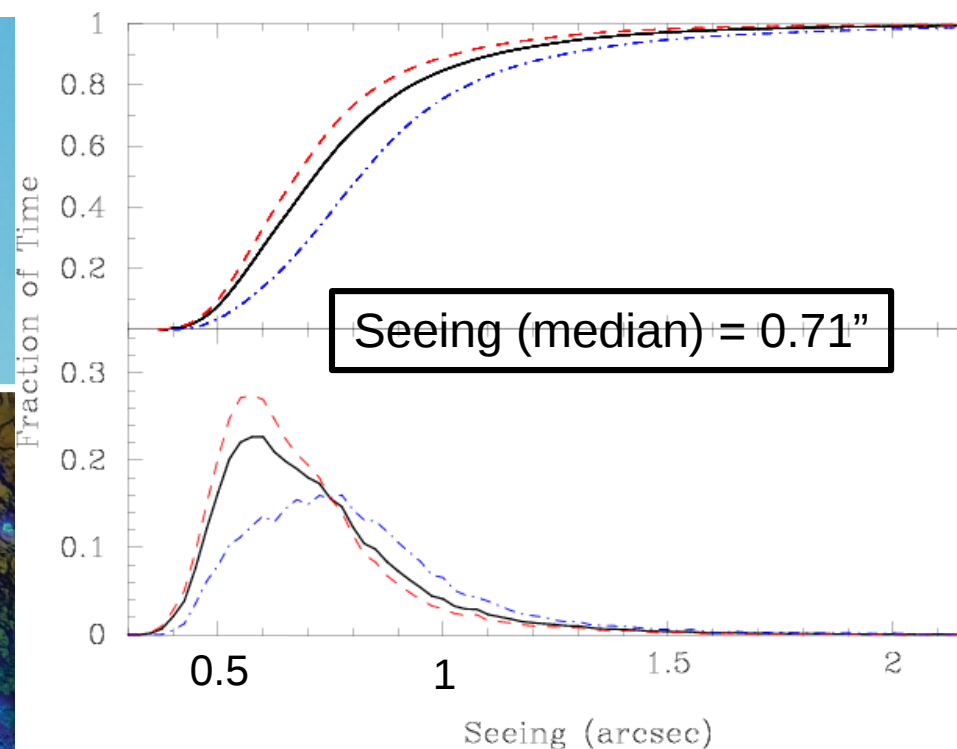


# The Javalambre Observatory

In the “Sierra de Javalambre” @1960m

now officially a Spanish “scientific and technical facility” (soon available for 20% open-time)

