



Nuclear Absorption Lines and Occulted Sources



Outline

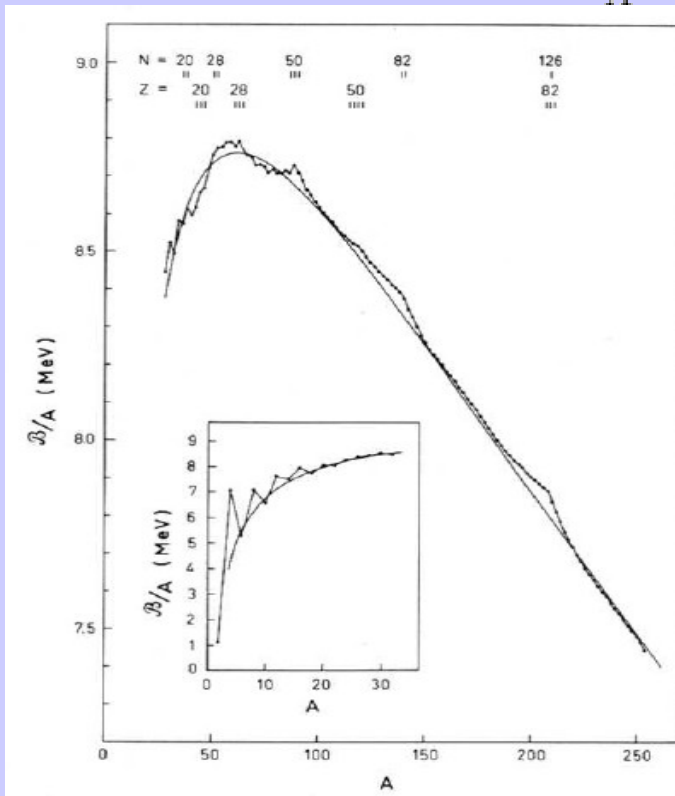
- 1. Nuclear Structure / Models**
- 2. Nuclear - Photon Interaction**
 - 2.1 Dipolinteraction - GDR**
 - 2.2 Pygmy Dipole Resonance - PDR**
 - 2.3 Particle Production**
 - 2.4 Total Cross Sections**
- 3. Photo absorption in γ -ray spectra**
 - 3.1 Obscured Sources**
 - 3.2 Absorption Lines**
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Nuclear Structure

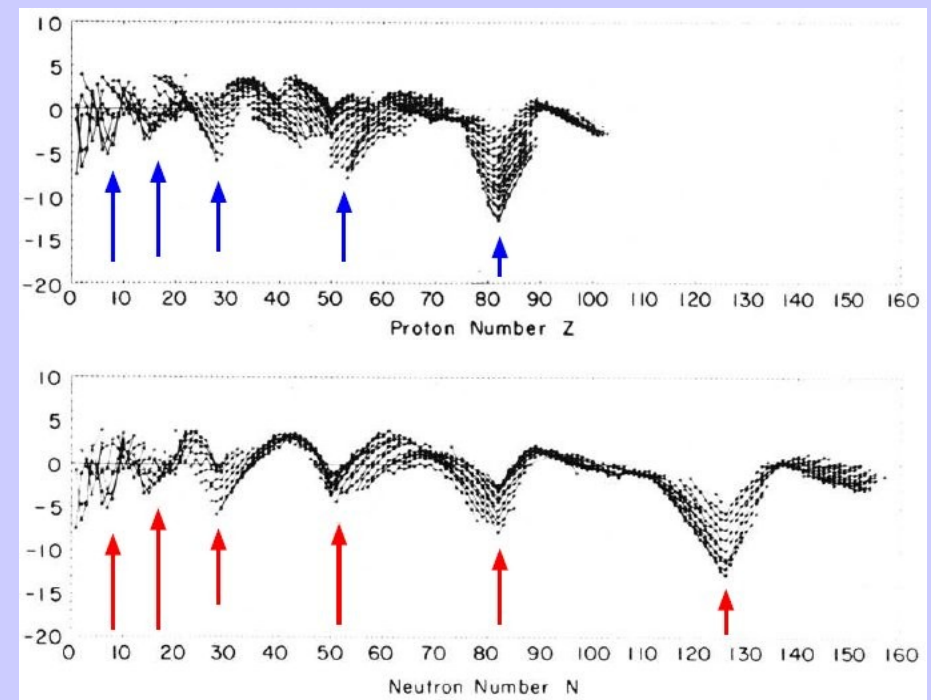
Binding Energies described in the Nuclear Drop Model:

-> Weizsäcker mass formula:

$$E_b = a_1 A - a_2 A^{2/3} - a_3 Z^2 A^{-1/3} + a_4 \frac{(Z - \frac{A}{Z})^2}{A} \pm f_5.$$



But: Deviations of the Binding Energy Depending on the Number of Protons/Neutrons



Nuclear Structure / Shell Model

Shell Modell: Describes the Nukleon in an effective potential of the others

- Spin – Orbit - Coupling

-> Magic Numbers

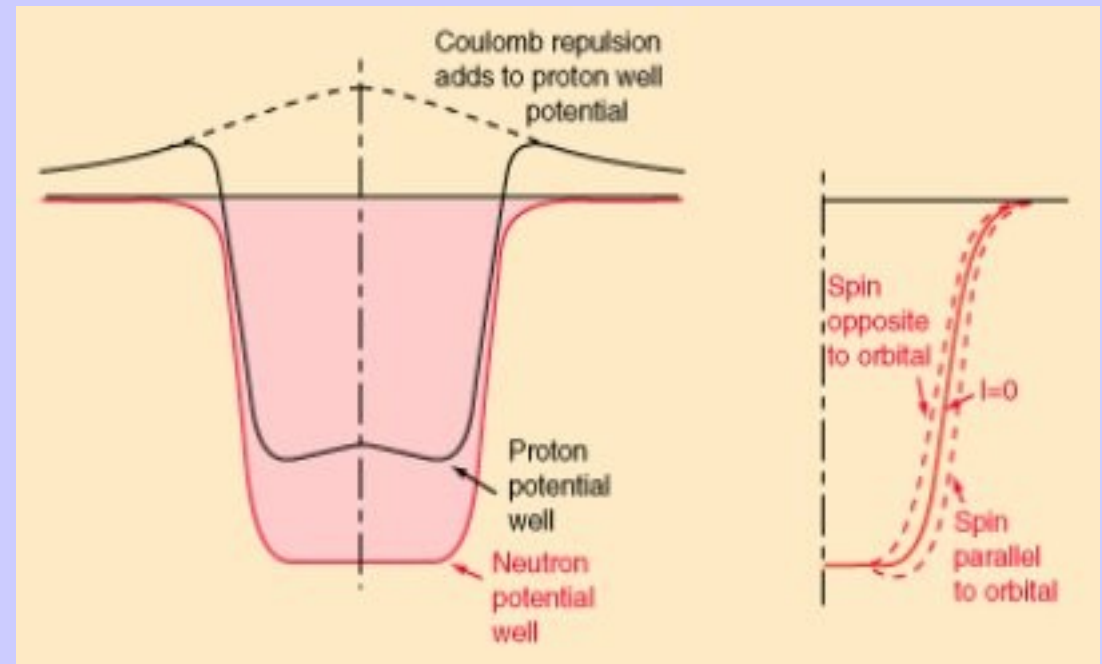
$$V_{S.O.}(\mathbf{r}; \beta_2) = -\lambda \left(\frac{\hbar}{2mc} \right)^2 [\nabla V(\mathbf{r}; \beta_2)_{S.O.} \times \mathbf{p}] \cdot \mathbf{s}$$

Woods Saxon Potential:

