

# **The Square Kilometre Array**

Lister Staveley-Smith<sup>1,2</sup> (with thanks to Minh Huynh and Joe Lazio)



- 1. International Centre for Radio Astronomy Research (ICRAR), University of Western Australia
- 2. ARC Centre for All-sky Astrophysics (CAASTRO)



SKA Kloster Seeon 2011 June 14

### The Centre for All-sky Astrophysics



The CAASTRO Vision: To be the international leader in wide-field astronomy, positioning Australia to address fundamental unsolved questions about the Universe with the dramatic capabilities of nextgeneration telescopes and advanced instrumentation.



#### Six Australian Universities, two national facilities and international partners

- →DISCOVER: Ground-breaking advances in understanding the Universe
- → INNOVATE: New ways of processing & visualising complex data sets
- → PERFORM: High-impact discoveries using SKA pathfinders
- → UNITE: A new network of talented researchers
- → EDUCATE: Exciting opportunities for students and young scientists

# Outline



- SKA Overview
- Science drivers
- Baseline design
  - phase 1 and 2
- Status
  - schedule



# **Great Observatories for the coming decade**



ALL-SKY ASTROPHYSIC

### **SKA Key Science Drivers**

#### ORIGINS

Probing the Dark Ages and the Epoch of Reionization - EoR, first galaxies and active galactic nuclei

# Galaxy Evolution, Cosmology, and Dark Energy - When did the first stars and galaxies form?

- How did galaxies evolve?Dark Energy, Dark Matter

#### Cradle of Life

- Organic molecules in interstellar space and protoplanetary disks
- SETI

#### **FUNDAMENTAL FORCES**

Pulsars, General Relativity & Gravitational Waves

Origin & evolution of cosmic magnetism

EXPLORATION OF THE UNKNOWN Transients and new phenomena



Science with the Square Kilometre Array (2004, eds. C. Carilli & S. Rawlings, New Astron. Rev., 48)

### Galaxy Evolution, Cosmology and Dark Energy

- What is the equation of state of dark energy?
- What is the efficiency of gas accretion onto galaxies at different epochs?
- How do baryons affect the distribution of dark matter?
- How does AGN feedback affect the growth of galaxies?
- What is the relative role of thermal, magnetic and CR pressure?
- How does gas content vary with redshift and environment?

# Probing the Dark Ages and the Epoch of Reionization

Imaging the transition from a neutral to an ionised IGM

- What reioinizes the Universe at the end of the dark ages?
- When does reionization end and how quickly did it proceed?
- How quickly did supermassive black holes form?
- What is the topology of the ionised IGM; how sharp are the edges around QSO HII regions?
- When did the first radio sources form?

# **Astrobiology at Long Wavelengths**

- Imaging of thermal dust emission from proto-planetary nebulae
- Astrochemistry of complex molecules
- Prebiotic molecules.
- Proto-planetary chemistry



# **Did Einstein Have the Last Word on Gravity?**

### **Relativistic binaries:**

- Equivalence principle
- Strong-field tests of gravity
- Black hole-neutron star binaries?

### **Direct detection of Gravity Waves:**

- Pulsar timing array
- LIGO: suspended masses
- LISA: freely falling mirrors



### **Cosmic Magnetism**

- How are magnetic fields generated and maintained?
- How do magnetic fields evolve as galaxies evolve?
- What is the strength and structure of the magnetic field of the intergalactic medium (IGM)
- Are magnetic fields in galaxies and clusters primordial or generated at later epochs?
- What is the efficiency of cosmic ray acceleration in the warm-hot intergalactic medium?

# Exploration of the Unknown

- Are there new physical phenomena?
- Gamma Ray Bursts/Hypernovae/Supernovae/Neutron stars
- What is the frequency of strong radio emission from sub-stellar objects?
- Are there other detectable civilisations?

# The SKA concept

#### a large radio telescope for transformational science

- up to 1 million m<sup>2</sup> collecting area distributed over a distance of 3000+ km (<0.1 arcsec resolution)</li>
- an interferometer at frequencies from 70 MHz to 10 GHz with two or more detector technologies
- wide field of view , approx:
  - 100 sq. deg (70 to 400 MHz),
  - 30 sq. deg (0.4 to 2 GHz)
  - 1 sq. deg. (2 to 10 GHz)
- connected to fast signal transport network and high performance computing system

### providing

- 40x sensitivity of EVLA
- up to 10,000x survey speed of today's facilities



# **SKA Phase 1**

- -10% of full SKA
- -sparse aperture arrays operating at 70 to 450 MHz
- -250×15m dishes operating between 0.45 and 3 GHz
- -Advanced instrumentation program
  - Phased Array Feeds (ASKAP/APERTIF)
  - Dense Aperture Arrays



### **SKA Key Science Drivers**

#### ORIGINS

Probing the Dark Ages and the Epoch of Reionization

- EoR, first galaxies and active galactic nuclei
- Galaxy Evolution, Cosmology, and Dark Energy When did the first stars and galaxies form?