Comparable (at similar altitudes) to  
Mauna Kea or La Silla



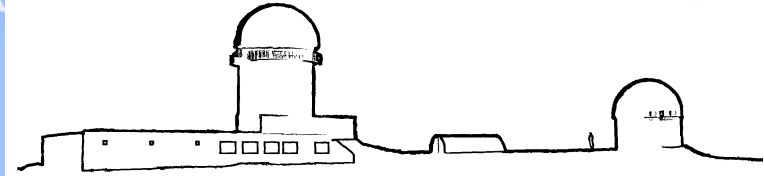
# Location



Teruel

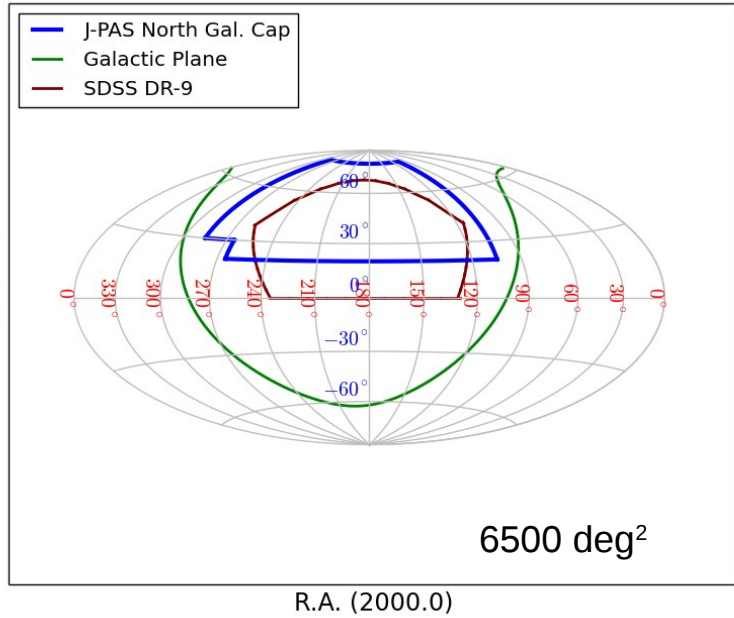


Javalambre

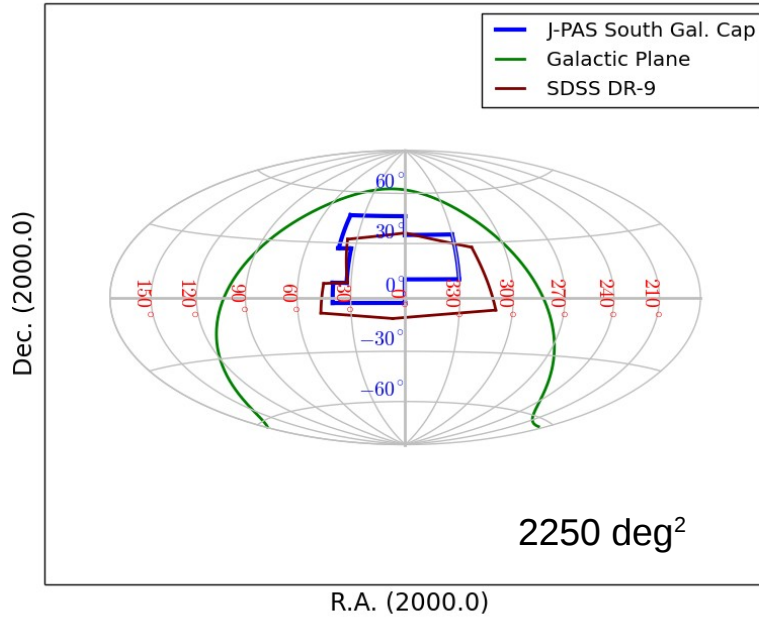


# Footprint and Survey strategy

Northern Galactic Hemisphere

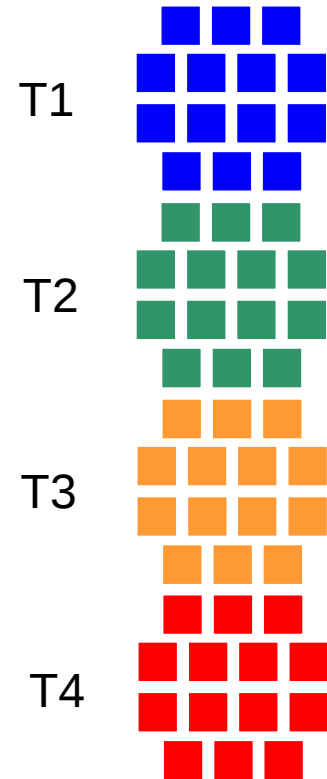
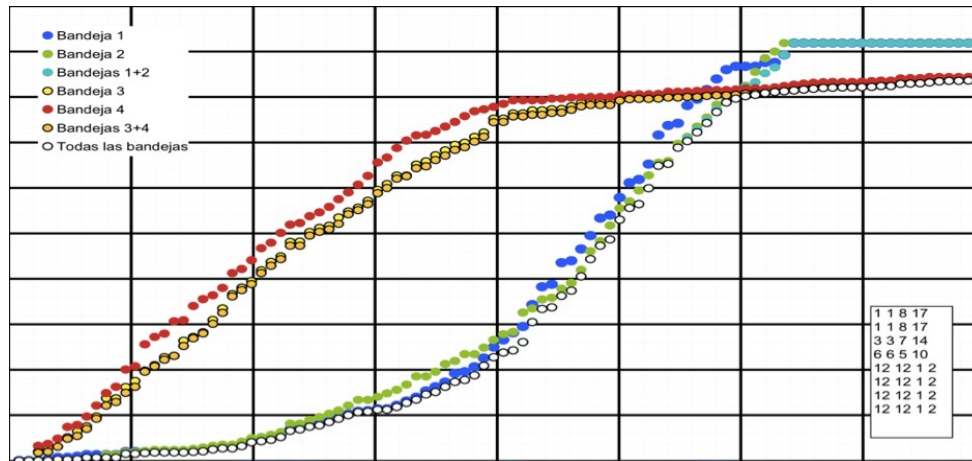


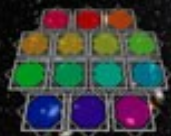
Southern Galactic Hemisphere



Compromise  
visibility from the  
OAJ and galactic  
extinction

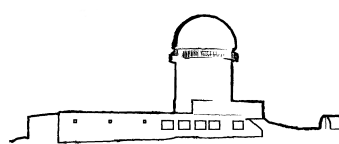
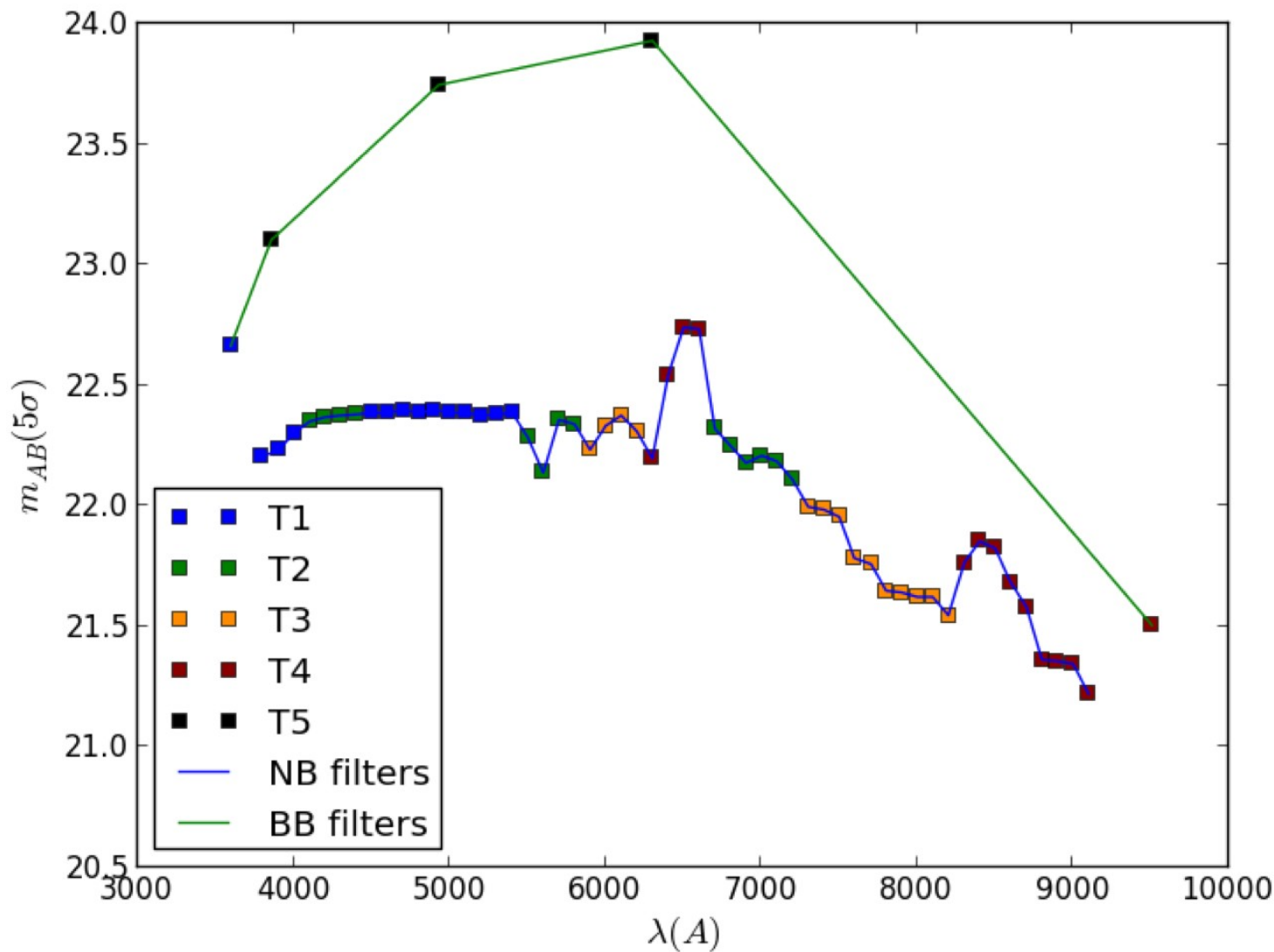
14 different filters in each tray  
at least 4 exposures in each filter





# Limiting magnitude

5 $\sigma$   
3arcsec aperture



# Data processing and storage

**J-PAS: 1.3Tb** of data per observing night

Data reduction, catalogs and storage  
managed @ **CEFCA**





**Type Ia  
Supernovae**

~4000 SNIa

**BAO**

90M galaxies (LRG,  
ELG, QSO)  
With photo-z precision  
of 0.003

**Clusters**

700k clusters

Sub arcs  
seeing

**Weak Lensing**

Type Ia  
Supernovae

BAO

~4000 SNIa

And...

