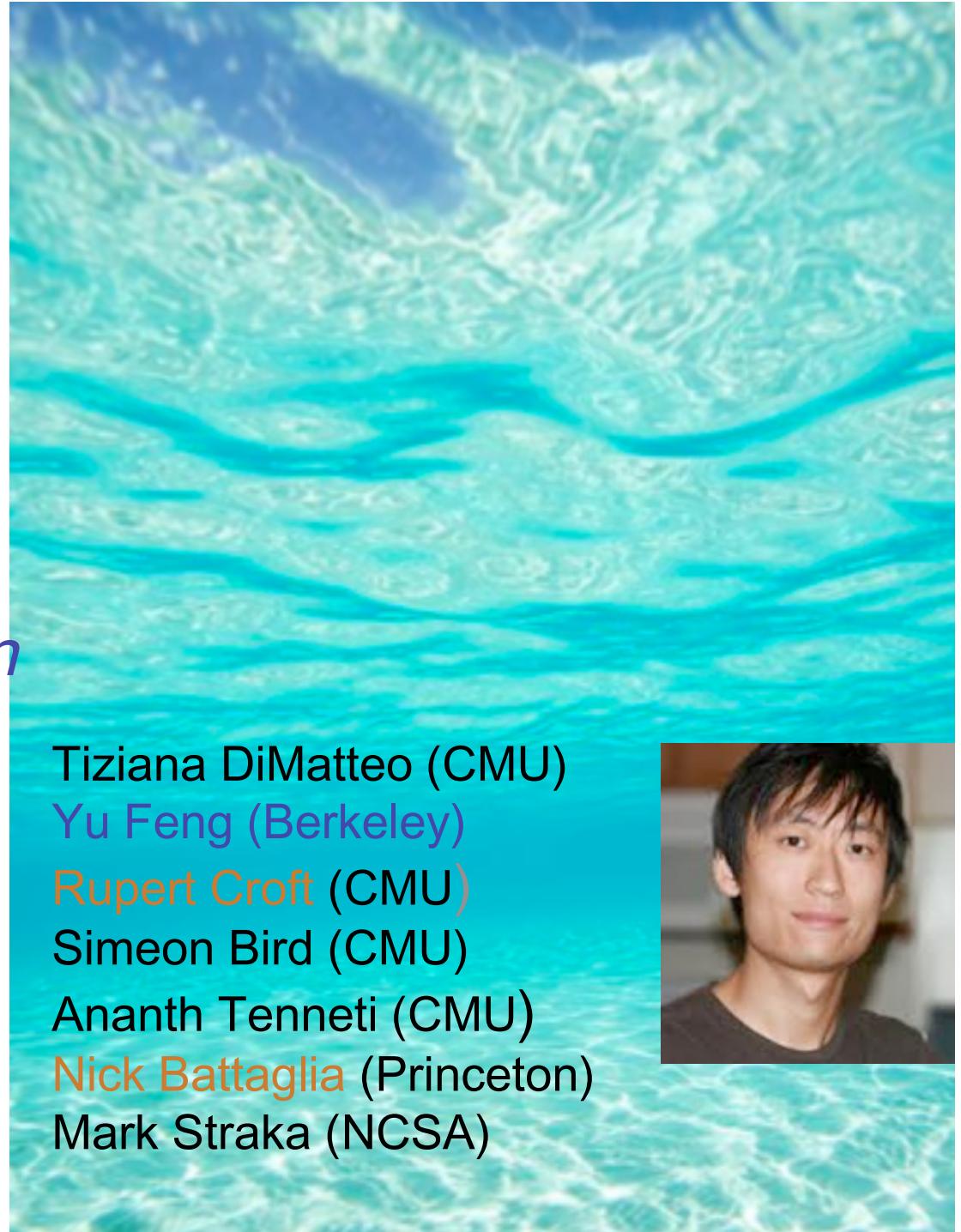


“The next frontier  
of first quasars  
and massive  
galaxies”

*Predictions from  
BlueTides Simulation*

<http://bluetides-project.org>



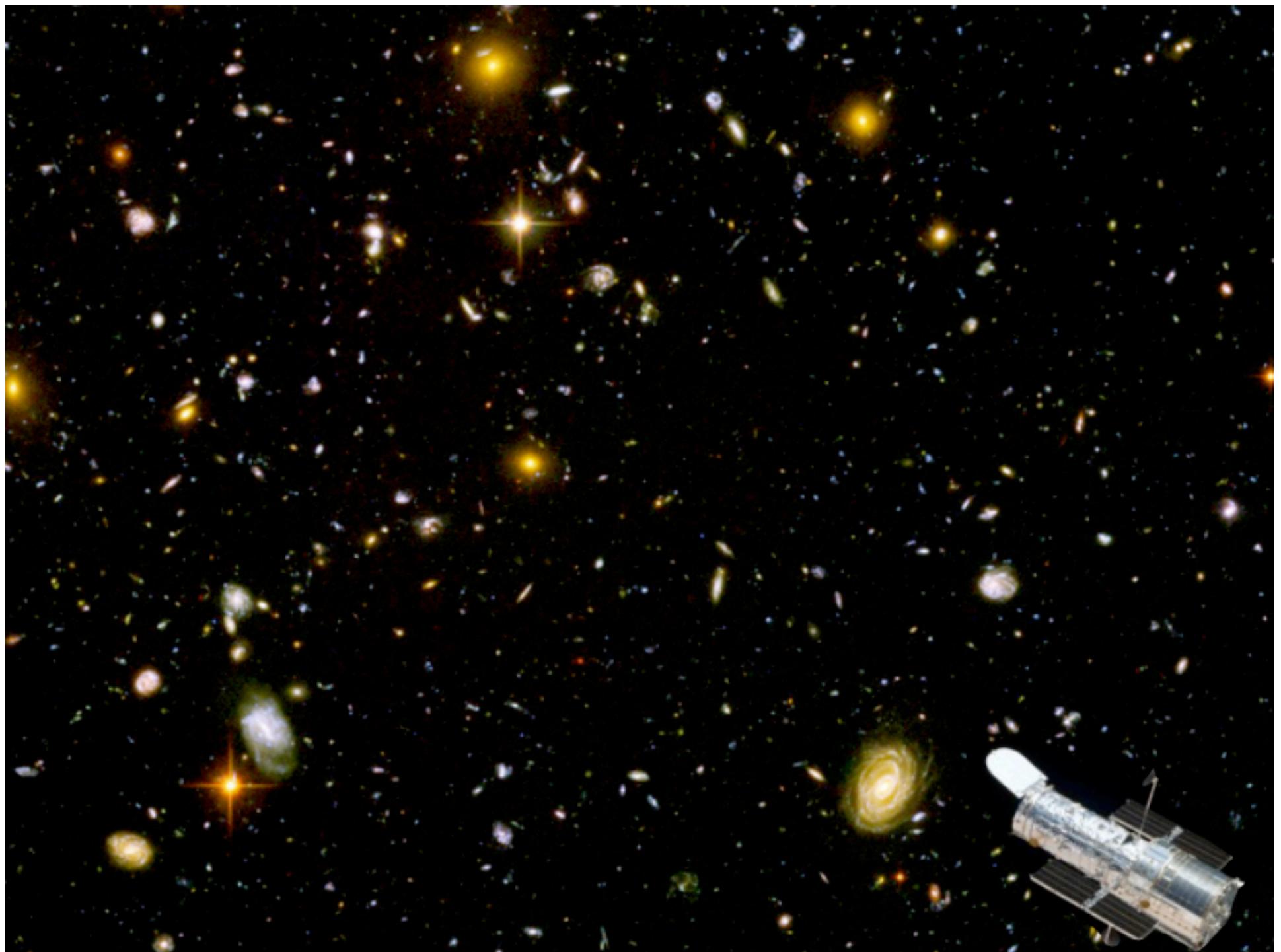
Tiziana DiMatteo (CMU)  
Yu Feng (Berkeley)  
Rupert Croft (CMU)  
Simeon Bird (CMU)  
Ananth Tennen (CMU)  
Nick Battaglia (Princeton)  
Mark Straka (NCSA)

