# Perturbation Theory and large deviation functions in cosmology 



## Self gravitating fluids

- A multi-component formulation
- Dynamical equations (in Fourier space)

$$
\Psi_{a}(\mathbf{k}, \eta)=\left(\begin{array}{c}
\delta(\mathbf{k}, \eta) \\
\theta(\mathbf{k}, \eta) \\
\cdots
\end{array}\right)
$$

$$
\frac{\partial}{\partial \eta} \Psi_{a}(\mathbf{k}, \eta)+\Omega_{a}^{b}(\eta) \Psi_{b}(\mathbf{k}, \eta)=\gamma_{a}^{b c}\left(\mathbf{k}_{1}, \mathbf{k}_{2}\right) \Psi_{b}\left(\mathbf{k}_{1}, \eta\right) \Psi_{c}\left(\mathbf{k}_{2}, \eta\right)
$$

convolution is implicit

- Explicit results will be given here for a single-component pressureless fluid
- detailed effects of baryons versus DM can be taken into account (Somogyi \& Smith 20I0; FB, Van de Rijt, Vernizzi '12) with a 4-component multiplet
- Same structure also for non-interacting relativistic particles (neutrinos) with multiple flow description (Dupuy and FB, 'I4,'I5)
- Diagrammatic representation

$$
\delta_{n}(\mathbf{k})=\int \mathrm{d}^{3} \mathbf{q}_{1} \ldots \int \mathrm{~d}^{3} \mathbf{q}_{n} \delta_{D}\left(\mathbf{k}-\mathbf{q}_{1 \ldots n}\right) F_{n}\left(\mathbf{q}_{1}, \ldots, \mathbf{q}_{n}\right) \delta_{1}\left(\mathbf{q}_{1}\right) \ldots \delta_{1}\left(\mathbf{q}_{n}\right)
$$

$\odot=\frac{k}{g(n) \phi(k)} \longleftrightarrow$



- Ensemble averages by glueing diagrams together




## Charting PT

number of loops in standard PT for Gaussian
Initial Conditions

|  | Tree order <br> LO | I-loop <br> NLO | 2-loops <br> NNLO | 3-loops | ...p-loops |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 2-point <br> statistics | OK | OK | OK | partial exact <br> results | partial resum |
| 3-point <br> statistics | OK | OK (but not <br> systematics) |  |  | partial <br> resummations |
| 4-point <br> statistics | OK | OK (but not <br> systematics) |  |  |  |
| N-point <br> statistics | OK, in specific <br> geometris (counts <br> in cells) |  |  |  |  |

- Not a single way of doing PT calculations
- change of variables or fields : most dramatic is Eulerian to Lagrangian
- re-organisation(s) of the perturbation series (for instance with multipoint propagators introduced in FB, Crocce, Scoccimarro, PRD, 2008)
- PT can then come in many different flavors : SPT, RPT, TRG, RegPT, gRPT, MPT
$\rightarrow$ Power spectra up to I-loop and 2-loop order



## An alternative to the power spectra: response functions



Of direct interest from $\mathrm{P}(\mathrm{k})$ predictions:

$$
P_{\mathcal{M}_{2}}^{\mathrm{nl}}(k) \approx P_{\mathcal{M}_{1}}^{\mathrm{nl}}(k)+\int \frac{\mathrm{dq}}{q} \mathcal{R}_{\mathcal{M}_{1}}(k, q)\left[P_{\mathcal{M}_{2}}^{\mathrm{lin}}(q)-P_{\mathcal{M}_{1}}^{\mathrm{lin}}(q)\right]
$$

How good can PTs be at predicting response functions?

## first measurement of the response function

Nishimichi, FB, Taruya, '/ 4


## Comparison with I- and 2-loop results



$$
T^{\text {eff. }}(k, q)=\left[T^{1-\operatorname{loop}}(k, q)+T^{2-\operatorname{loop}}(k, q)\right] \frac{1}{1+\left(q / q_{0}\right)^{2}}
$$



- From PT perspective, UV regularization is necessary
- existence of damping is good news (it reduces sensitivity to small scale physics)
- origin is unclear (associated to shell-crossings ?)