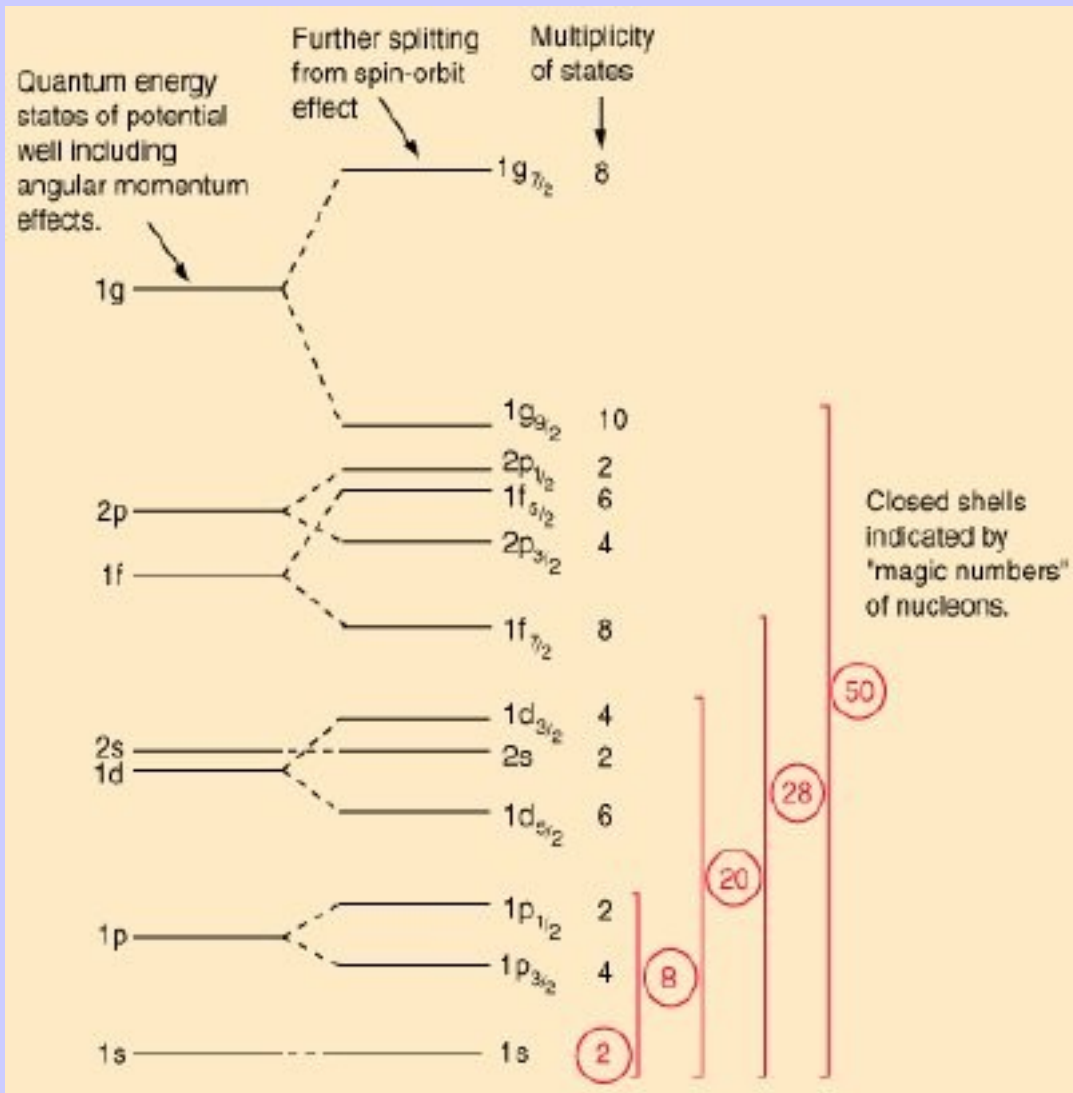
depth of the well is proportional to the density of nucleons

spherical Woods-Saxon potential:

$$V = -\frac{V_0}{1 + \exp\left[\frac{r-R}{a}\right]}$$



Nuclear Structure / Shell Model



But: Shell Model does not explain Quadrupol Moment for a number of nuclei, e.g. ^{17}O

-> Collective vibrations and rotations

-> Collective Motion of all nucleons



Nuclear Structure / Collective Model

Collective model combines liquid drop and shell model:

- Net nuclear potential due to filled core states
- Nucleons in the unfilled shells move independently within the core potential
- Outer-shell nucleons disturb the potential

$$V(\mathbf{r}; \beta_2) = -\frac{V_0}{1 + \exp[r - R(\mathbf{r}; \beta_2)/a]}$$

- Potential deformation parameter β
- Two types of low lying (0..2 MeV), collective effects:
 - Excitation of surface vibrations
 - Rotations of the deformed shape



Electronic Excitations

Fermis Golden Rule yields transition propability:

$$dW = \frac{2\pi}{\hbar} |\langle \psi_f | H_{\text{int}} | \psi_i \rangle|^2 \cdot d\rho(E)$$

Hamiltonian for electromagnetic interactions:

$$H_{\text{int}} = -\frac{e}{m} \vec{p} \vec{A} + e\phi \equiv \mathbf{j} \cdot \mathbf{A}$$

Incoming Photon:

$$dW_{\text{fi}} = \frac{e^2}{8\pi^2 \epsilon_0 \hbar^4 c^3} E_\gamma^3 |\vec{\epsilon} \int d^3r \psi_f^* \vec{r} e^{i\vec{k}\vec{r}} \psi_i|^2 d\Omega$$

Multipolexpansion:

$$e^{i\vec{k}\vec{r}} = 1 + i\vec{k}\vec{r} + \dots$$

Dipol transmission propability (Photon spin parity : $J^\pi = 1^-$)

$$W_{\text{fi}}^{(\text{E1})} = \frac{1}{\tau} = \frac{e^2}{3\pi\epsilon_0 \hbar^4 c^3} E_\gamma^3 \left| \int d^3r \psi_f^* \vec{r} \psi_i \right|^2$$

Photo-absorption

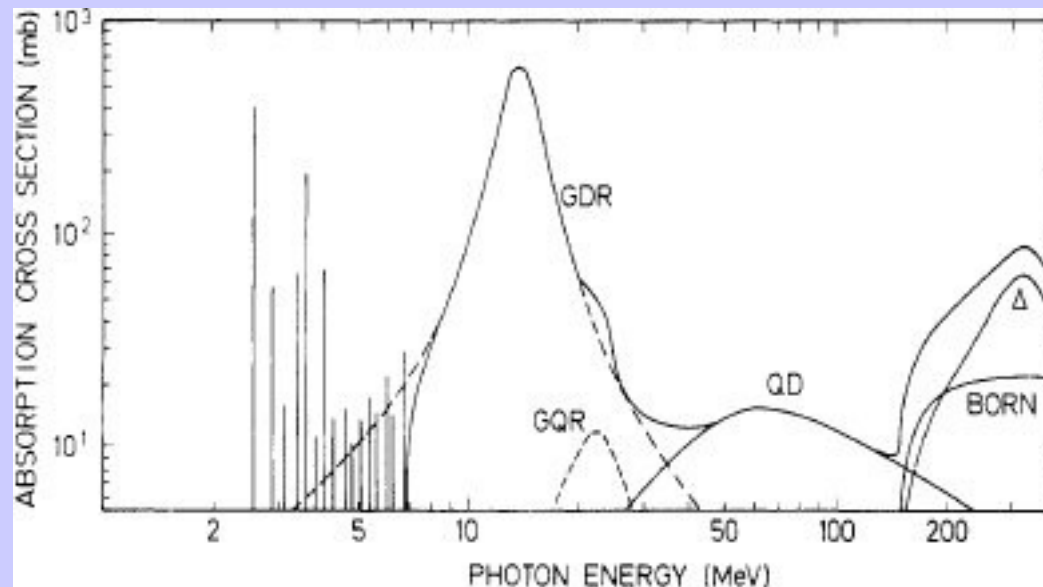
Pygmy Dipol Resonance PDR

Giant Resonances:

- Giant Dipole Resonance GDR
- Giant Quadrupol Resonance GQR
- Quasi-Deuteron Excitations QD