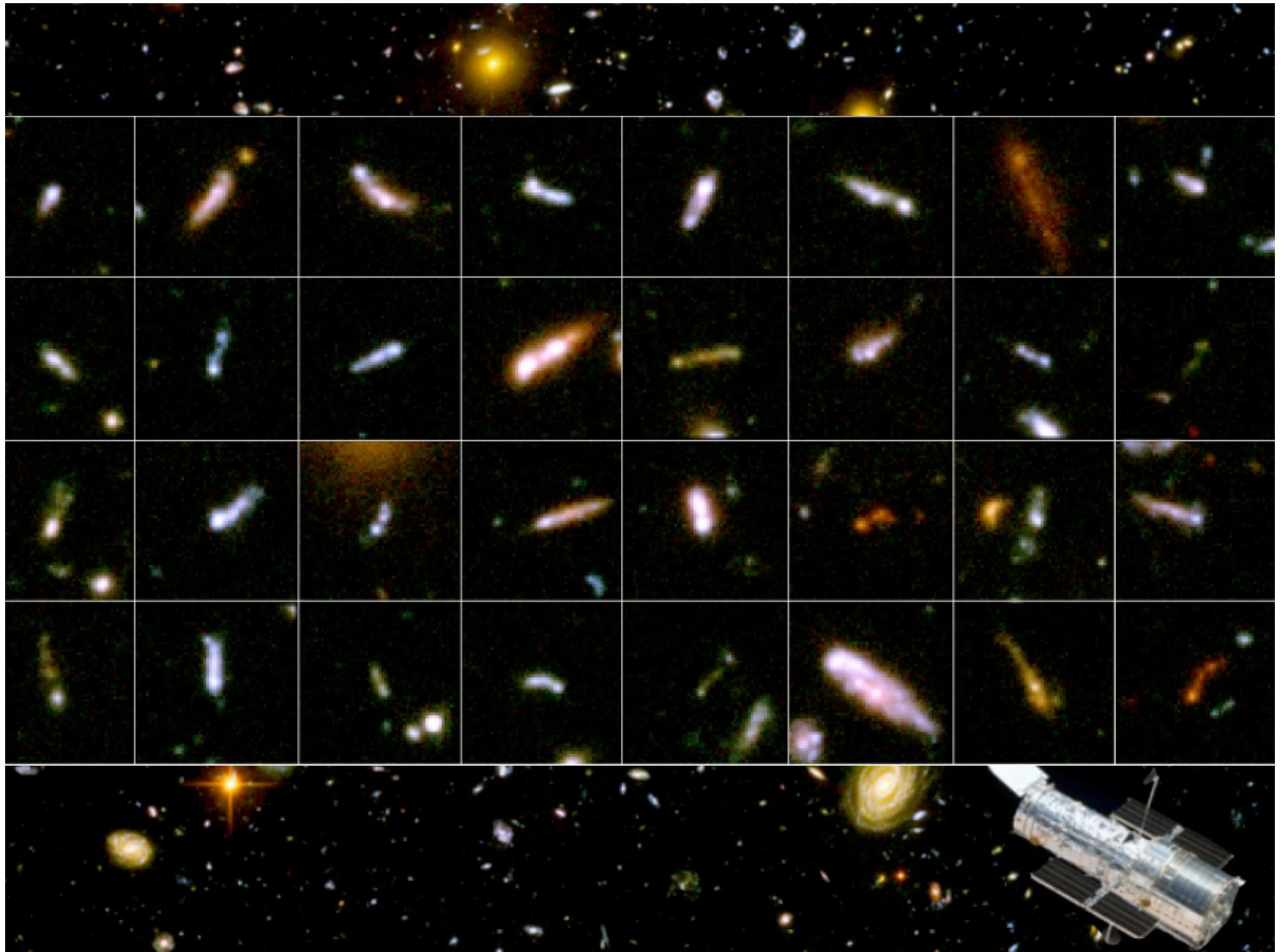


The first 600 million years ( $z=7+$ ):

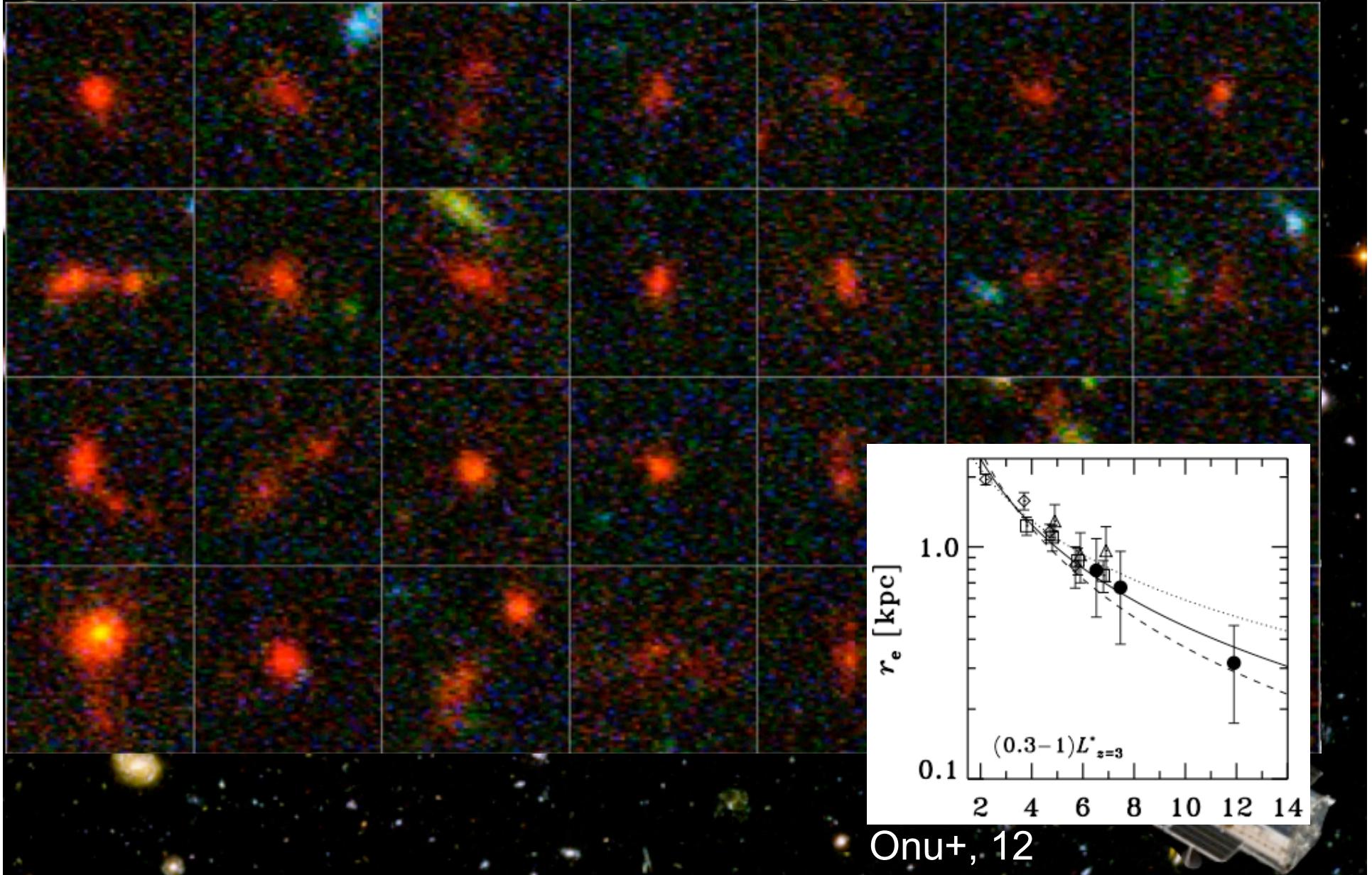
What are galaxies like?

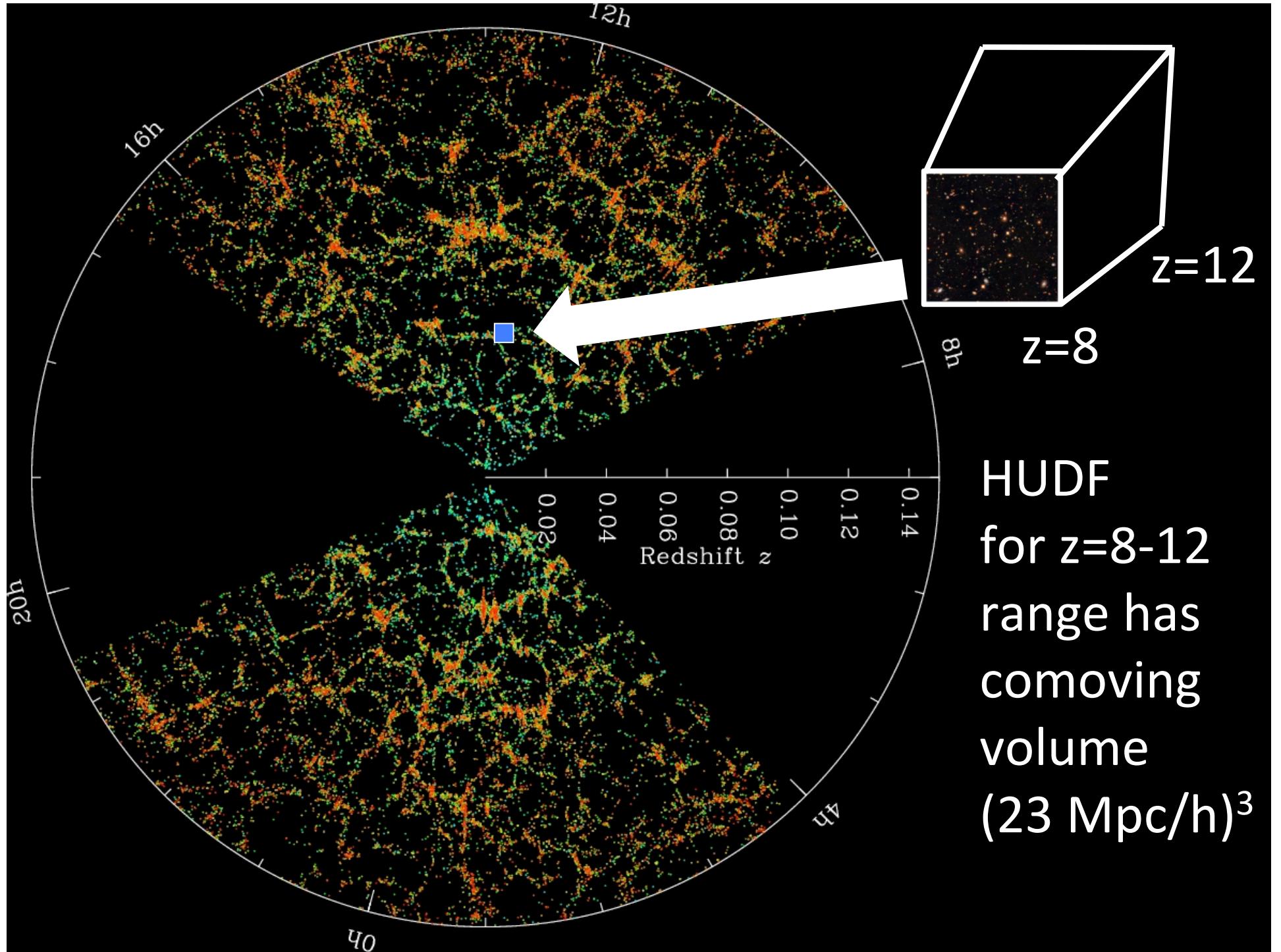
Observations:



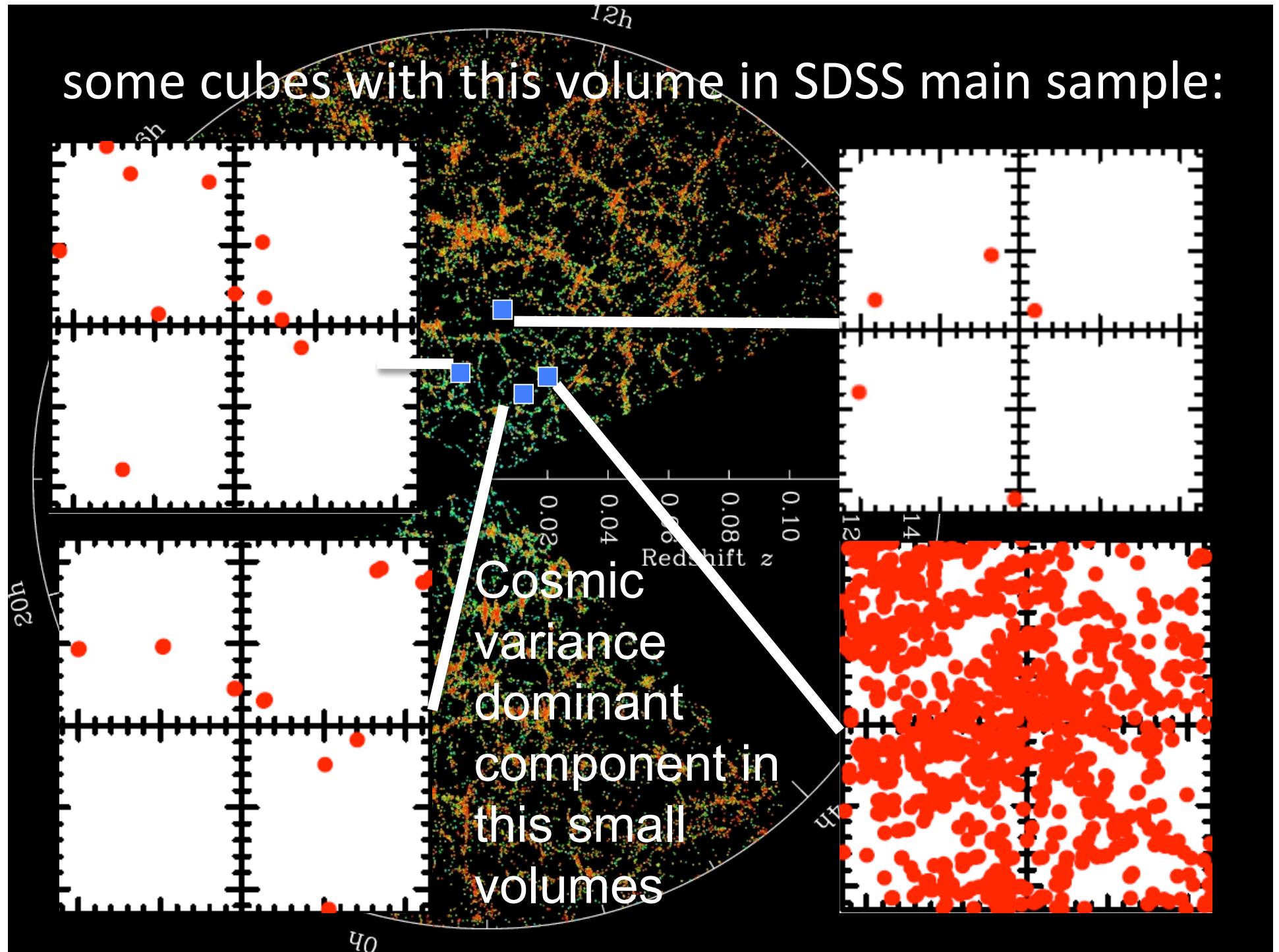


Tiny (few arcmin<sup>2</sup>) Hubble Legacy Ultra Deep Fields  
show compact, clumpy and irregular morphologies..





some cubes with this volume in SDSS main sample:



# The first and last ‘first quasar’?

LETTER TO NATURE

## A luminous quasar at a redshift of $z = 7.085$

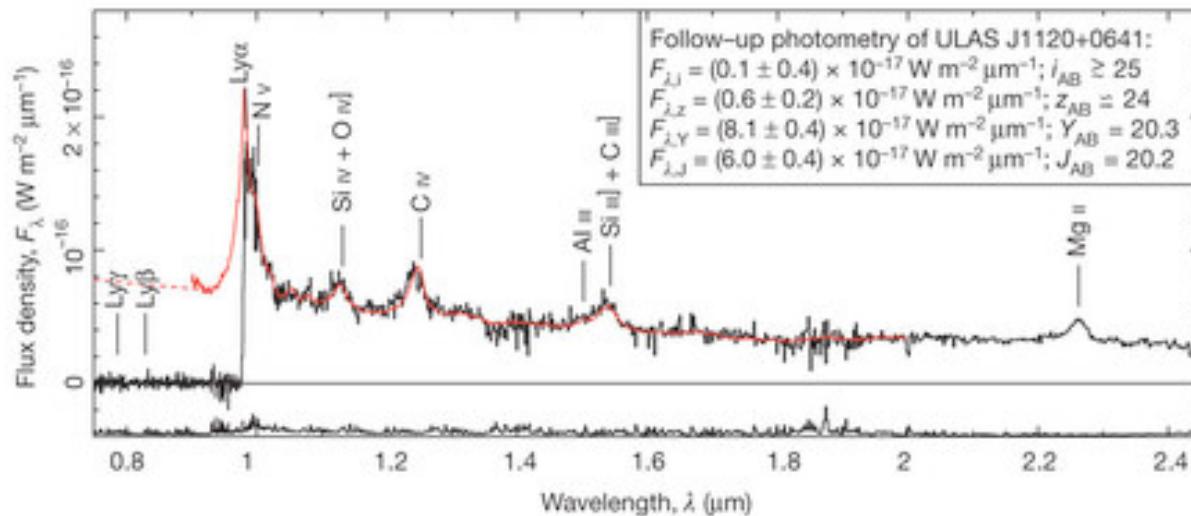
At  $z=7$ , 1 in a  $\text{Gpc}^3$

Daniel J. Mortlock<sup>1</sup>, Stephen J. Warren<sup>1</sup>, Bram P. Venemans<sup>2</sup>, Mitesh Patel<sup>1</sup>,  
Paul C. Hewett<sup>3</sup>, Richard G. McMahon<sup>3</sup>, Chris Simpson<sup>4</sup>, Tom Theuns<sup>5,6</sup>,  
Eduardo A. González-Solares<sup>3</sup>, Andy Adamson<sup>7</sup>, Simon Dye<sup>8</sup>, Nigel C. Hambly<sup>9</sup>,  
Paul Hirst<sup>10</sup>, Mike J. Irwin<sup>3</sup>, Ernst Kuiper<sup>11</sup>, Andy Lawrence<sup>9</sup>  
& Huub J. A. Röttgering<sup>11</sup>

| 30 Jun 2011

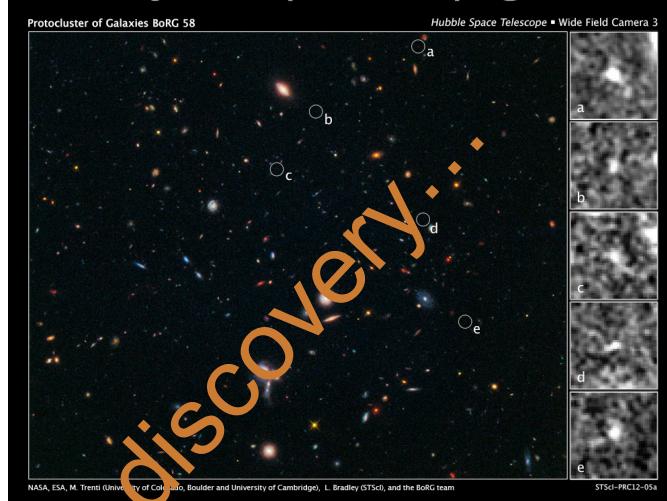
The intergalactic medium was not completely reionized until approximately a billion years after the Big Bang, as revealed<sup>1,2</sup> by observations of quasars with redshifts of less than 6.5. It has been difficult to probe to higher redshifts, however, because quasars have historically been identified<sup>3,4</sup> in optical surveys, which are insensitive to sources at redshifts exceeding 6.5. Here we report observations of a quasar (ULAS J112001.48+064124.3) at a red-