## Charting PT

| number of loops in standard PT for Gaussian |  |  |  |  |  |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| Initial Conditions |  |  |  |  |  |  |

## Basics of theory of large deviation functions

Review paper by Hugo Touchette, '09
Beyond the central limit theorem
One exemple : tossing, coins and counting the number of heads $\quad x=\frac{1}{n} \sum_{n} t_{n}$



Central limit theorem : $\quad I(x)=2(x-0.5)^{2}$
Exact result : $\quad I(x)=x \log [x]+(1-x) \log [1-x]+\log [2]$

The cumulant generating function: $\varphi(\lambda)=\log \left(e^{\lambda} / 2+1 / 2\right)$ Cramér's Theorem : both are Legendre transform of one-another


## Key theorems: relation between rate function and cumulant generating function

Consider a random variable $x$ such that, $\quad x=\frac{1}{n} \sum_{n} v_{i} \quad\left(n \equiv \frac{1}{\sigma^{2}}\right)$
The (scaled) cumulant generating function of $x$ is defined as,

$$
\varphi(\lambda)=\lim _{n \rightarrow \infty} \frac{1}{n} \log \left[\left\langle e^{n \lambda x}\right\rangle\right]=\log \left[\left\langle e^{\lambda v}\right\rangle\right] \text { (in case vi are IID) }
$$

The Gärtner-Ellis Theorem (Cramér's Theorem for IID): the rate function is the Legendre-Fenchel transform of the (scaled) cumulant generating function

$$
I(\rho)=\sup _{\lambda}[\lambda \rho-\varphi(\lambda)]
$$

Under some regularity conditions, this relation can be inverted in

$$
\varphi(\lambda)=\sup _{\rho}[\lambda \rho-I(\rho)]
$$

## The Contraction Principle

$$
\text { For a mapping } x \rightarrow y \text { we have }, \quad I(y)=\inf _{x, x \rightarrow y} I(x)
$$

that is the rate function for $y$ is the smallest rate function (the most probable) of the values (configurations) that lead to $y$.

## Applications

- Shannon entropy (as rate function) and free energy (as cumulant generating function) in statistical mechanics ;
- Natural generalization for non-equilibrium systems (rate function for configurations) ;
- escape time in dynamical systems in presence of noise ;
- Queuing systems;
- etc..

Consequences in the context of LSS cosmology are at least 2 folds

- you do not need to impose $\delta(x)$ to be small everywhere, only the variance has to be small;
- you have a possible working procedure provided you can identify the leading initial configuration and its probability (rate function).

In practice, such an identification can be done only for configurations with enough symmetries

## Density PDFs in concentric cells

description of full joint PDF de densities in concentric cells $\quad P\left(\rho\left(R_{1}\right), \rho\left(R_{2}\right)\right) \mathrm{d} \rho\left(R_{1}\right) \mathrm{d} \rho\left(R_{2}\right)$


For spherical symmetry there exists a function $\zeta$ that gives the density $\rho$ as a function of the linear density contrast $T$

The rate functions (from the contraction principle)

$$
\begin{array}{r}
\Psi\left(\left\{\rho_{i}\right\}\right)=\frac{1}{2} \sum_{i j} \Xi_{i j} \tau_{i} \tau_{j} \\
\text { with } \\
\sigma^{2}\left(R_{i} \rho_{i}^{1 / 3}, R_{j} \rho_{j}^{1 / 3}\right) \Xi_{j k}=\delta_{i k}
\end{array}
$$

The matrix $\equiv$ is given by the inverse of correlation matrix of the density between cells at Lagrangian radius.
initially implemented in FB' 94, FB \& Valageas '00 and developed in Valageas '02

The final expression of the scaled cumulant generating function is then given by

$$
\varphi\left(\left\{\lambda_{i}\right\}\right)=\sum_{i} \lambda_{i} \rho_{i}-\Psi\left(\left\{\rho_{i}\right\}\right) \quad \begin{aligned}
& \text { with stationary } \\
& \text { conditions }
\end{aligned} \quad \lambda_{i}=\frac{\partial \Psi\left(\left\{\rho_{i}\right\}\right)}{\partial \rho_{i}}
$$