90M galaxies  
With photo-z precision  
of 0.003

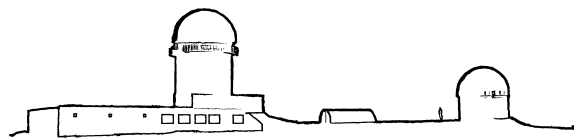
- Galaxy evolution
- AGN and quasars
- “IFU”-type science for extended objects
  - Milky-Way
- Transient objects
- ...

700K clusters

Sub-arcsec  
seeing

Clusters

Weak Lensing

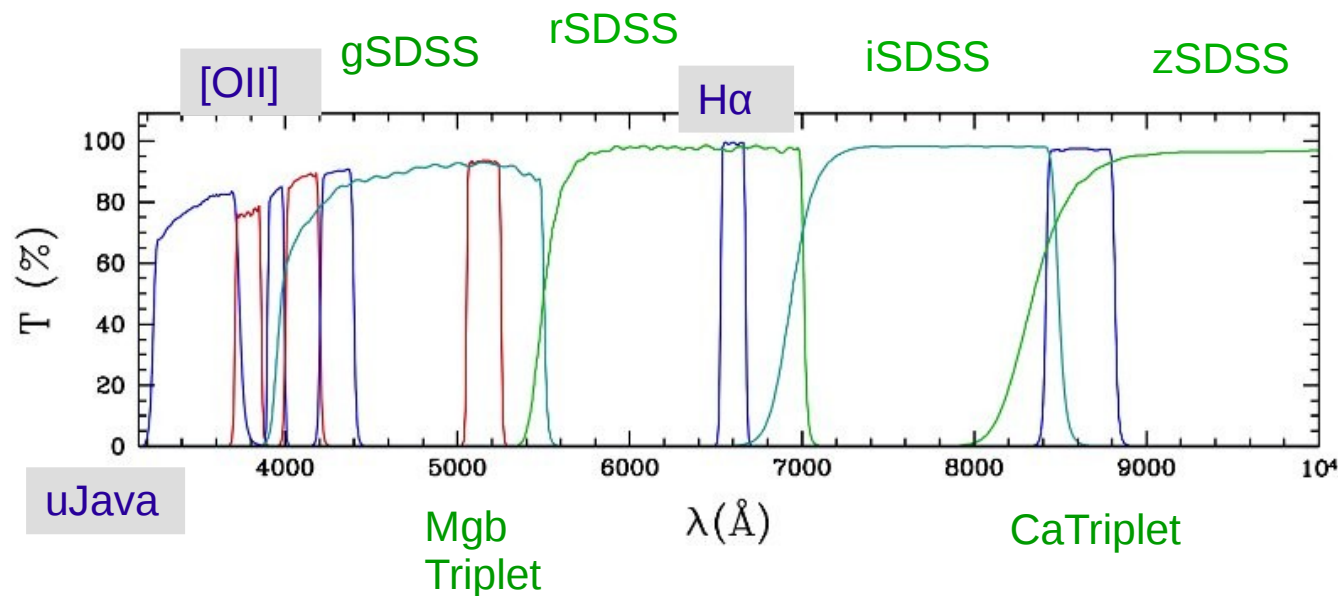
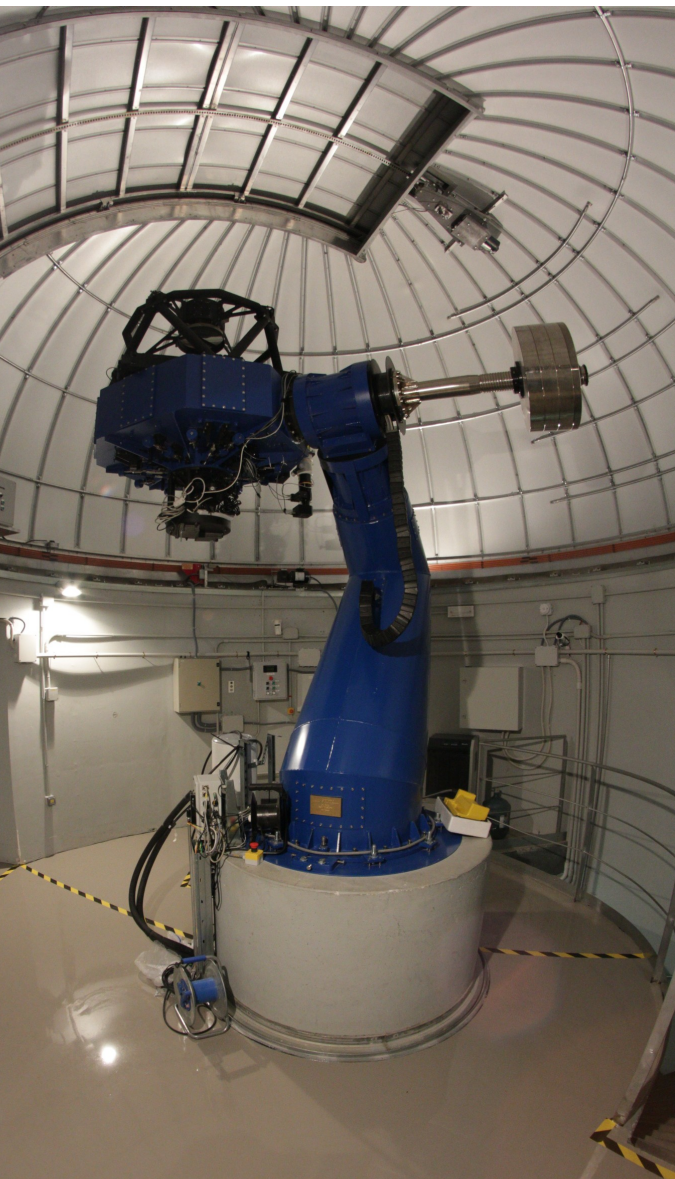