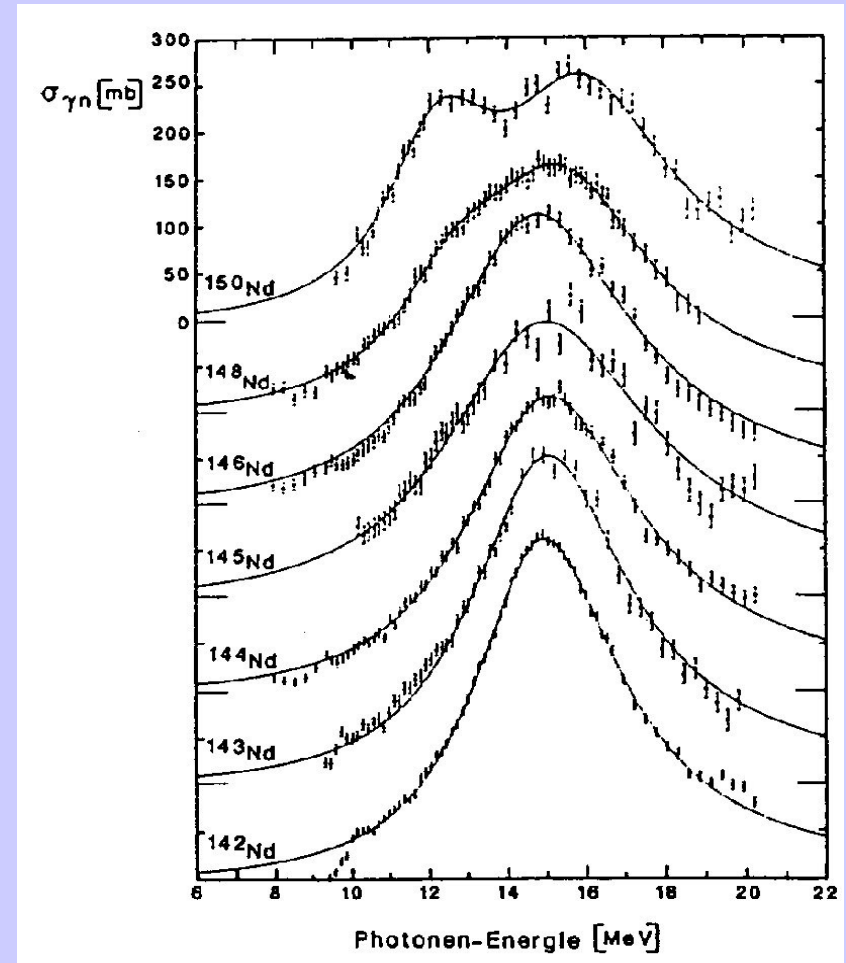
Particle Excitations:

- Delta Resonance
- Born Meson Production



Giant Resonances

- Oszillation of Neutrons against Protons
- Above the particle production energy (at 8..10 MeV)
- Spherical Nuclei: Vibrations along all three perpendicular axes are equivalent and the strength is distributed equally.
- Deformed Nuclei: two peaks: Vibrations of protons against neutrons along the symmetry axis vs. Vibrations perpendicular to the symmetry axis.
- Exact description complicated, collective excitation of a many-body system
- GDR (Giant Dipole Resonance dominates total cross section for E1 transitions)





Giant Resonances

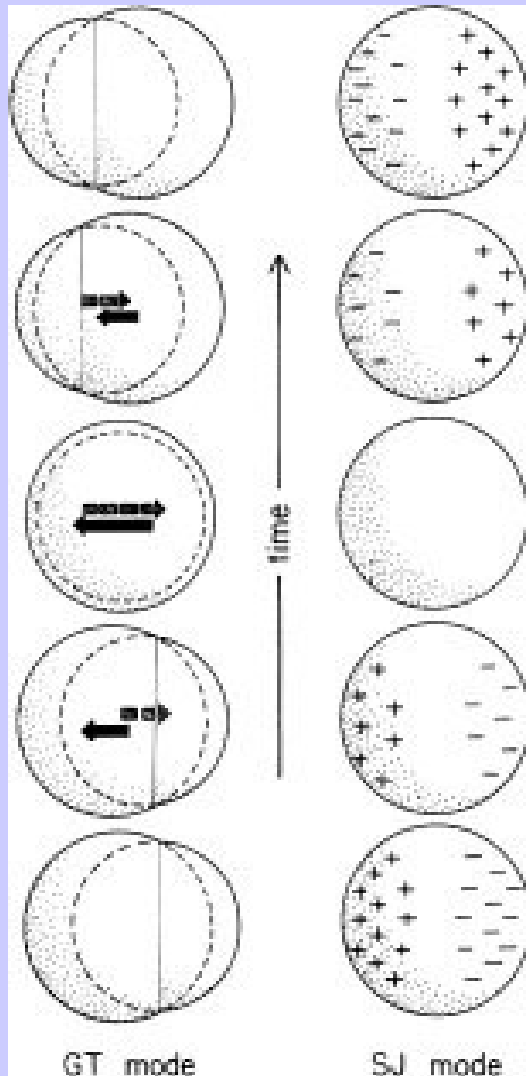
- Generally described in RPA (Random Phase Approximation):
 - > harmonic, mean-field description of collective excitations of a many-body system
 - > collective motion is contained in the one-body density matrix $\rho(r, r', t)$
 - > Solution of the time dependent Hartree Fock Equation gives the equation of motion for $\rho(r)$ and the corresponding wave functions

$$i\hbar \frac{\partial \rho(r, r', t)}{\partial t} = \frac{\delta E[\rho]}{\delta \rho(r, r', t)}$$

- > particle-vibration coupling Hamiltonian $\delta U = k \alpha f$
 - α collective coordinate of the harmonic oscillator
 - f collective states exciting Operator (e.g. for GDR = D)
- > Theoretical Models described in

K. Goeke and J. Speth – Theory of Giant Resonances
Aumann – Multiphonon Giant Resonances in Nuclei

Giant Resonances / More intuitive



- Giant Dipole Resonance divided in

GT Mode (Changing Surface)
SJ Mode (Charge Fluctuations)

- Describes fairly well position and width of the GDR

$$\frac{1}{\omega_{-}^2} = \frac{m r_0^2}{8J} A^{2/3} + \frac{3m r_0^2}{8Q} A^{1/3}$$

SJ Mode

J = nuclear symm. coeff.

GT Mode

Q = stiffness coeff.

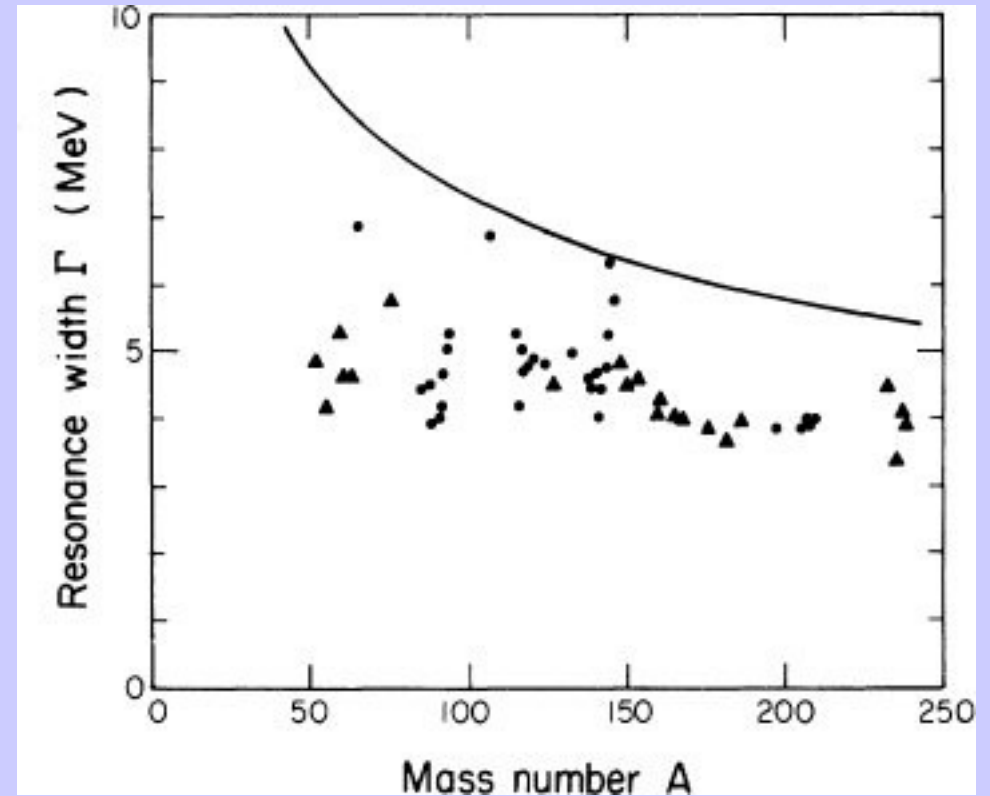
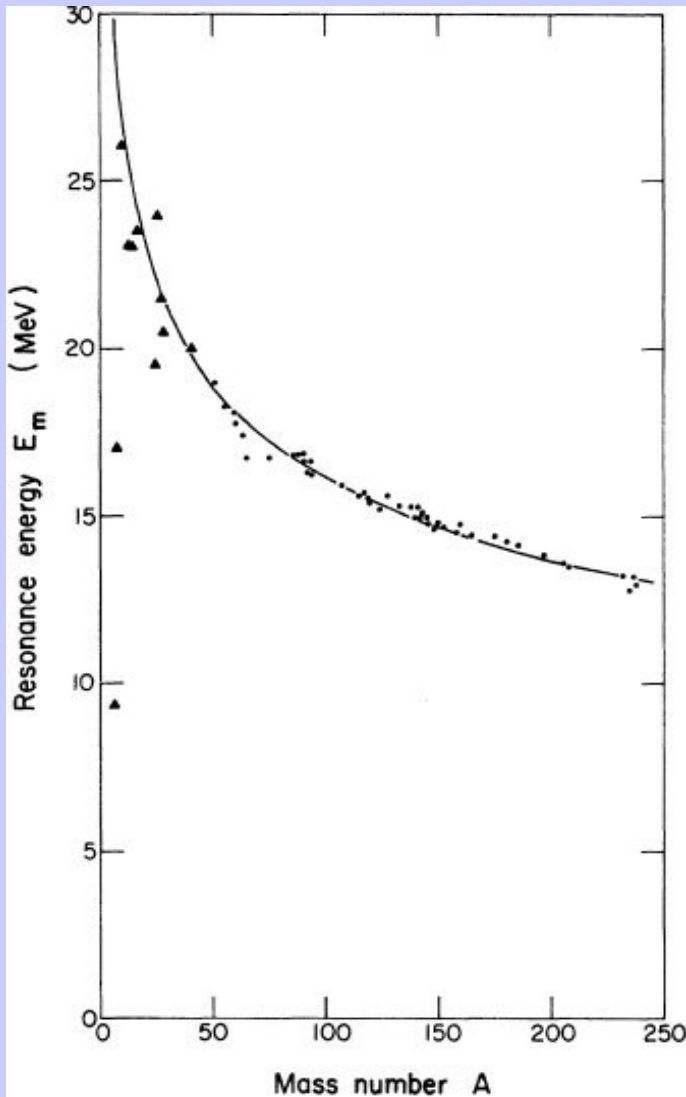
$$E_{GDR} = 31.2 A^{-1/3} + 20.6 A^{-1/6}$$