shift spectroscopically confirmed to have even higher redshifts, two are faint  $J_{AB} \gtrsim 26$  galaxies<sup>5,6,7</sup> and the other is a  $\gamma$ -ray burst which has since faded<sup>8,9</sup>. Indeed, it has not been possible to obtain high signal-to-noise ratio spectroscopy of any sources beyond the most distant quasars previously known: CFHQS J0210–0456<sup>10</sup> ( $z = 6.44$ ), SDSS 1148+5251<sup>11</sup> ( $z = 6.42$ ) and CFHQS J2329+0301<sup>12</sup> ( $z = 6.42$ ). Follow-up measurements of ULAS J1120+0641 will provide the first opportunity to explore the 0.1 Gyr between  $z = 7.08$  and  $z = 6.44$ ,



At  $z > 7$ :

Plenty of ('faint') galaxies in tiny volumes



Hubble Legacy  
Fields

1 quasar

LETTER TO NATURE

A luminous quasar at a redshift of  $z = 7.085$

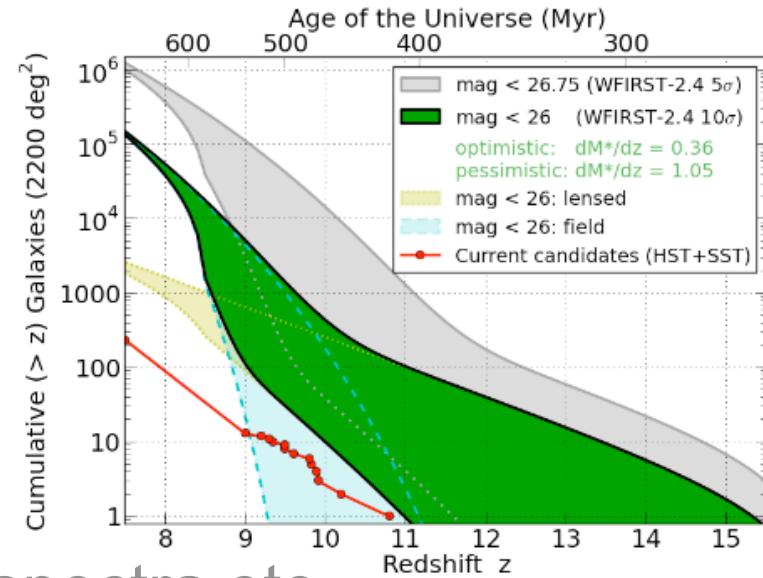
Daniel J. Mortlock<sup>1</sup>, Stephen J. Warren<sup>1</sup>, Bram P. Venemans<sup>2</sup>, Mitesh Patel<sup>1</sup>, Paul C. Hewett<sup>3</sup>, Richard G. McMahon<sup>3</sup>, Chris Simpson<sup>4</sup>, Tom Theuns<sup>5,6</sup>, Eduardo A. González-Solares<sup>3</sup>, Andy Adamson<sup>7</sup>, Simon Dye<sup>8</sup>, Nigel C. Hambly<sup>9</sup>, Paul Hirst<sup>10</sup>, Mike J. Irwin<sup>3</sup>, Ernst Kuiper<sup>11</sup>, Andy Lawrence<sup>9</sup> & Huub J. A. Röttgering<sup>11</sup>

01 Jun 2011

The intergalactic medium was not completely reionized until approximately a billion years after the Big Bang, as revealed<sup>12</sup> by observations of quasars with redshifts of less than 6.5. It has been difficult to probe to higher redshifts, however, because quasars have historically been identified<sup>2,3</sup> in optical surveys, which are insensitive to sources at redshifts exceeding 6.5. Here we report observations of a quasar (ULAS J112001.48+064124.3) at a redshift of 7.085, which is 0.77 billion years after the Big Bang. ULAS J1120+0641 had a luminosity of

troscopically confirmed to have even higher redshifts, two are faint  $JAB \gtrsim 26$  galaxies<sup>13,14</sup> and the other is a  $\gamma$ -ray burst which has since faded<sup>12</sup>. Indeed, it has not been possible to obtain high signal-to-noise ratio spectroscopy of any source beyond the most distant quasars previously known: CFHQS J0210-045<sup>15</sup> ( $z = 6.44$ ), SDSS 1148+5251<sup>16</sup> ( $z = 6.42$ ) and CFHQS J2329+0301<sup>17</sup> ( $z = 6.42$ ). Follow-up measurements of ULAS J1120+0641 will provide the first opportunity to explore the 0.1 Gyr between  $z = 7.08$  and  $z = 6.44$ , a significant cosmological epoch about which little is cur-

WFIRST > 50,000 galaxies  
+quasars



JWST spectra etc..

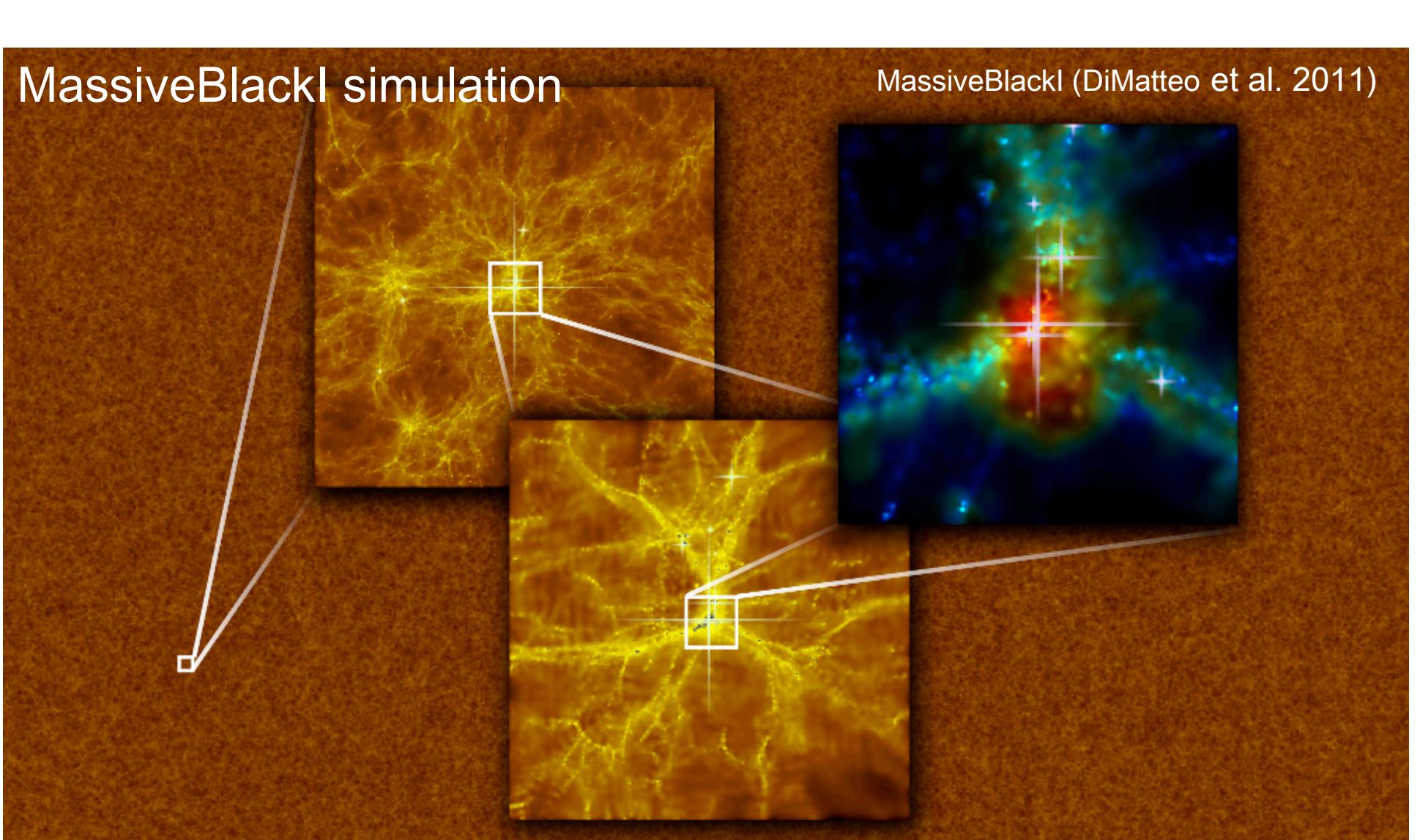
*Room for discovery.*

The first 500 million years ( $z=7+$ ):