# The J-PLUS survey

*Javalambre Photometric Local Universe Survey*



## T80



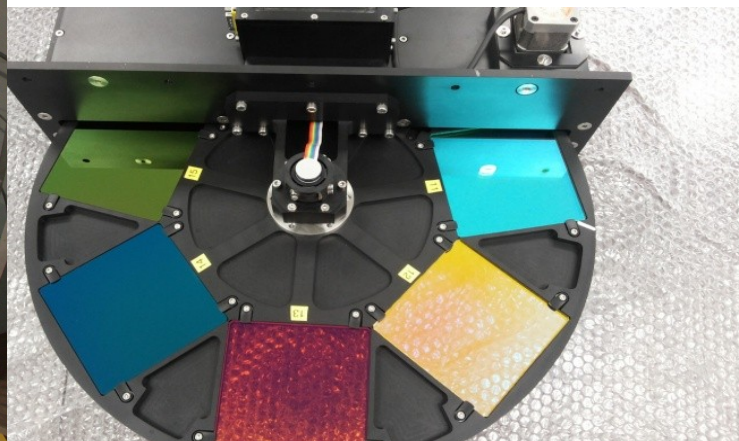
12 filters

1 CCD of 9200x9200 pixels

0.55 arcsec/pixel

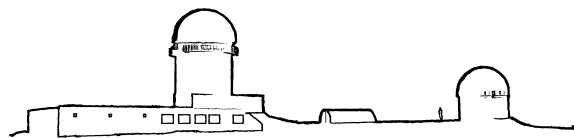
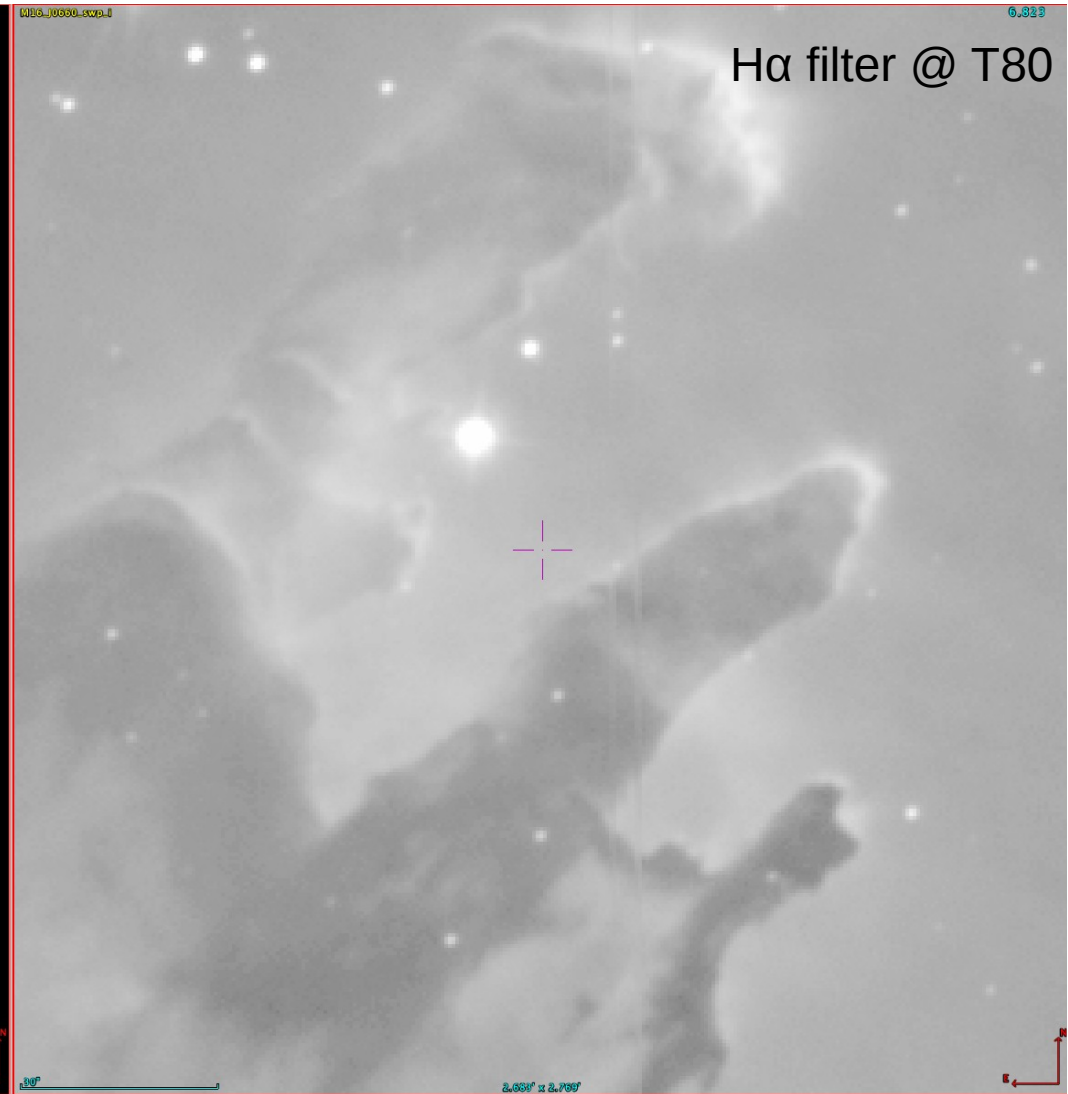
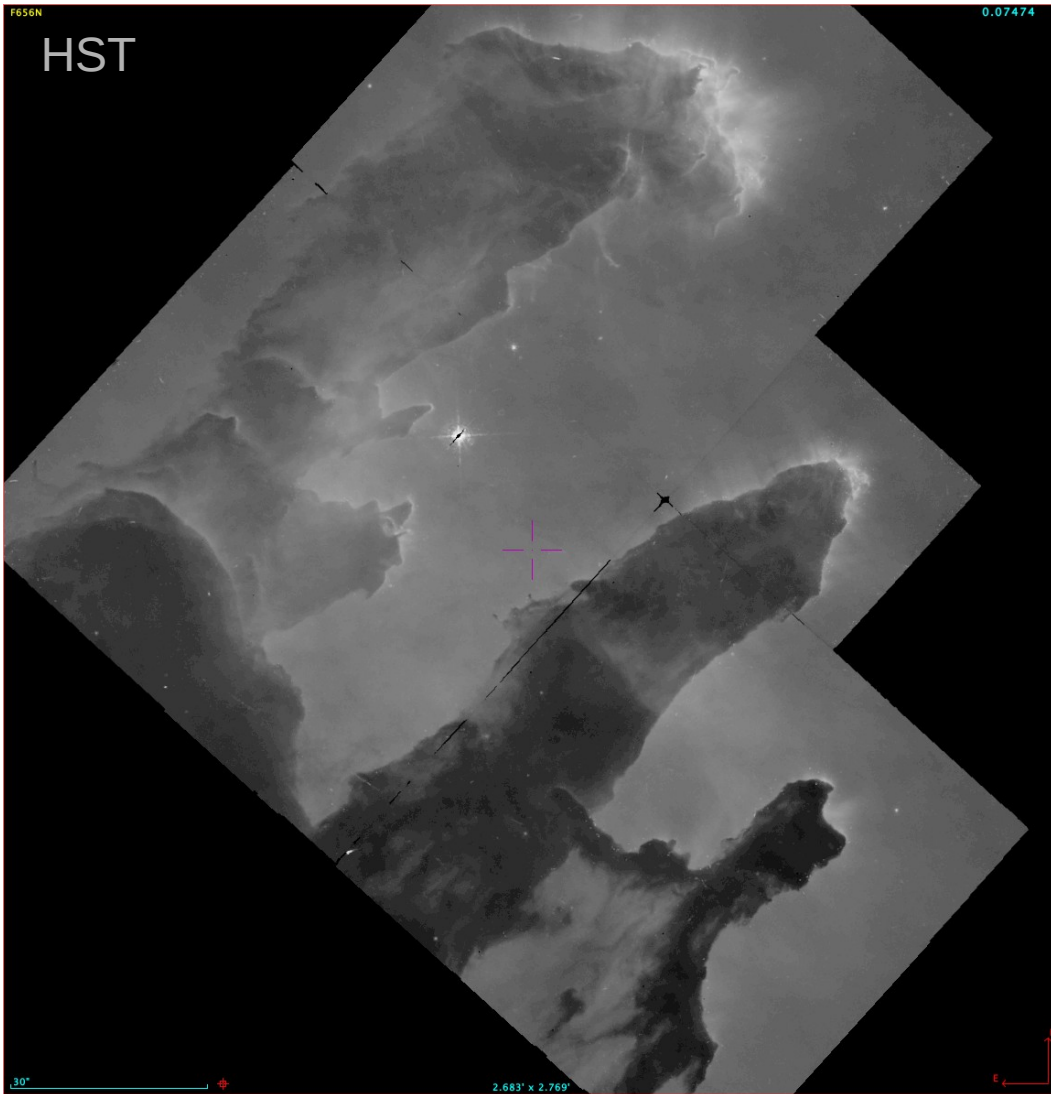
1.4x1.4 deg<sup>2</sup>