## Connexion with diagrams in standard PT

scaled cumulant GF: $\quad \varphi(\lambda)=\lim _{\left\langle\rho^{2}\right\rangle_{c} \rightarrow 0}\left\langle\rho^{2}\right\rangle_{c} \sum_{p=1}^{\infty} \frac{\left\langle\rho^{p}\right\rangle_{c}}{p!}\left(\frac{\lambda}{\left\langle\rho^{2}\right\rangle_{c}}\right)^{p}=\lambda+\frac{\lambda^{2}}{2}+S_{3} \frac{\lambda^{3}}{3!}+\ldots$
Average of (combination of) tree order expression of the $p$-point correlation functions in spherical cells.

$$
\text { Expression of } S_{p}=\lim _{\left\langle\delta^{2}\right\rangle_{c} \rightarrow 0} \frac{\left\langle\delta^{p}\right\rangle_{c}}{\left\langle\delta^{2}\right\rangle_{c}^{p-1}}=\text { tree order expr. }
$$



$$
\begin{aligned}
\left\langle\delta^{3}\right\rangle= & 6 \int \frac{\mathrm{~d} \mathbf{k}_{1}}{(2 \pi)^{3}} P\left(k_{1}\right) P\left(k_{2}\right) \\
& \times F_{2}\left(\mathbf{k}_{1}, \mathbf{k}_{2}\right) W\left(k_{1} R\right) W\left(k_{2} R\right) W\left(\left|\mathbf{k}_{1}+\mathbf{k}_{2}\right| R\right)
\end{aligned}
$$

$$
\begin{aligned}
& S_{3}=\frac{34}{7}+\gamma_{1} \\
& S_{4}=\frac{60712}{1323}+\frac{62 \gamma_{1}}{3}+\frac{7 \gamma_{1}^{2}}{3}+\frac{2 \gamma_{2}}{3}
\end{aligned}
$$

$$
S_{5}=\frac{200575880}{305613}+\frac{1847200 \gamma_{1}}{3969}+\frac{6940 \gamma_{1}^{2}}{63}+\frac{235 \gamma_{1}^{3}}{27}
$$

$$
+\frac{1490 \gamma_{2}}{63}+\frac{50 \gamma_{1} \gamma_{2}}{9}+\frac{10 \gamma_{3}}{27}
$$

$$
\gamma_{p}=\frac{\mathrm{d}^{p} \log \sigma^{2}\left(R_{0}\right)}{\mathrm{d} \log ^{p} R_{0}} .
$$

1-cell density
cumulants (FB '94)
it has a non trivial dependence on the wave vectors through the functions F3 and F2

## Identification of initial configuration for general profiles

Considering the statistical properties of

$$
\begin{aligned}
& \rho_{w}=\sum_{i} w_{i} \rho\left(<R_{i}\right) \\
& \varphi(\lambda)=\sup _{\left\{\tau_{i}\right\}}\left[\lambda \sum_{i} w_{i} \zeta\left(\tau_{i}\right)-\Psi\left(\left\{\tau_{i}\right\}\right)\right]
\end{aligned}
$$

its scaled cumulant generating function is
(looking for most likely configuration with Lag. mult)

## Consequences

$$
S_{3}=3 \nu_{2} \frac{\int d x w(x) \Pi^{2}(x)}{\left[\int d x w(x) \Pi(x)\right]^{2}}+\frac{\int d x w(x) x \frac{d}{d x} \Pi^{2}(x)}{\left[\int d x w(x) \Pi(x)\right]^{2}} \quad \begin{array}{ll}
\text { with } \\
\Pi(x)=\int d y \Xi(x, y) w(y)
\end{array}
$$



Gaussian filter, points by Juszkiewicz, Bouchet, Colombi '93 obtained from direct calculation for specific power law spectra.

Top-hat filter, FB '94

## The 2 cell cumulant generating function

The global shape of the joint cumulant generating function

critical lines $=$ stationary constraint is singular $/$ signal to noise $>10 \%$

## From cumulant to PDFs





Figure 3. Density profiles in underdense (solid light blue), overdense (dashed purple) and all regions (dashed blue) for cells of radii $R_{1}=10 \mathrm{Mpc} / \mathrm{h}$ and $R_{2}=11 \mathrm{Mpc} / \mathrm{h}$ at redshift $z=0.97$. Predictions are successfully compared to measurements in simulations (points with error bars).