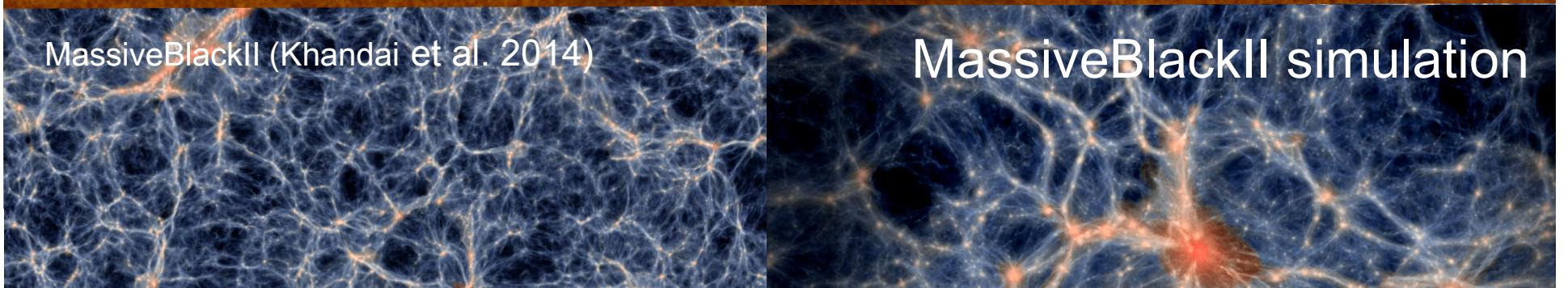
What are galaxies like?

Predictions:

MassiveBlackI simulation



MassiveBlackI (DiMatteo et al. 2011)



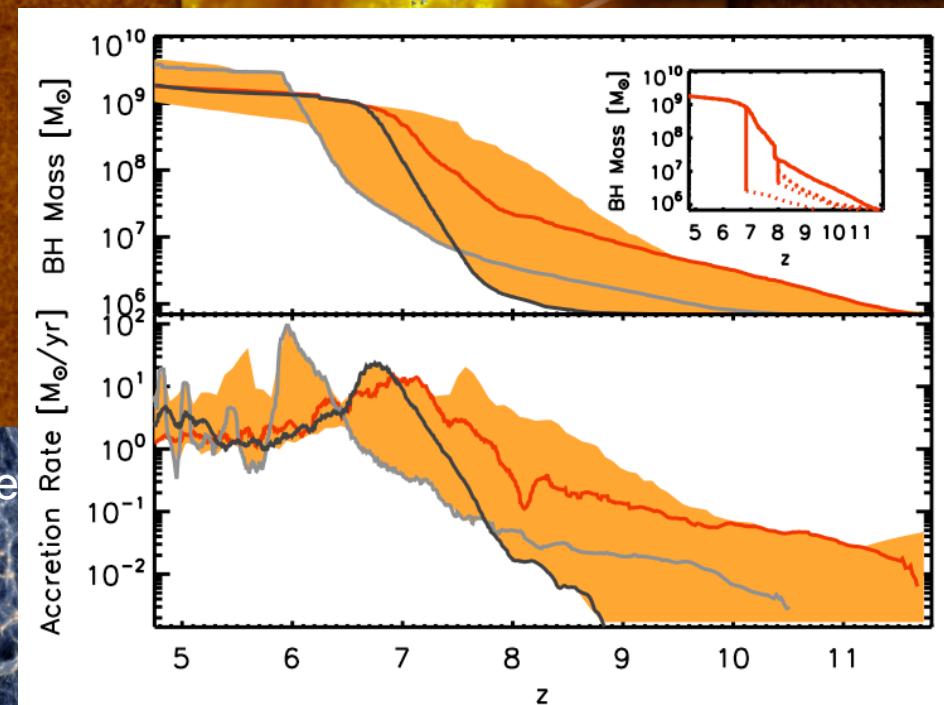
MassiveBlackII (Khandai et al. 2014)

MassiveBlackII simulation

MassiveBlackI simulation

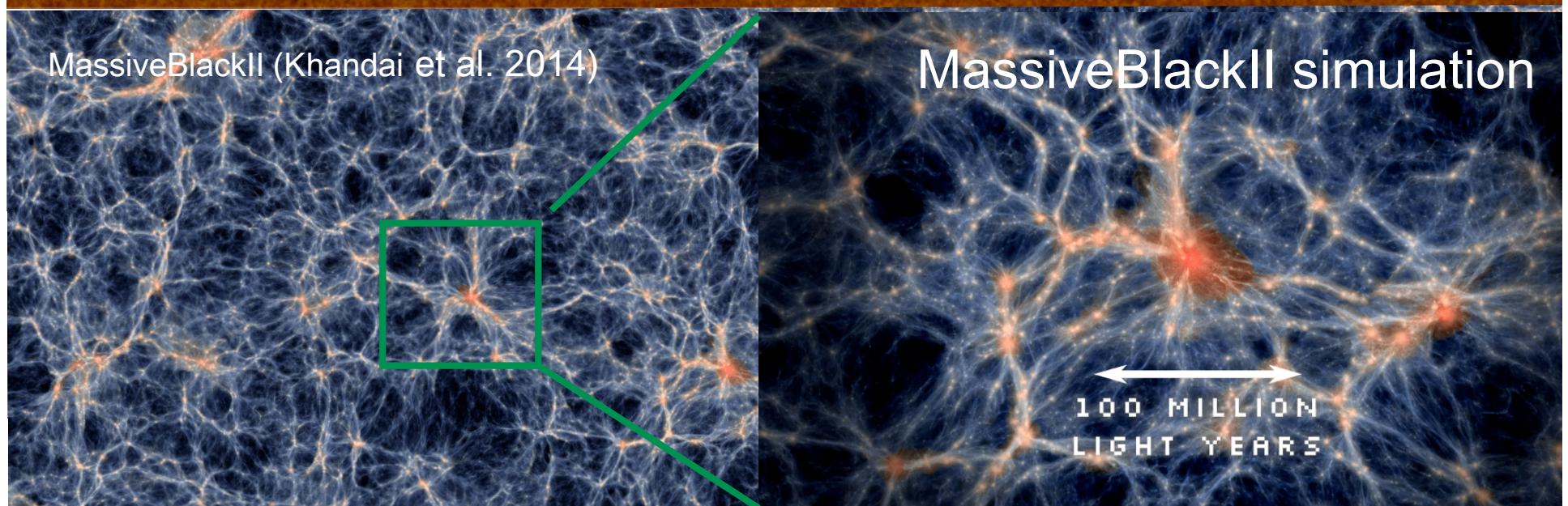
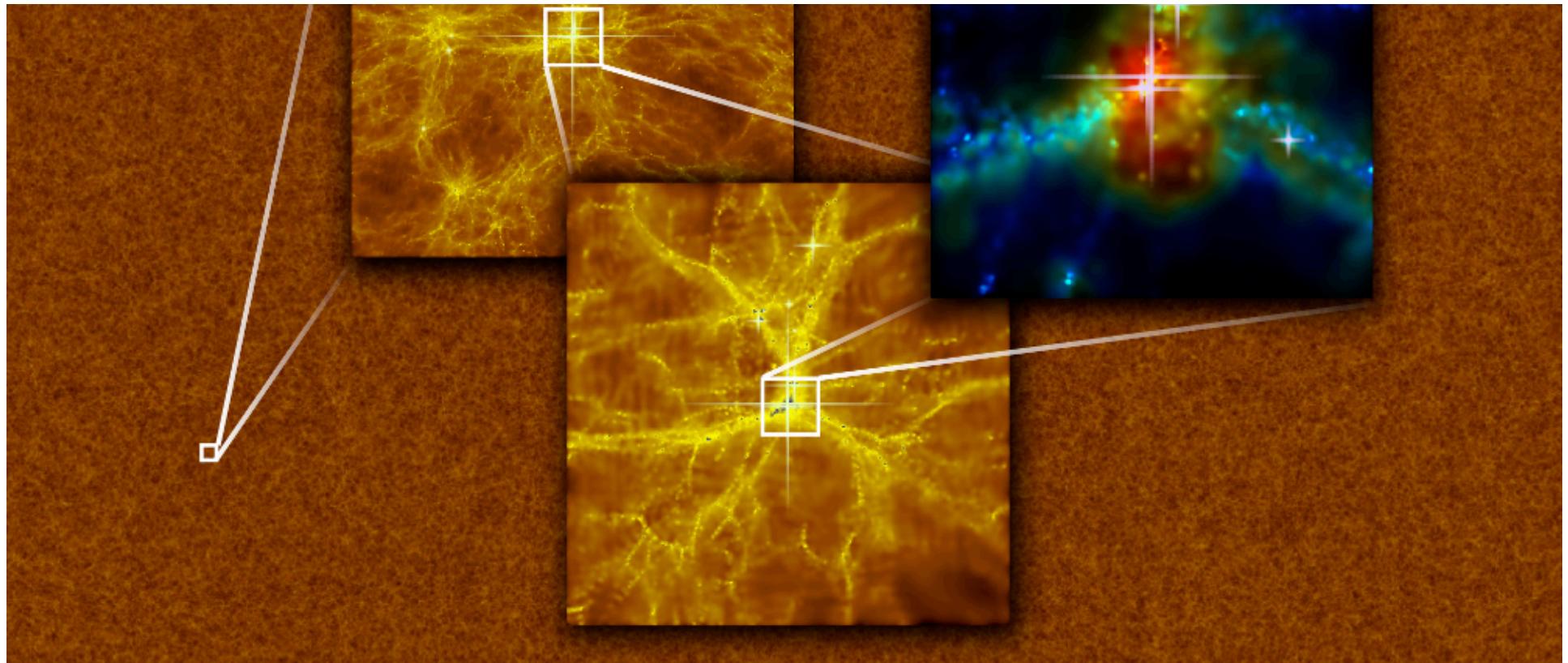
MassiveBlackI (DiMatteo et al. 2012)