~ 1 mag deeper than SDSS



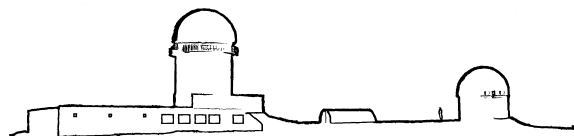
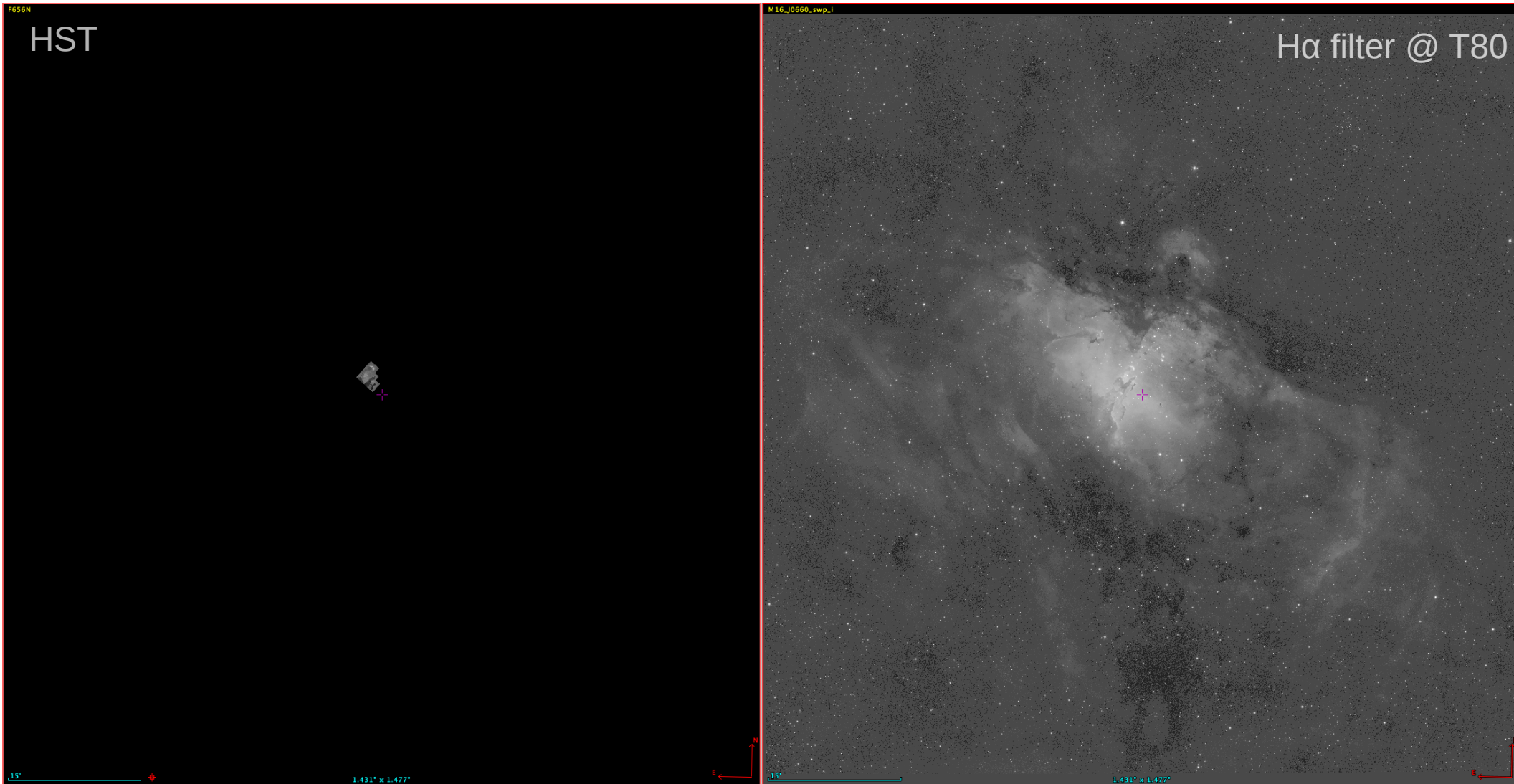
# T80 + T80Cam