  - How did galaxies evolve?Dark Energy, Dark Matter
- ➢Cradle of Life
  - Organic molecules in interstellar space and proto-planetary disks
  - SFTI

#### **FUNDAMENTAL FORCES**

>Pulsars, General Relativity & Gravitational Waves

Origin & evolution of cosmic magnetism

**EXPLORATION OF THE UNKNOWN** Transients and new phenomena



Science with the Square Kilometre Array (2004, eds. C. Carilli & S. Rawlings, New Astron. Rev., 48)

### **SKA Phase 1 Key Science Drivers**

#### ORIGINS

Neutral hydrogen in the universe from the Epoch of Reionization to now

- EoR
- first galaxies and active galactic nuclei
- Galaxy Evolution
- Dark energy, Dark Matter

#### Cradle of Lif

- Organic molecules in interstellar space and protoplanetary disks

#### FUNDAMENTAL FORCES

>Pulsars, General Relativity & Gravitational Waves

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#### EXPLORATION OF THE UNKNOWN

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### Phase 1 (SKA<sub>1</sub>) baseline design



ALL-SKY ASTROPHYSIC

Swinburne Astronomy Productions

# **Sparse Aperture Arrays**

Omnidirectional dipoles No moving parts. Cheap!





LOFAR

SKA



### **Dishes**

#### 15m offset Gregorians



Allen Telescope Array



SKA



# **Advanced Instrumentation Program (AIP)**



### **SKA Phase 2 including AIP technologies**



### **Australia and New Zealand**



### **South Africa + 8 countries**



# The Path to the SKA

#### EVLA

NON



APERTIF

#### MeerKAT



#### FAST



#### SKA phase 2



2024







#### SKA phase 1



# **Top-level schedule for the SKA**

Technical		
2008-12	Preparatory Phase (system design and cost)	•
2013-15	Project Execution Phase (€90M pre-construction)	
2016-19	Phase 1 construction	
2016	Advanced Instrumentation Program (AIP) decisions	
2018-23	Phase 2 construction	
2020+	Science operations with Phase 1	
2024+	Science operations with Phase 2	
Programmatic		
2011	Founding Board and SPO location announced	
2011	SPO director selected (Schilizzi retires)	
2012	SKA site selection	
2014	Phase 1 construction approval (350 M€, 2007)	
2017	Phase 2 construction approval (1.2 B€, 2007)	X



# **Current Project Status**

#### SKA Founding Board formed:

- Nine countries: Australia, China, France, Germany, Italy, Netherlands, New Zealand, South Africa, UK
- Other countries likely to join in SKA2011 (Banff)
- €90M committed to pre-construction funding

#### **SKA Project Office location selected:**

Jodrell Bank Observatory

#### Siting roadmap drafted by SKA siting group:

- Siting decision in Feb 2012

#### **Project Execution Plan developed**

Work to be done and resources needed for next phase 2012 to 2015

#### **RFI** measurements taken at both candidate sites





News



#### HOME

#### NEWS

**Royal Wedding** 

Forces

Sun City

Captain Crunch

- Sun Says
- The Green House
- Weather
- SunVote
- VIDEO
- SPORT

Football

**Dream Team** 

Cricket

F1 & Motorsport

#### ÷

#### SHOWBIZ

Bizarre

Film



**Biz Sessions** 

+

TV



#### / T aps



BRITAIN is set to lead the search for alien life after being chosen as the base for the world's biggest collection of radio telescopes.

Jodrell Bank Observatory has been placed at the heart of a new £1.3billion project to investigate the mysteries of

#### GOT A STORY? EMAIL : TALKBACK@THE-SUN.CO.UK



Long wave ... Jodrell Bank Anthony Holloway



Soaps

X Factor

# Some Engineering Challenges remain...

- Low cost, high performance dishes, receivers and digital
- Phased array feed technology for large field of view
- Wideband optical-fibre signal transport systems (~20 Tbps, raw data to correlator in Phase 1)
- High performance computing (exaflop imaging machines needed)
- Data storage and retrieval (1 to 10 PB/day of image data)
- Remote/renewable power generation (~100 MW needed to run full SKA)



### **Summary**

- SKA Project Office begins operation 1 Jan 2012
- €90M committed to pre-construction
- Site decision early-2012
- Phase 1 construction 2016-19
- SKA will improve our understanding of the formation and evolution of the gaseous Universe from the EOR to the present day.







# Thank you



