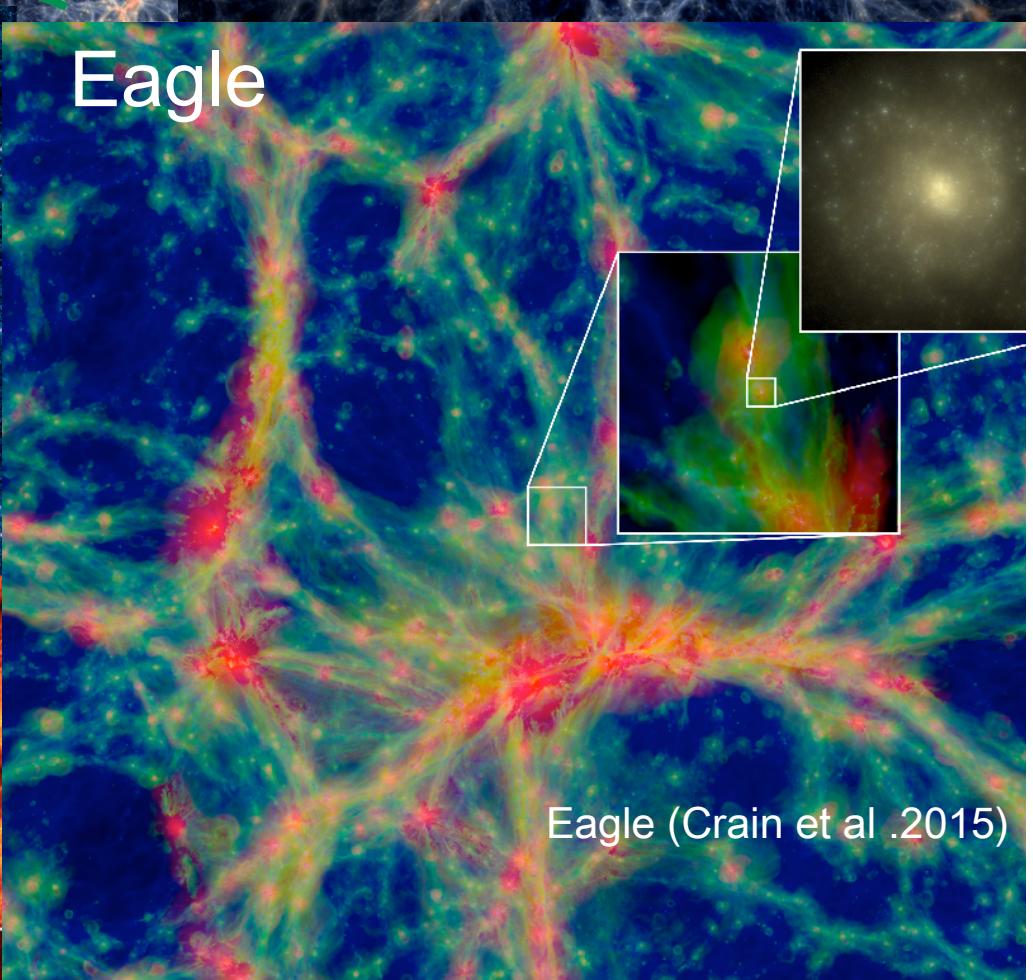
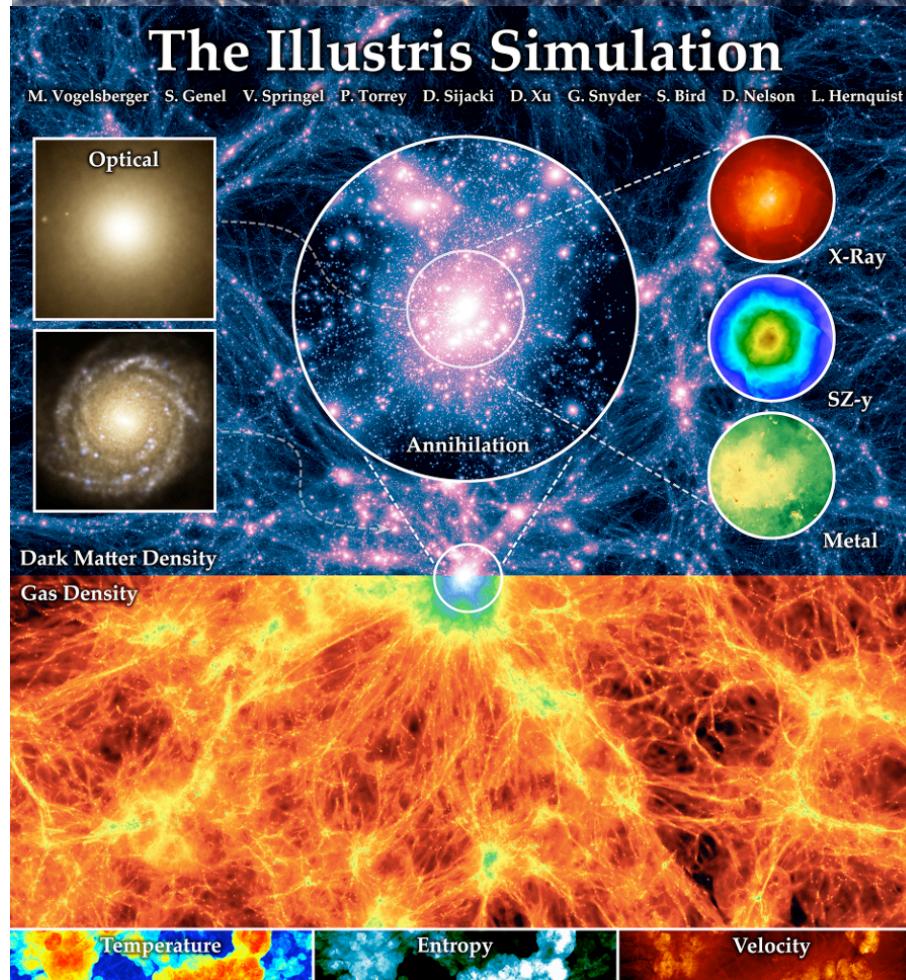
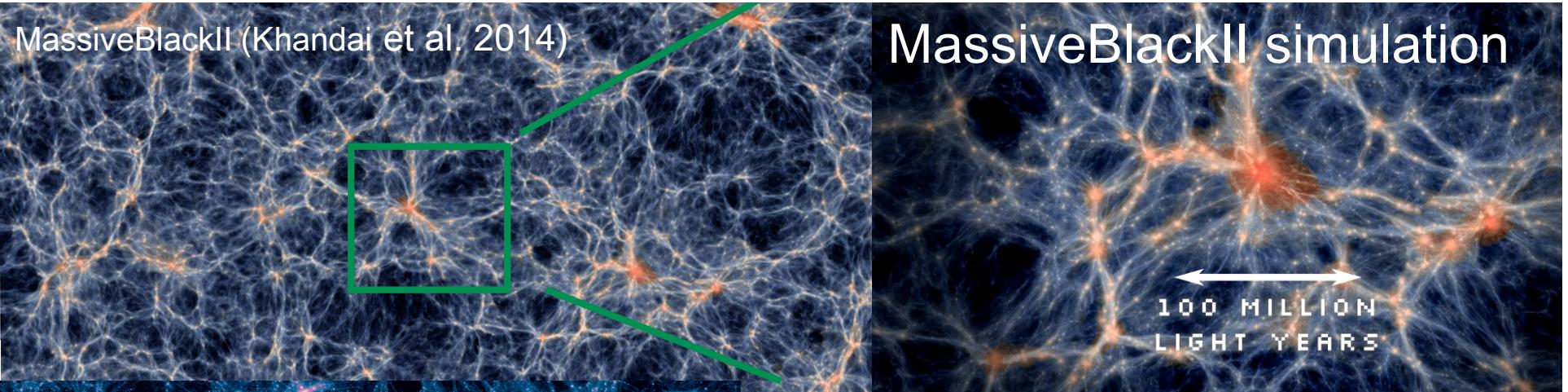
Black holes grow to  $10^9 M_{\odot}$  by  $z=7-6$