M16



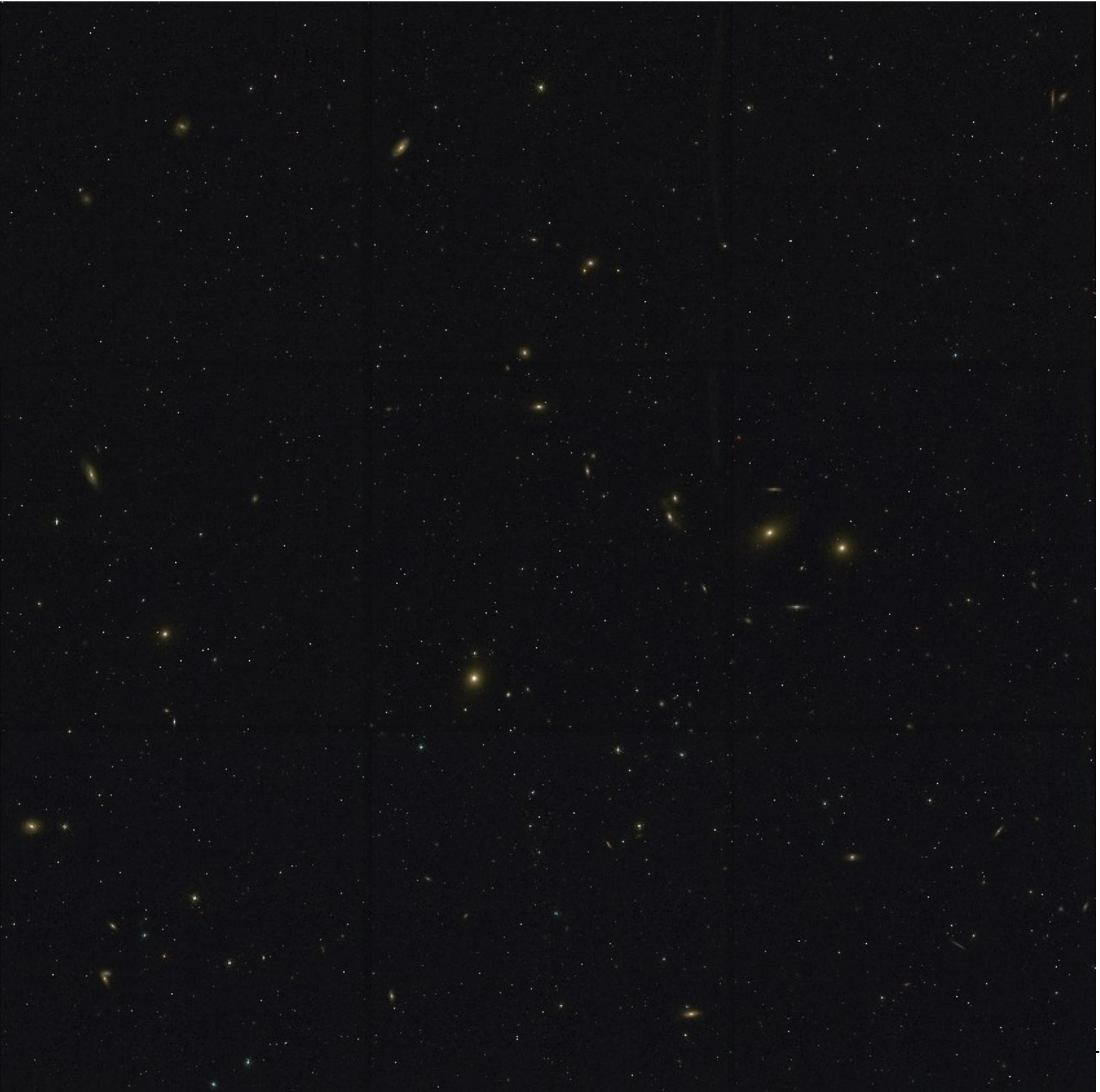
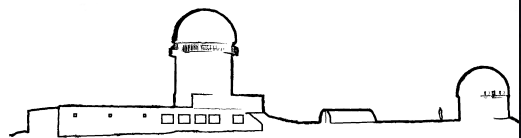
# T80 + T80Cam

M16



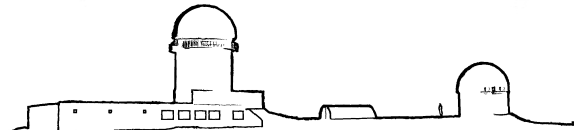
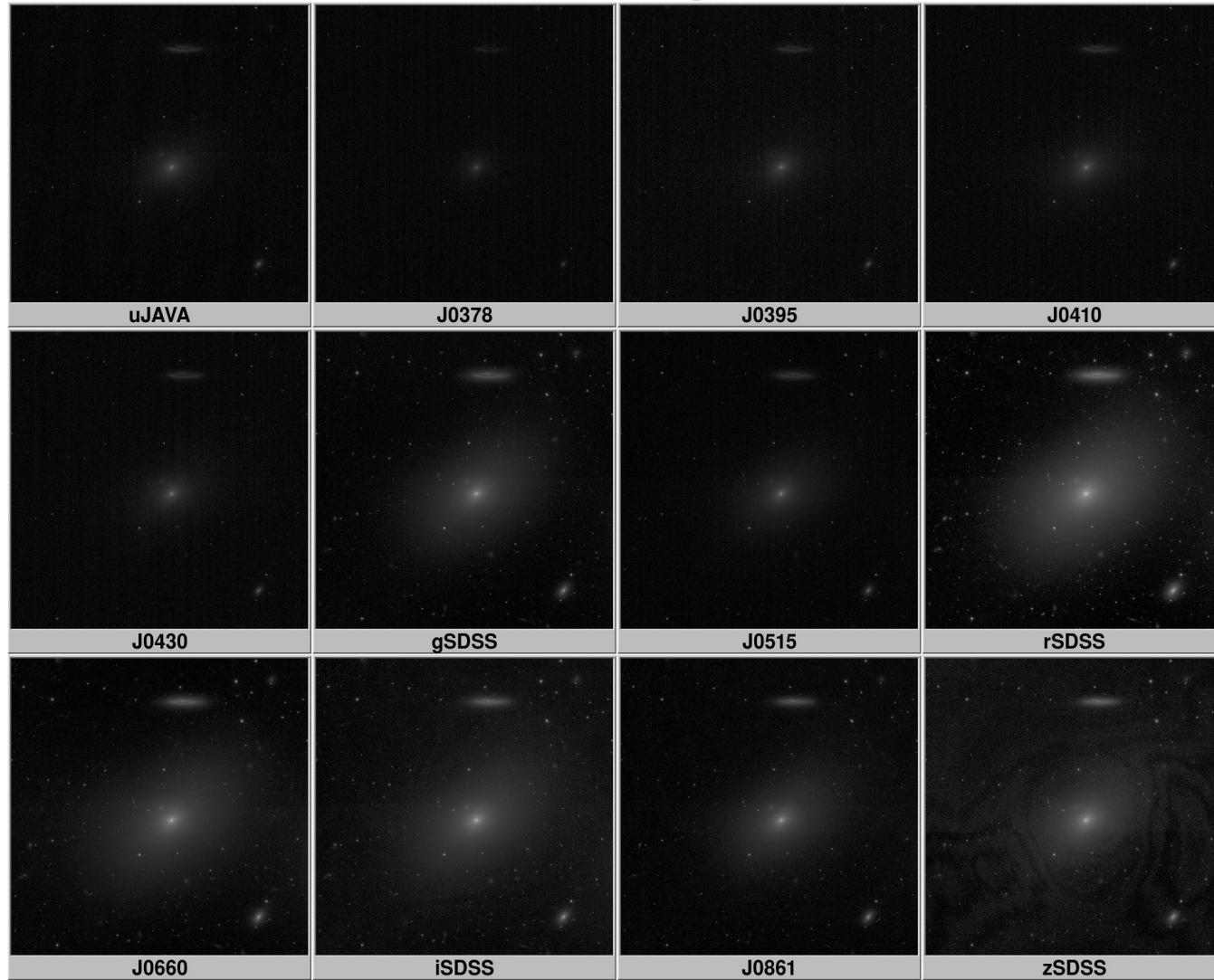
# T80 + T80Car