-> Mainly GT mode, with SJ mixing at high As

Width: $\Gamma = \hbar \left(\frac{\bar{v}}{\bar{R}} \right)$

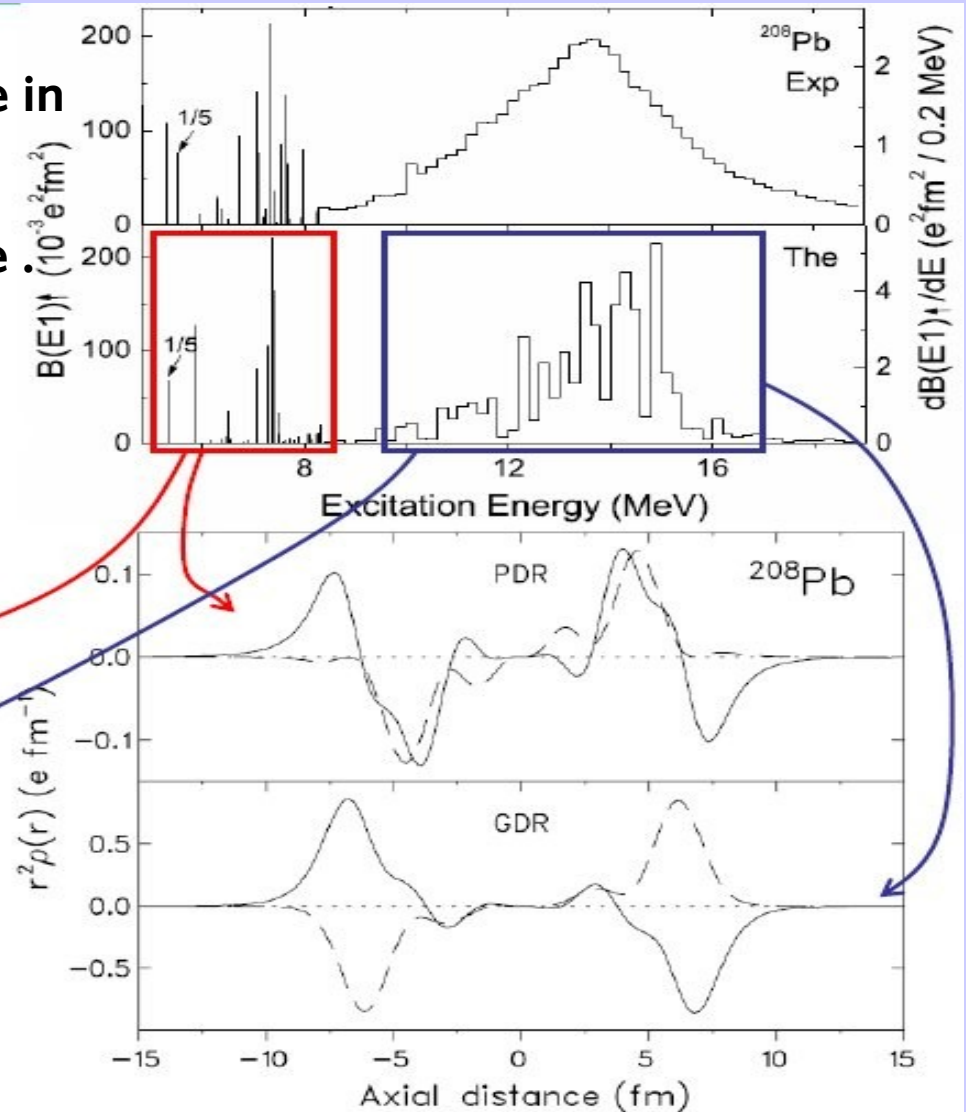
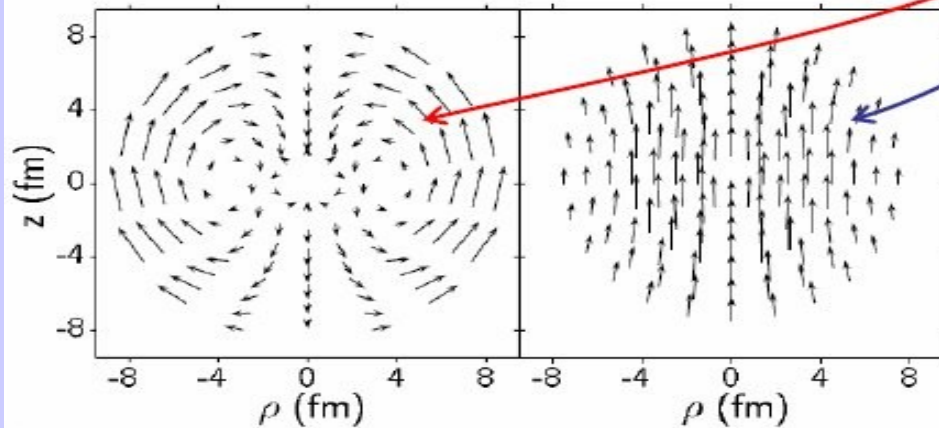