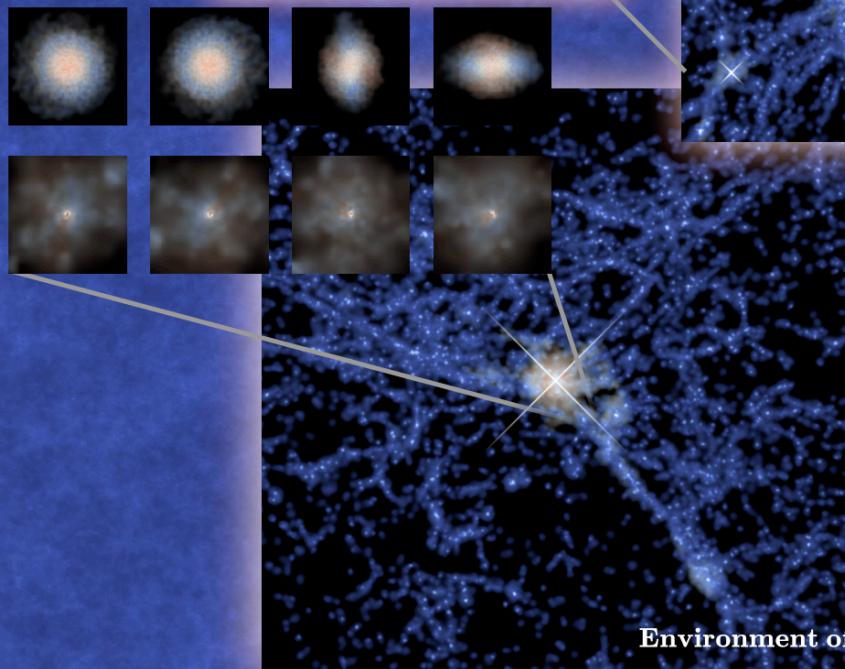
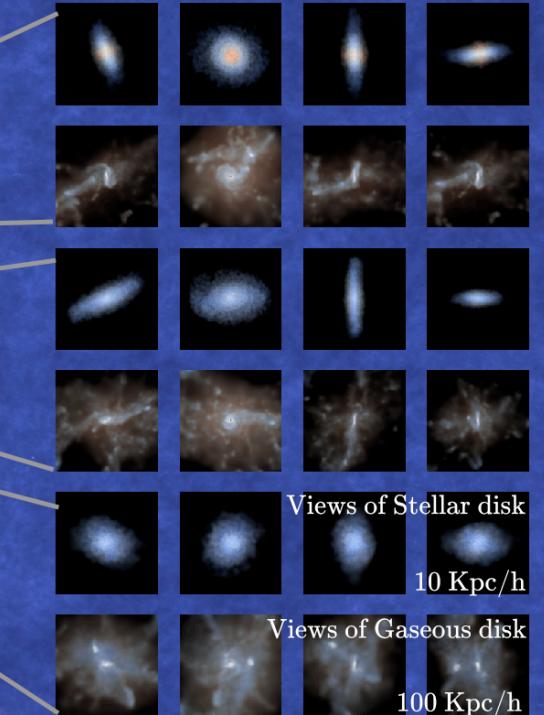
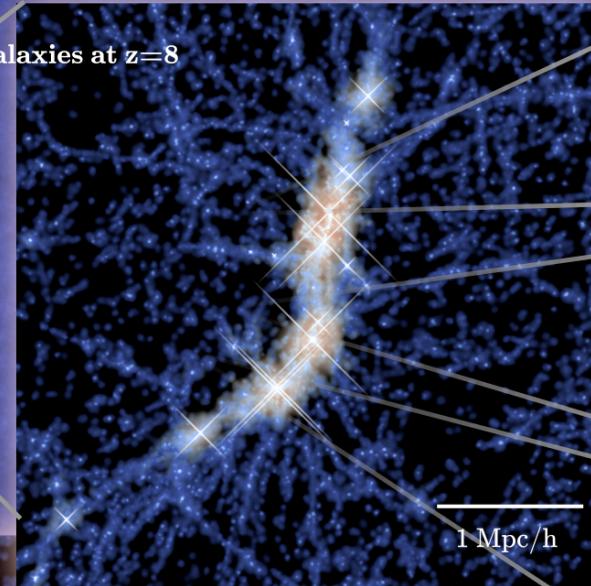
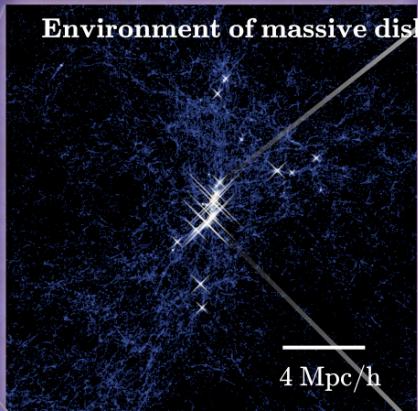
MassiveBlackII (Khandai et al. 2012)

MassiveBlackII simulation





Eagle (Crain et al .2015)



Environment of most massive blackhole at  $z=8$

The **Blue Tides** Simulation

0.7 trillion particles  
0.65 million cores

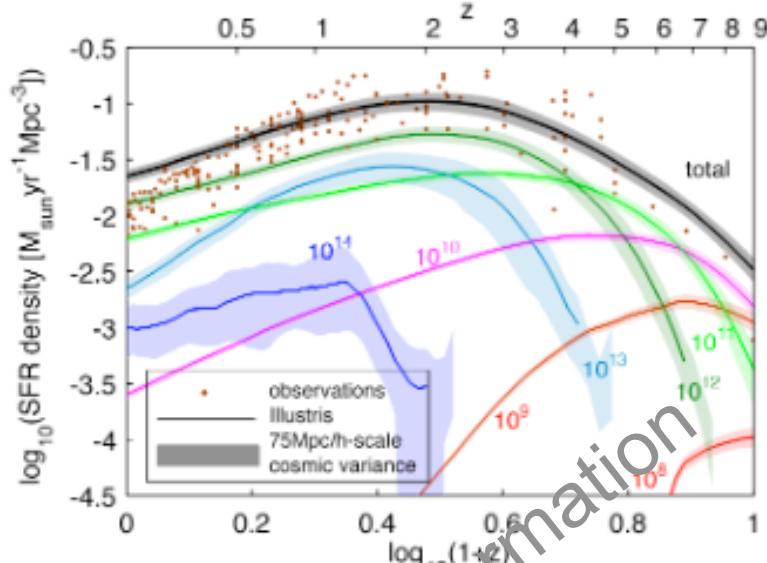
BLUE WATERS  
SUSTAINED PETASCALE COMPUTING



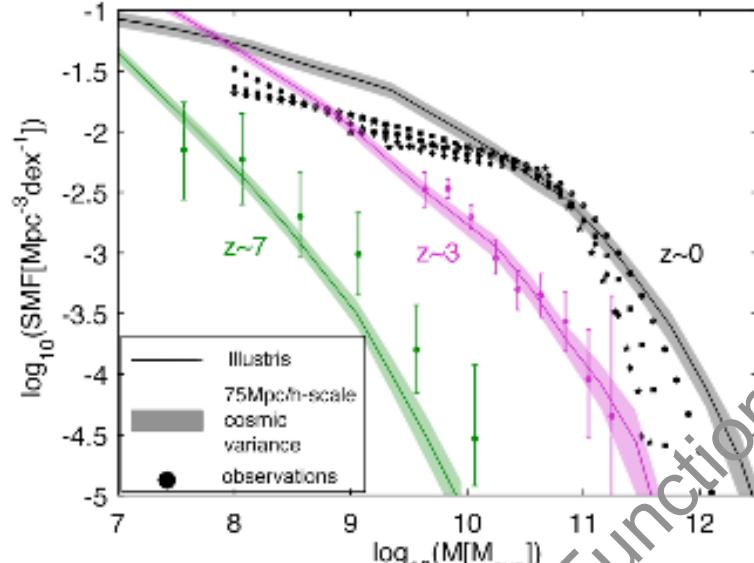
blueTides

Feng et al. 2015

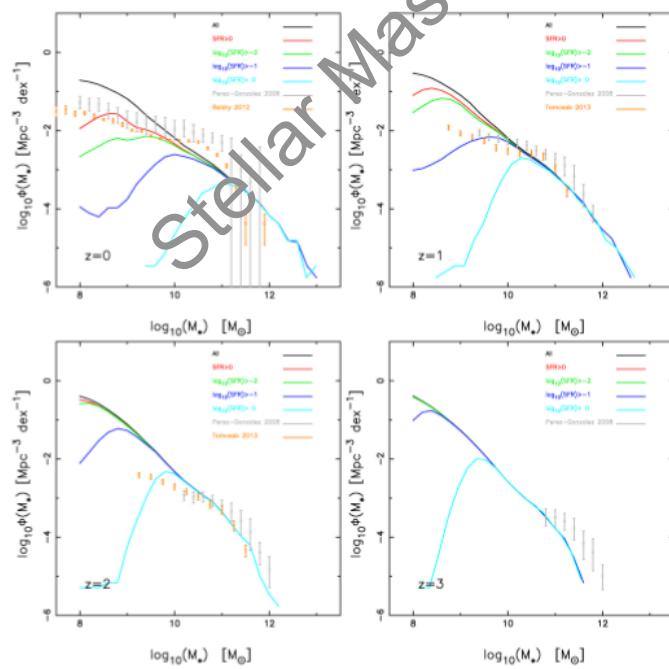
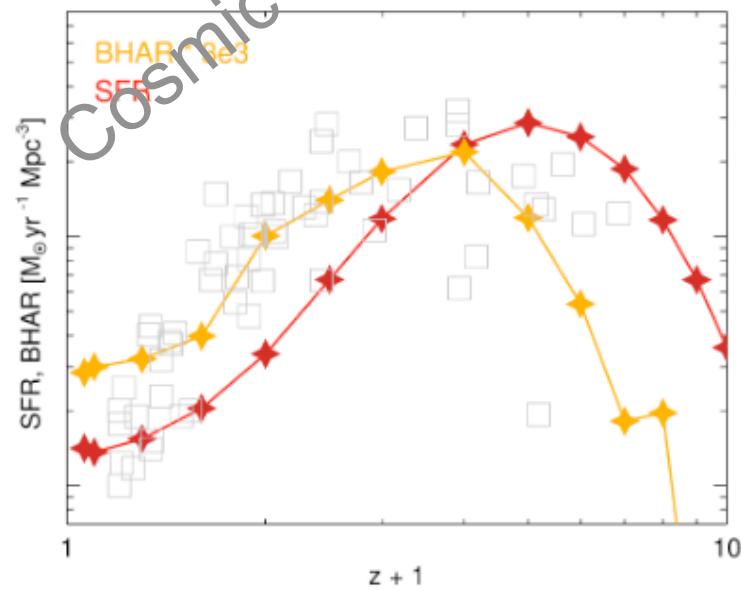
# Simulations reproduce statistics of galaxy formation