Mosaic of Virgo



# T80 + T80Cam

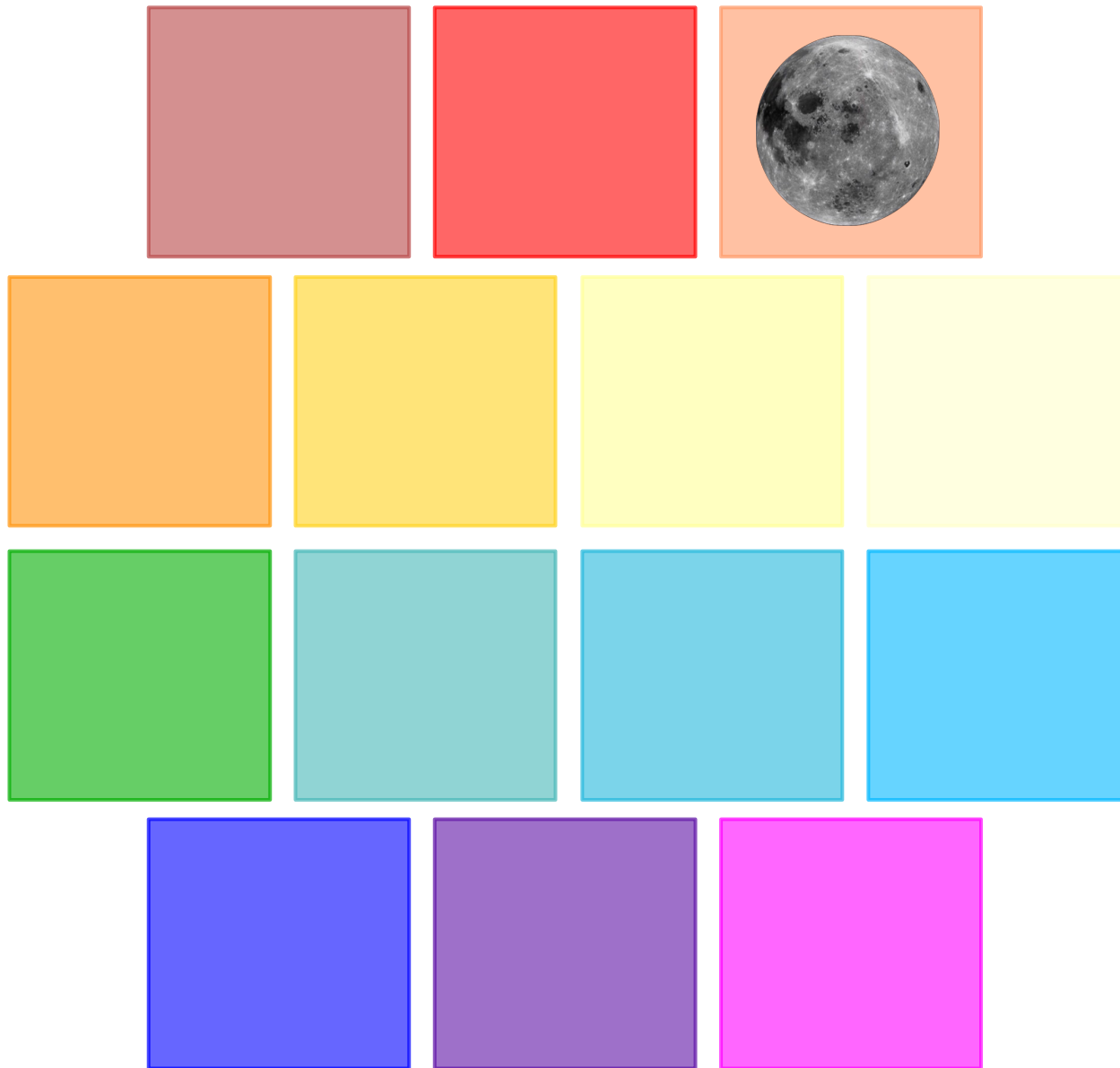
## M86 - T80cam@JAST





# JPCam

# T80Cam

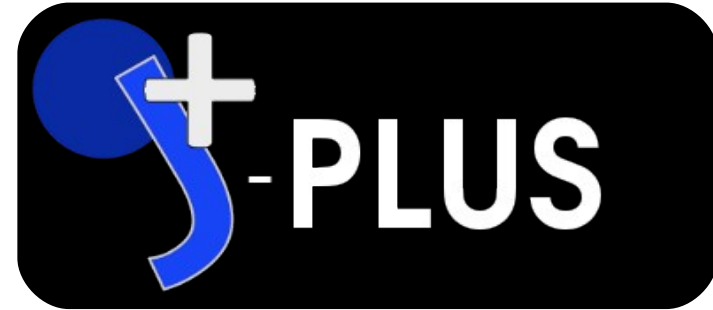


8500 deg<sup>2</sup> observed from Spain with 2 telescopes



- 54 NB + 5 MB/BB filters
- 4.5 deg<sup>2</sup> FoV
- Up to mag ~**24**
- 90M ELG and LRG
- Reaching **0.3%** photo-z precision
- Millions of quasars
- 200M of galaxies
- 4000 SNIa
- 700k of groups and clusters
- Starts in **2016**

Possible replica in the South



- 12 NB/MB/BB filters
- 2 deg<sup>2</sup> FoV
- Up to mag ~**23**
- ~ 1 mag deeper than SDSS
- SEDs of 5M stars
- Starts in the fall of **2015**

Already replica in the South  
@Cerro Tololo  
S-PLUS – start in 2016



## J-PAS: The Javalambre-Physics of the Accelerated Universe Astrophysical Survey

[arXiv:1403.5237](https://arxiv.org/abs/1403.5237)