Giant Resonances / More intuitive

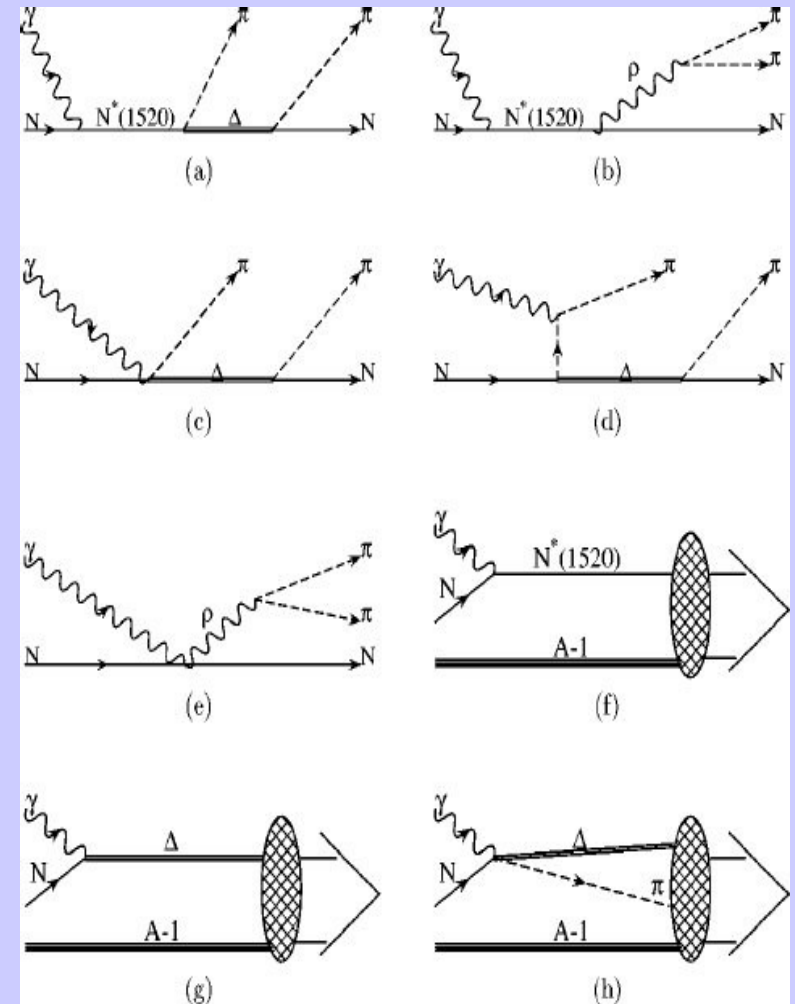


Pygmy Dipole Resonance / PDR

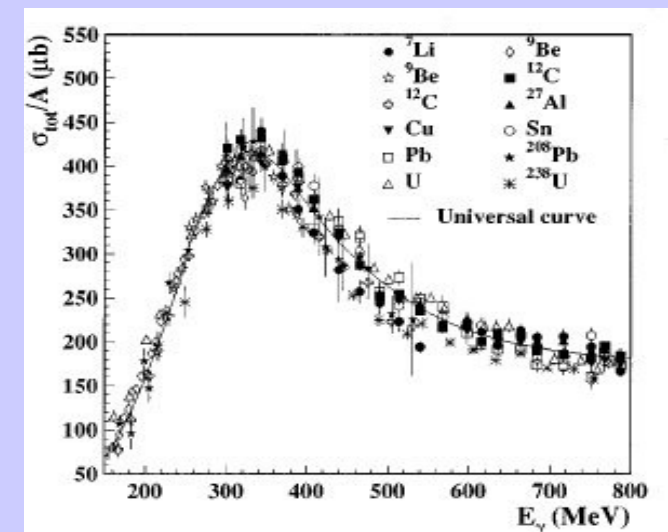
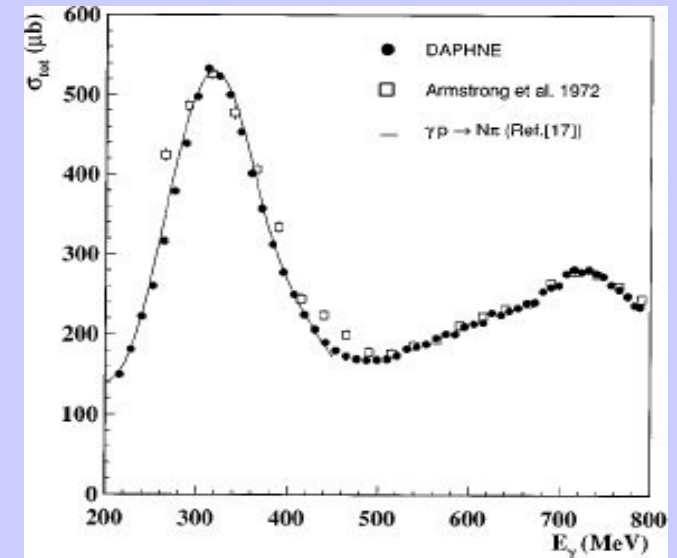
- Low-energy electric dipole resonance in stable nuclei
- Between 5 and 8 MeV (at or below the particle threshold)
- Small total cross sections compared to GDR



- Photon induced particle production
- Two prominent in between states
Delta and N^* between 300 and 800 MeV
- Produced in different reaction channels
- No Position change (327 MeV peak for Δ)
- Δ = Spinflip of one quark
- N^* = Excited state of one nuclei

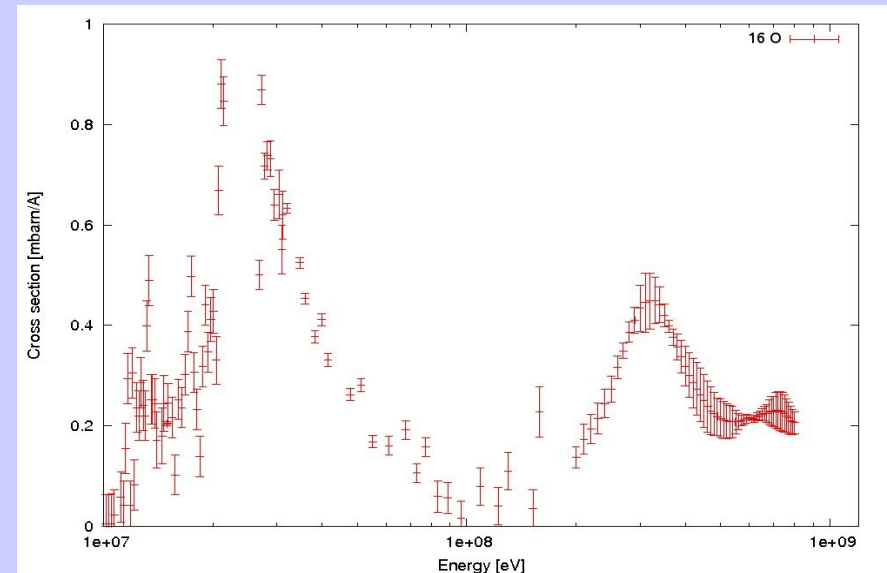
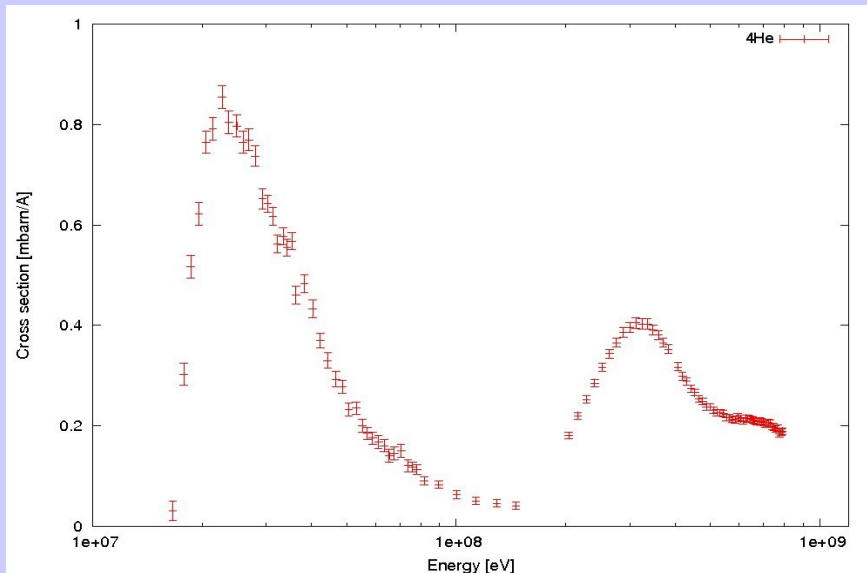
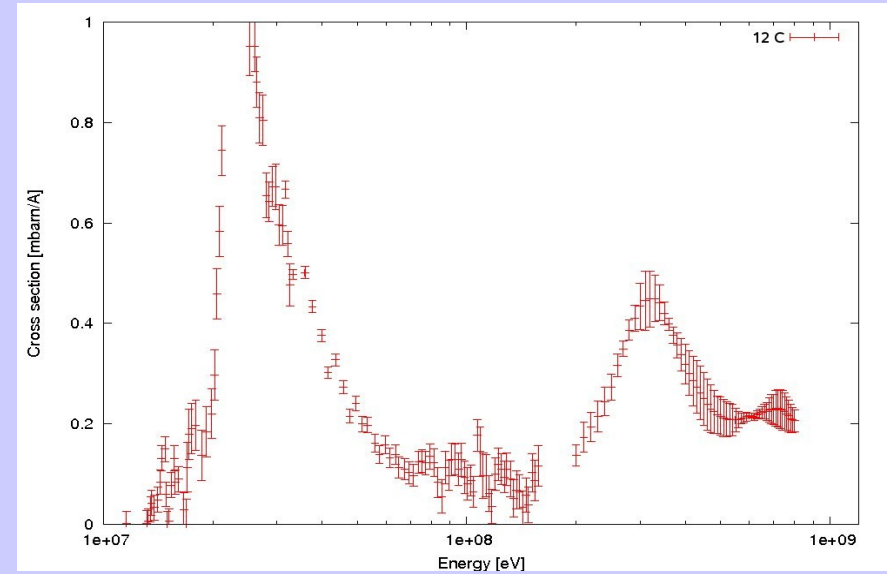
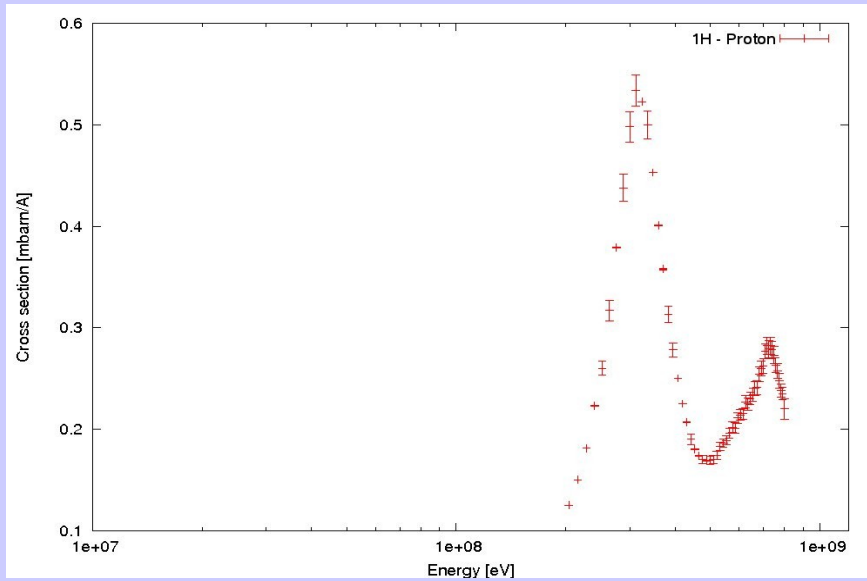