## Towards a complete theory of count-in-cell statistics...



Figure 1. the configuration of multiple concentric count in cell statistics.

$$
\begin{aligned}
& \mathcal{P}\left(\left\{\hat{\rho}_{k}\right\},\left\{\hat{\rho}_{k}^{\prime}\right\} ; r_{e}\right)= \\
& \quad \mathcal{P}\left(\left\{\hat{\rho}_{k}\right\}\right) \mathcal{P}\left(\left\{\hat{\rho}_{k}^{\prime}\right\}\right)\left[1+\xi\left(r_{e}\right) b\left(\left\{\hat{\rho}_{k}\right\}\right) b\left(\left\{\hat{\rho}_{k}^{\prime}\right\}\right)\right]
\end{aligned}
$$

A regime of large-deviation functions can be identified in LSS cosmology.

- Observables can be related to joint PDFs of the density in concentric cells but also to the cumulant generating function.
- Natural application of these approaches is the density and profile PDFs


## Perspectives:

- These calculations can be applied to 3D and projected mass maps, and to join density of multiple tracers;
- biasing of over-dense/under-dense regions can also be computed = statistical properties of clipped regions;


## Charting PT

number of loops in standard PT for Gaussian Initial
Order of observable in field expansion Conditions

|  | Tree order <br> LO | I-loop <br> NLO | 2-loops <br> NNLO | 3-loops | $\ldots$..p-loops |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 2-point <br> statistics | OK | OK | OK | partial exact <br> results | partial resum | | where most of |
| :---: |
| the action is |

## Kernels in Perturbation Theory calculations



Expression of the density kernel for the propagator at I-loop order

$$
f(k, q)=\frac{3\left(2 k^{2}+7 q^{2}\right)\left(k^{2}-q^{2}\right)^{3} \log \left[\frac{(k-q)^{2}}{(k+q)^{2}}\right]+4\left(6 k^{7} q-79 k^{5} q^{3}+50 k^{3} q^{5}-21 k q^{7}\right)}{2016 k^{3} q^{5}}
$$

## The spherical collapse: the solution for specific initial conditions

The radius evolution

The exact non-linear mapping for spherically symmetric initial field (for growing mode setting)

$$
\frac{\mathrm{d}^{2} R}{\mathrm{~d} t^{2}}=-\frac{G M(<R)}{R^{2}}
$$



For spherical symmetry perturbations there exists a function $\zeta$ that gives the density at time $\eta$ knowing the density $\rho_{0}$ within the same Lagrangian radius at time $\mathrm{n}_{0}$.

$$
\zeta_{\rho}\left(\eta ; \rho_{0}, \eta_{0}\right)
$$

## The result

The rate functions, Legendre Transform of the cumulant generating function,

$$
\varphi\left(\left\{\lambda_{k}, R_{k}\right\}, \eta\right)=\sum_{p_{i}=0}^{\infty}\left\langle\Pi_{i} \hat{\rho}_{R_{i}}^{p_{i}}\right\rangle c \frac{\Pi_{i} \lambda_{i}^{p_{i}}}{\Pi_{i} p_{i}!} \quad \psi\left(\left\{\rho_{k}, R_{k}\right\}, \eta\right)=\sum_{i} \lambda_{i} \rho_{i}-\varphi\left(\left\{\lambda_{k}, R_{k}\right\}, \eta\right)
$$

have, according to the contraction principle, the following time dependence,

$$
\Psi\left(\left\{\rho_{k}, R_{k}\right\}, \eta\right)=\Psi\left(\left\{\zeta\left(\rho_{k}, \eta, \eta^{\prime}\right), R_{k} \frac{\zeta^{1 / 3}\left(\rho_{k}, \eta, \eta^{\prime}\right)}{\rho_{k}^{1 / 3}}\right\}, \eta^{\prime}\right)
$$

In other words we know how to compute the cumulant generating function of densities in concentric cells starting with specific initial conditions.

## The mathematical part, construction of the cumulant generating function

from ideas in FB' 94 see also FB \& Valageas '00 and fully developed in Valageas '02
Can we get the whole generating function of the cumulants?