N. Benítez<sup>a,b</sup>, R. Dupke<sup>b,c,d</sup>, M. Moles<sup>e,a</sup>, L. Sodré<sup>f</sup>, J. Cenarro<sup>e</sup>, A. Marín-Franch<sup>e</sup>, K. Taylor<sup>b</sup>, D. Cristóbal<sup>e</sup>,  
A. Fernández-Soto<sup>g</sup>, C. Mendes de Oliveira<sup>f</sup>, J. Cepa-Nogué<sup>h</sup>, L.R. Abramo<sup>i</sup>, J.S. Alcaniz<sup>b</sup>, R. Overzier<sup>b</sup>,  
C. Hernández-Monteagudo<sup>e</sup>, E. J. Alfaro<sup>a</sup>, A. Kanaan<sup>j</sup>, J. M. Carvano<sup>b</sup>, R.R.R. Reis<sup>k</sup>, E. Martínez González<sup>l</sup>,  
B. Ascaso<sup>a</sup>, F. Ballesteros<sup>g</sup>, J. Varela<sup>e</sup>, H.S. Xavier<sup>i</sup>, T. Broadhurst<sup>n</sup>, E. Cypriano<sup>f</sup>, R. Angulo<sup>e</sup>, J. M. Diego<sup>l</sup>,  
A. Zandivarez<sup>o</sup>, E. Díaz<sup>o</sup>, P. Melchior<sup>p</sup>, K. Umetsu<sup>q</sup>, P. F. Spinelli<sup>r</sup>, A. Zitrin<sup>s</sup>, D. Coe<sup>an</sup>, G. Yepes<sup>t</sup>, P. Vielva<sup>l</sup>,  
V. Sahni<sup>u</sup>, A. Marcos-Caballero<sup>l</sup>, F. Shu Kitaura<sup>v</sup>, A. L. Maroto<sup>w</sup>, M. Masip<sup>at</sup>, S. Tsujikawa<sup>x</sup>, S. Carneiro<sup>y</sup>,  
J. González Nuevo<sup>l</sup>, G. C. Carvalho<sup>b</sup>, M. J. Rebouças<sup>av</sup>, J. C. Carvalho<sup>b,z</sup>, E. Abdalla<sup>i</sup>, A. Bernui<sup>b</sup>,  
C. Pigozzo<sup>y</sup>, E.G. Ferreira<sup>i</sup>, N. Chandrachani Devi<sup>b</sup>, C.A.P. Bengaly Jr.<sup>b</sup>, M. Campista<sup>b</sup>, A. Amorim<sup>g</sup>,  
N. V. Asari<sup>aa</sup>, A. Bongiovanni<sup>h</sup>, S. Bonoli<sup>e</sup>, G. Bruzual<sup>ab</sup>, N. Cardiel<sup>l</sup>, A. Cava<sup>ac</sup>, R. Cid Fernandes<sup>j</sup>,  
P. Coelho<sup>ai</sup>, A. Cortesi<sup>f</sup>, R. G. Delgado<sup>a</sup>, L. Díaz Garcia<sup>e</sup>, J. M. R. Espinosa<sup>h</sup>, E. Galliano<sup>b</sup>,  
J. I. González-Serrano<sup>l</sup>, J. Falcón-Barroso<sup>h</sup>, J. Fritz<sup>ad</sup>, C. Fernandes<sup>b</sup>, J. Gorgas<sup>l</sup>, C. Hoyos<sup>e</sup>,  
Y. Jiménez-Teja<sup>a,b</sup>, J. A. López-Aguerri<sup>h</sup>, C. López-San Juan<sup>f</sup>, A. Mateus<sup>j</sup>, A. Molino<sup>a</sup>, P. Novais<sup>f</sup>, A. O'Mill<sup>f</sup>,  
I. Oteo<sup>h</sup>, B. Poggianti<sup>af</sup>, R. Proctor<sup>b</sup>, E. Ricciardelli<sup>g</sup>, P. Sánchez-Blázquez<sup>l</sup>, T. Storchi-Bergmann<sup>ag</sup>,  
E. Telles<sup>b</sup>, W. Schoenell<sup>a</sup>, N. Trujillo<sup>h</sup>, A. Vazdekis<sup>h</sup>, K. Viironen<sup>e</sup>, S. Daflon<sup>b</sup>, T. Aparicio<sup>b</sup>, D. Rocha<sup>ah</sup>,  
A. Ederoclite<sup>e</sup>, H. Vázquez Ramió<sup>f</sup>, T. Ribeiro<sup>ai</sup>, M. Borges<sup>b</sup>, S. L. Martins<sup>ah</sup>, W. Marcolino<sup>ah</sup>, D.  
Martínez-Delgado<sup>i,aj</sup>, M.A. Pérez-Torres<sup>f</sup>, B.B. Siffert<sup>k</sup>, M.O. Calvão<sup>k</sup>, M. Sako<sup>m</sup>, R. Kessler<sup>ak</sup>,  
A. Álvarez-Candal<sup>b</sup>, M. De Prá<sup>b</sup>, F. Roig<sup>b</sup>, D. Lazzaro<sup>b</sup>, J. Gorosábel<sup>a</sup>, R. Lopes de Oliveira<sup>al</sup>,  
G. B. Lima-Neto<sup>f</sup>, J. Irwin<sup>d</sup>, J. F. Liu<sup>aj</sup>, E. Álvarez<sup>t</sup>, I. Balmés<sup>i</sup>, A. A. da Costa<sup>f</sup>, S. Chueca<sup>e</sup>, A. Y. Díaz<sup>e</sup>,  
M. C. Díaz-Martín<sup>e</sup>, M. V. C. Duarte<sup>i</sup>, J. Fabregat<sup>g</sup>, F. Ferrari<sup>ao</sup>, B. Gavela<sup>t</sup>, S. G. Gracia<sup>f</sup>, N. Gruel<sup>ae</sup>,  
J. L. L. Gutiérrez<sup>f</sup>, R. Guzmán<sup>ap</sup>, J. D. Hernández-Fernández<sup>e</sup>, D. Herranz<sup>h</sup>, L. Hurtado-Gil<sup>q</sup>, F. Jablonsky<sup>au</sup>,  
R. Laporte<sup>au</sup>, J. Licandro<sup>h</sup>, M. Lima<sup>i</sup>, E. Martín<sup>aq</sup>, V. Martínez<sup>g</sup>, J. J. C. Montero<sup>f</sup>, P. Penteado<sup>f</sup>, C.B. Pereira<sup>b</sup>,  
V. Peris<sup>g</sup>, V. Quilis<sup>g</sup>, N. M. Sacristán<sup>f</sup>, M. Sánchez-Portal<sup>ar</sup>, A. C. Soja<sup>f</sup>, E. Solano<sup>ao</sup>, J. Torra<sup>as</sup>, L. Valdivielso<sup>e</sup>