- For light elements: $A=1..4 \rightarrow N^*$
- For heavy elements: $A>7$ No N^* contribution
- Broadening of the Δ Resonance from ~ 160 MeV for a proton/neutron to ~ 300 MeV for a nuclei
 -> More open decay channels





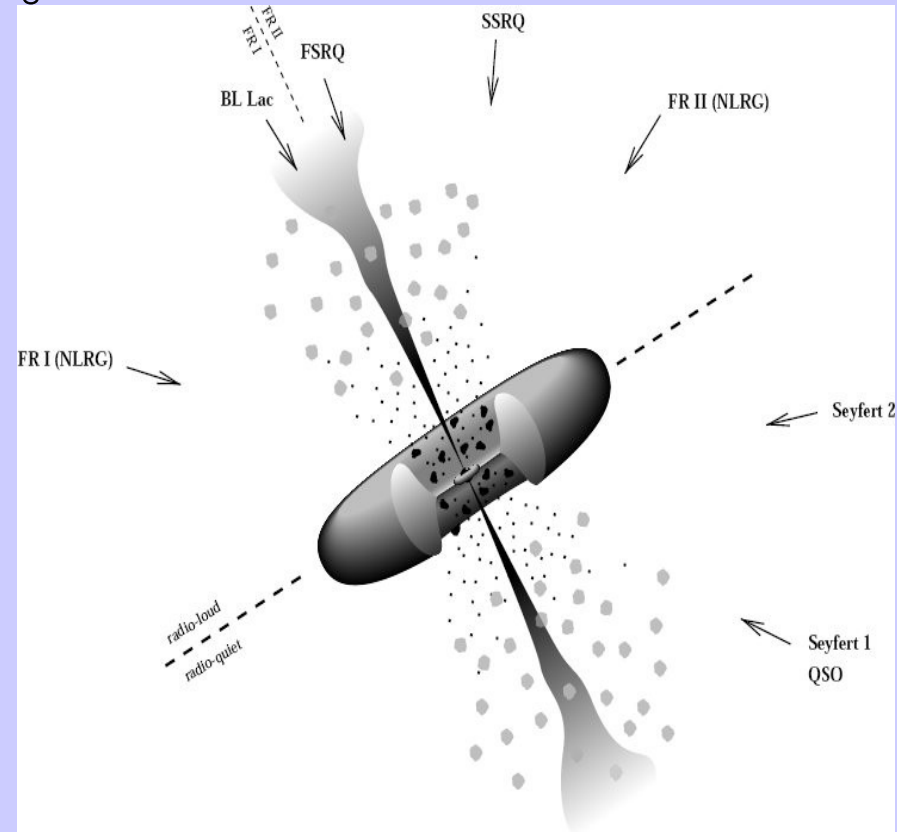
Total Cross sections



γ -ray Sources / AGNs

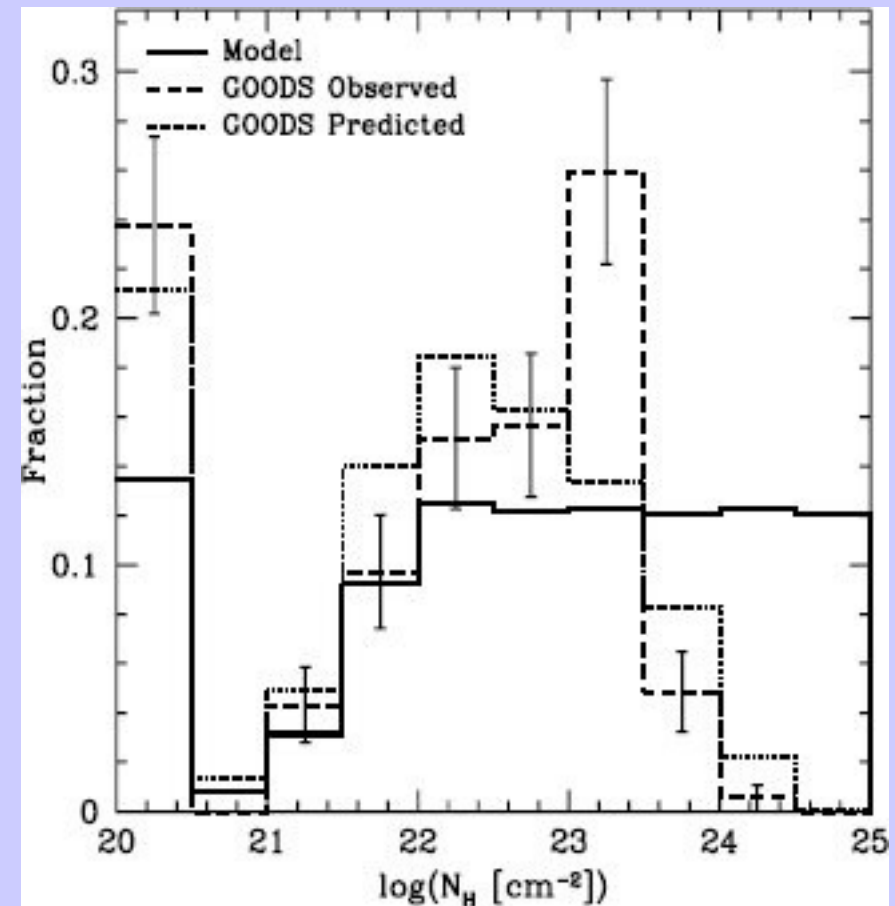
Active Galactic Nuclei

- > ~3% of all Galaxies are considered as active
- > Large luminosity released in small core regions
- > Supermassive Black hole in the center $>10^6 M_{\odot}$
- > Accretion of matter from an accretion disc feeds the black hole
- > Highly relativistic jet along the rotation axis
- > Seen along the jet axis ($<10^{\circ}$) : Blazar
- > Quasar or Seyfert 1: AGN seen at $\sim 30^{\circ}$
- > Seyfert 2: Seen at the higher angles $>60^{\circ}$
- > Radio galaxy: Seen perpendicular to the jet