It is given by the following relation (multidimensional Laplace transform of joint-PDFs)

$$
\begin{array}{r}
\varphi\left(\left\{\lambda_{k}\right\}\right)=\sum_{p_{i}=0}^{\infty}\left\langle\Pi_{i} \rho_{i}^{p_{i}}\right\rangle_{c} \frac{\Pi_{i} \lambda_{i}^{p_{i}}}{\Pi_{i} p_{i}!} \\
\exp \left[\varphi\left(\left\{\lambda_{i}\right\}\right)\right]=\left\langle\exp \left(\sum_{i} \lambda_{i} \rho_{i}\right)\right\rangle \\
=\int_{0}^{\infty} \Pi_{i} \mathrm{~d} \rho_{i} P\left(\left\{\rho_{i}\right\}\right) \exp \left(\sum_{i} \lambda_{i} \rho_{i}\right)
\end{array}
$$

Formal solution

$$
\begin{gathered}
=\int_{0}^{\infty} \Pi_{i} \mathrm{~d} \rho_{i} P\left(\left\{\rho_{i}\right\}\right) \exp \left(\sum_{i} \lambda_{i} \rho_{i}\right) \\
\exp \left[\varphi\left(\left\{\lambda_{i}\right\}\right)\right]=\int \mathcal{D}[\tau(\vec{x})] \mathcal{P}[\tau(\vec{x})] \exp \left(\lambda_{i} \rho_{i}[\tau(\vec{x})]\right)
\end{gathered}
$$

Principle of the calculations : in the small variance approximation one can look for the most probable configuration - for fixed $\rho_{i}$ - and compute the resulting cumulant generating function using the

The (conjectured) solution for spherical cells: an initial steepest-descent method. spherical perturbation the profile of which can be computed from spherical collapse solution.

$$
\rho_{i}=\zeta_{\mathrm{SC}}\left(\tau_{i}\right)
$$

The 1-cell rate function and cumulant generating function


The 1-cell rate function compared to Gaussian approximation.
critical point


## Example of contribution to the 3 - to $n$-point

 cumulants at tree order

$$
\begin{aligned}
\left\langle\delta^{3}\right\rangle= & 6 \int \frac{\mathrm{~d} \mathbf{k}_{1}}{(2 \pi)^{3}} P\left(k_{1}\right) P\left(k_{2}\right) \\
& \times F_{2}\left(\mathbf{k}_{1}, \mathbf{k}_{2}\right) W\left(k_{1} R\right) W\left(k_{2} R\right) W\left(\left|\mathbf{k}_{1}+\mathbf{k}_{2}\right| R\right) \\
\propto & \left\langle\delta^{2}\right\rangle^{2}
\end{aligned}
$$

it has a non trivial dependence on the wave vectors through the functions F3 and F2

$$
\left\langle\delta^{p}\right\rangle_{c} \propto\left\langle\delta^{2}\right\rangle_{c}^{p-1}
$$

## Predictions for cumulants and PDFs...

Baugh \& Gaztañaga '95
Prediction at tree order is very accurate

Let us assume that,

$$
\varphi(\lambda ; \sigma)=\frac{1}{\sigma^{2}} \varphi_{c}\left(\sigma^{2} \lambda\right)
$$



## Application 1: 1-cell PDF and stats

The inverse Laplace transform,

$$
\mathcal{P}\left(\hat{\rho}_{1}\right)=\int_{-\mathrm{i} \infty}^{+\mathrm{i} \infty} \frac{\mathrm{~d} \lambda_{1}}{2 \pi \mathrm{i}} \exp \left(-\lambda_{1} \hat{\rho}_{1}+\varphi\left(\lambda_{1}\right)\right)
$$

requires integration into complex plane.
$P(\rho)=\frac{1}{\sqrt{2 \pi}} \sqrt{\frac{\partial^{2} \Psi(\rho)}{\partial \rho^{2}}} \exp [-\Psi(\rho)]$
low density approximation
$P(\rho)=\frac{3 a \frac{3}{2}}{4 \sqrt{\pi}} \exp \left(\varphi^{(c)}-\lambda^{(c)} \rho\right) \frac{1}{\left(\rho+r_{1}+r_{2} / \rho+\ldots\right)^{5 / 2}}$,
large density approximation


## Comparison with simulations: the 1-point PDF shape (500 h^(-I) Mpc) ${ }^{\wedge} 3$




FB, Pichon, Codis 'I3

## Residuals as a function of $R$



Effect of scale dependent index