γ -ray Sources / AGNs

- The viscous disc is heated up and shines brightly in the UV and soft X-ray region (Hard X- and γ -rays are produced very near the black hole)
- The central regions of many AGNs contain obscuring material (probably dust)
- AGNs account for most of the XRB.
- Geometry of the dust torus determines the amount of obscuring material along the observers line of sight the central emitting region
- Bimodal distribution between unabsorbed/absorbed sources ($N_H=10^{22} \text{ cm}^{-2}$)
- Chandra (to 10 keV) and XMM Newton (12 keV) find $\frac{3}{4}$ of all AGNs are absorbed
- Deep X-ray surveys miss a fraction of sources due to large obscuration



γ -ray Sources / AGNs

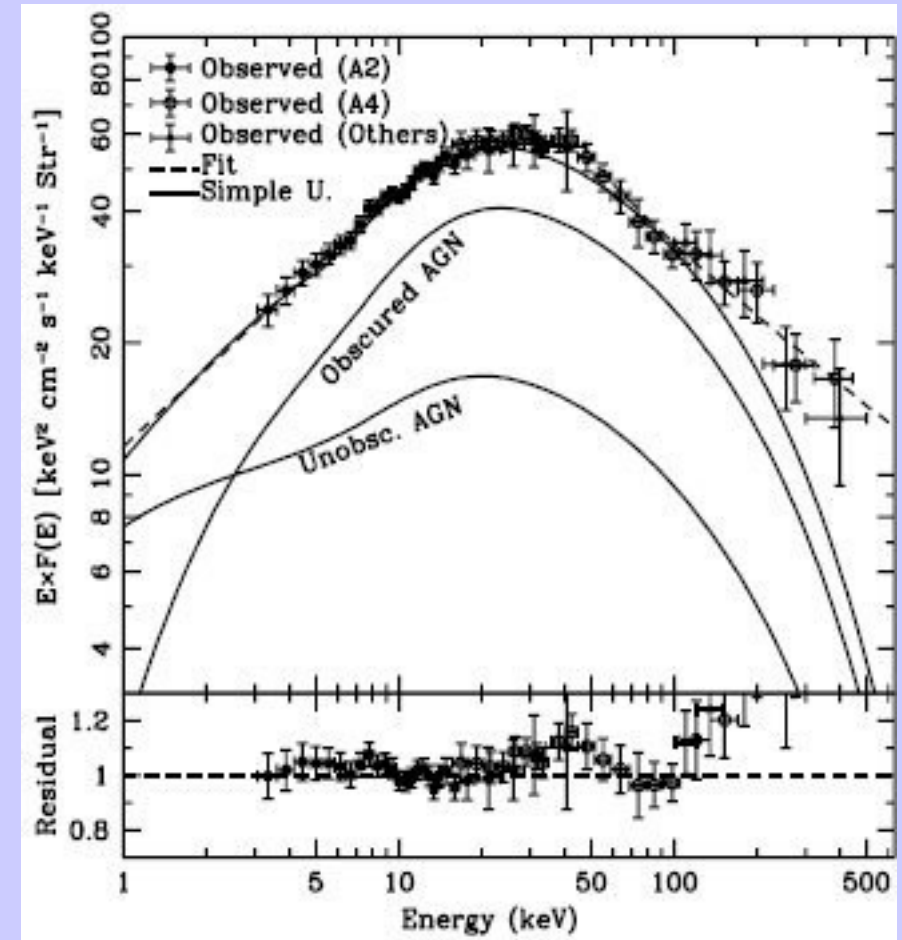
- Need for hard X-ray, γ ray studies sensitive $> N_H = 10^{25} \text{ cm}^{-2}$
- Soft X-rays surveys are sensitive to max $N_H = 10^{24} \text{ cm}^{-2}$
- Above 15 keV, Compton-thick sources can be probed

Compton-thick: $N_H > 1.5 \cdot 10^{24} \text{ cm}^{-2}$

Mildly Compton-thick : $N_H < 10^{25} \text{ cm}^{-2}$

Heavily Compton-thick : $N_H > 10^{25} \text{ cm}^{-2}$

-> Source is visible only above 10 keV





γ -ray Troughs in QSOs environments

- γ -ray studies with COMPTEL (1 MeV to 30 MeV) and EGRET (20~MeV to 30~GeV) from CGRO
- Analysis of the absorption troughs from bright γ -ray sources PKS 0528, 3C279 and some others
- Photon flux on detectable on Earth

$$\frac{dN}{dE} = \left(\frac{dN}{dE} \right)_{\text{unabsorbed}} \cdot e^{-\tau(E,z)}$$

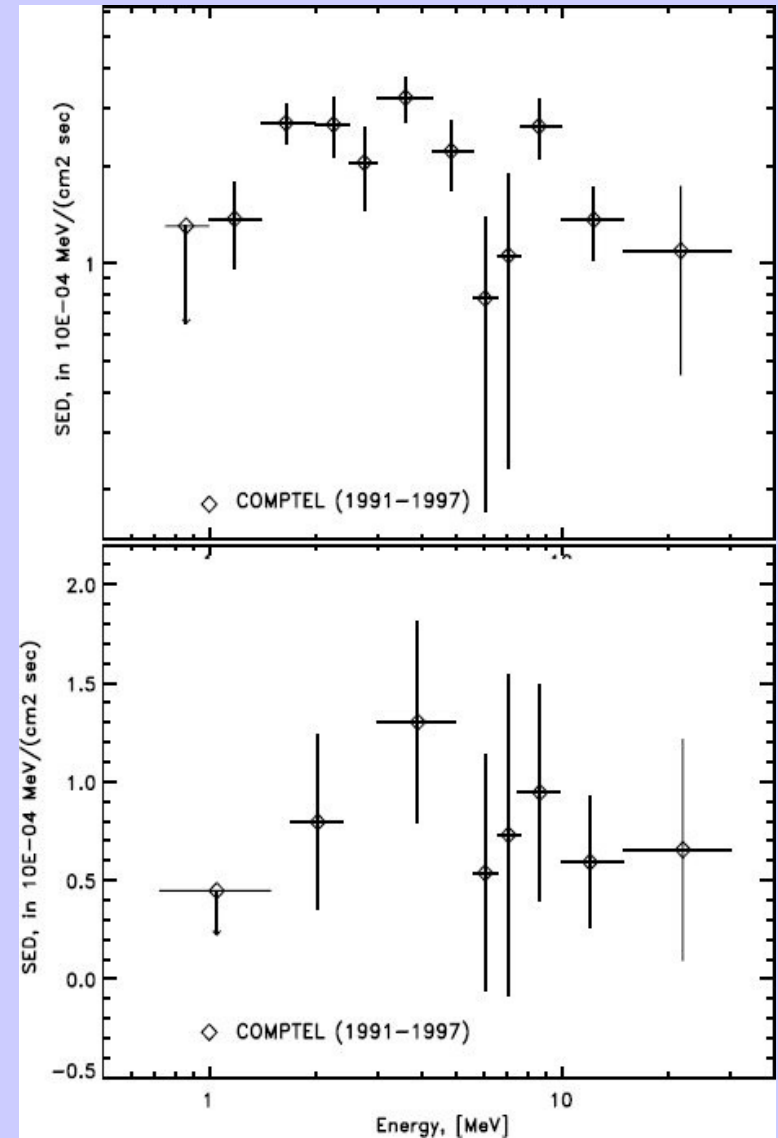
- Spectrum is fitted by two smoothly connected Power laws („Band function“)

$$N_E(E) = A \left(\frac{E}{100 \text{ keV}} \right)^\alpha \exp\left(-\frac{E}{E_0}\right),$$

for $(\alpha - \beta)E_0 \geq E$,

$$N_E(E) = A \left[\frac{(\alpha - \beta)E_0}{100 \text{ keV}} \right]^{\alpha - \beta} \exp(\beta - \alpha) \left(\frac{E}{100 \text{ keV}} \right)^\beta,$$

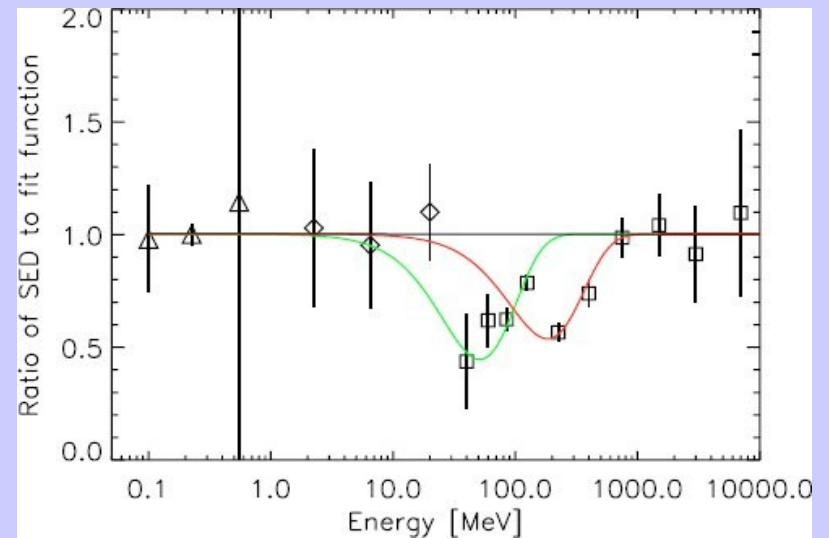
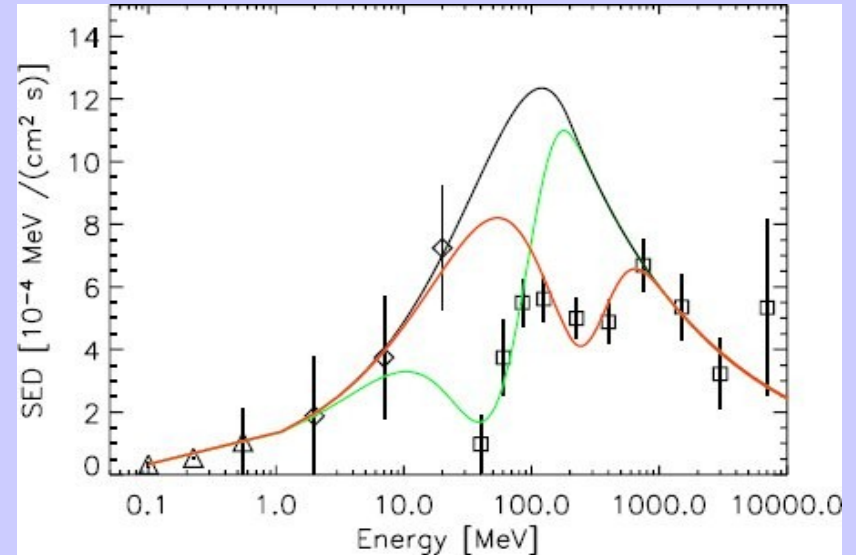
for $(\alpha - \beta)E_0 \leq E$.





γ -ray Troughs in QSOs environments

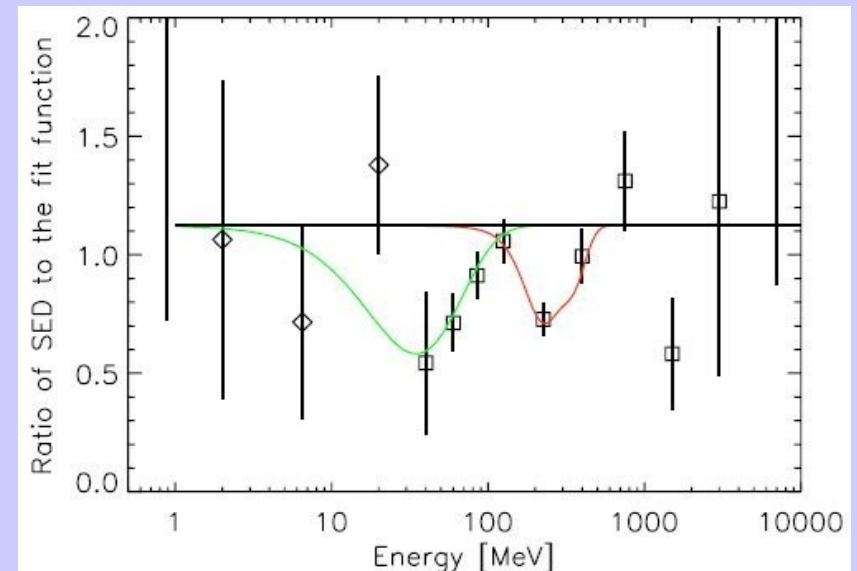
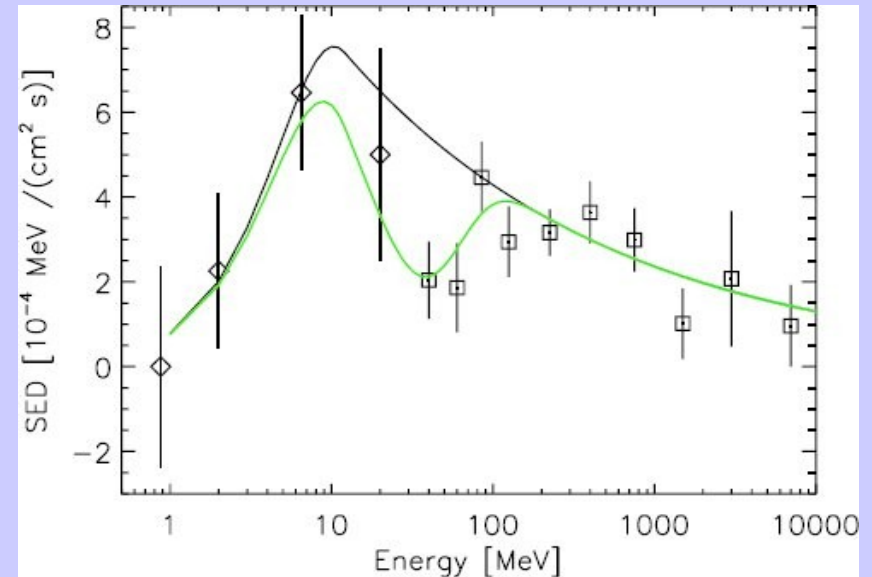
- 3C279: Optical variable Quasar:
Luminosity variation in the optical,
radio and X-ray band
Redshift $z=0.54$
- Different Observations between 91 and 96
- Identification of γ -ray absorption
from GDR and Δ -Resonance
- Flux depression of 40% for Δ -
and 93% for GDR
- Δ Resonance at 204 ± 4 MeV
 - > $E(\Delta)^{3C279} = 313 \pm 10$ MeV
 - > $E(\Delta)^{\text{real}} = 327$ MeV
- GDR at ~ 23 MeV
 - > Depending on the metallicity of
the environment





γ -ray Troughs in QSOs environments

- PKS 0528+134
Redshift $z=2.08$
- Observations from March 1993
- Again, identification of γ -ray absorption from GDR and Δ -Resonance (1bin)
- Flux depression of 34% for Δ - and 60% for GDR
- Δ Resonance at 106 ± 49 MeV
 - > $E(\Delta)^{3C279} = 326 \pm 152$ MeV
 - > $E(\Delta)^{\text{real}} = 327$ MeV
- GDR at $\sim 34 \pm 20$ MeV
 - > Depending on the metallicity of the environment





γ -ray Troughs in QSOs environments

- Absorption column densities
Assume cross sections of 0.5 mbarn for Δ and 3.6 mbarn for GDR
-> $N_{\text{H}} = 10^{27} \text{ cm}^{-2}$

- Depending on the sensitivity , source brightness and exposure time
- Smearing effect – Compton down scattering in Resonance regions
- Assume metallicity, geometry ->

$$F_{\text{Obs}}(E) = F_0(E) \left[\exp\left(-\sigma_{\text{eff}} N_{\text{H}}^{\Delta r}\right) \right]$$

$$\frac{\Delta F_j}{F_j} = \left[1 - \exp\left(-\sigma_{\text{eff}} N_{\text{H}}^{\Delta r}\right) \right]$$

- With flux depression as quoted, follows:

-> for 3C279: $N_{\text{H}} = 8 \cdot 10^{26} \text{ cm}^{-2}$

-> for PKS0528+134: $N_{\text{H}} = 7 \cdot 10^{26} \text{ cm}^{-2}$



What to remember ?

- Nuclear Models
- Nuclear Photoabsorption in the MeV region
- GDR, PDR and Δ Resonance
- AGNs
- γ ray absorption promising tool for probing high column densities, the content and metallicity of QSOs host galaxies and verify/measuring the redshift of known/unknown objects
- Upcoming mission GLAST (2007) with its instruments LAT (~ 20 MeV - ~ 300 GeV) and BAT (~ 20 keV - ~ 30 MeV) should provide evidence of γ -ray absorption troughs



Thank you for your attention