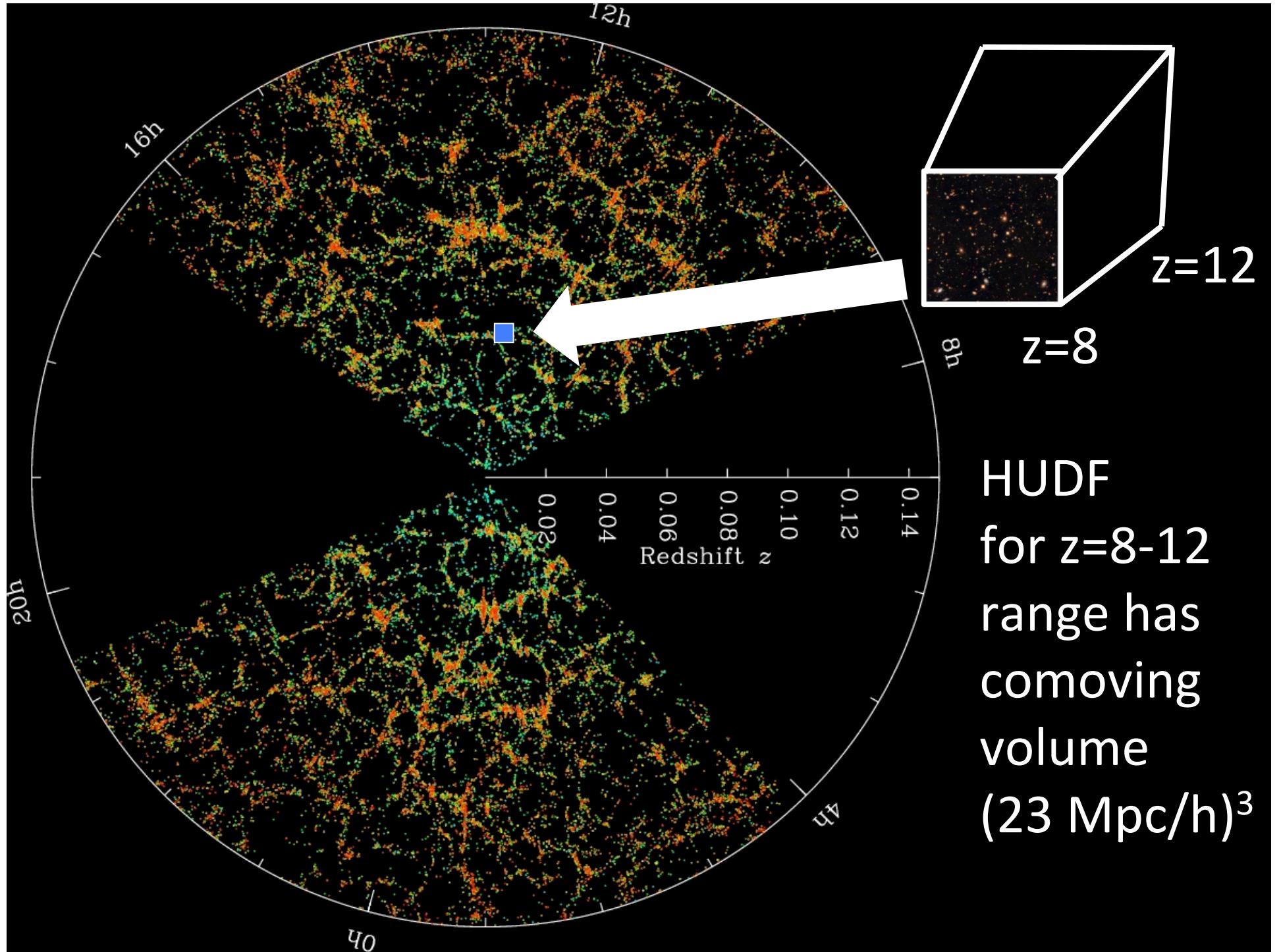


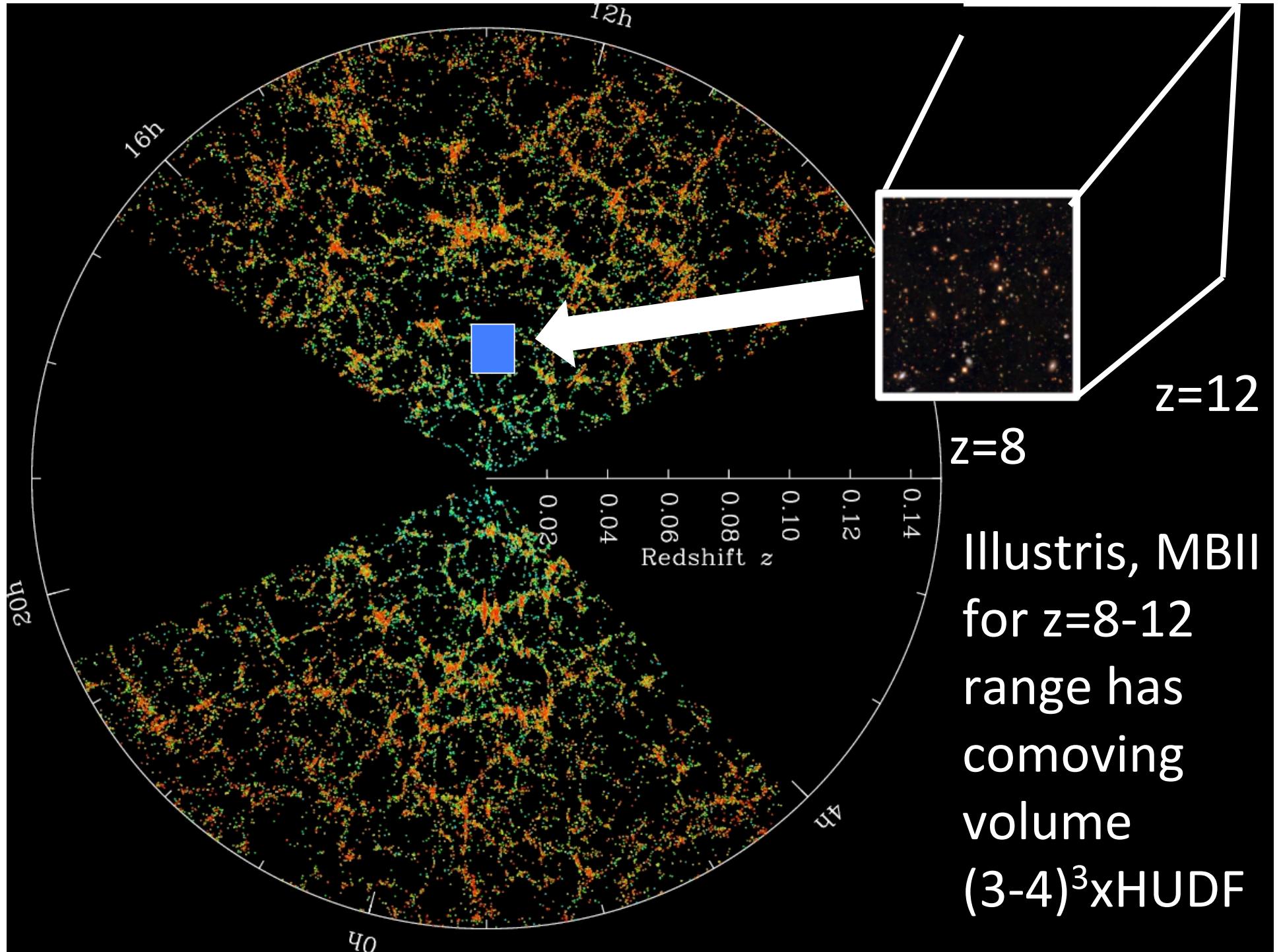
(c) Cosmic SFR density, 106.5 Mpc cosmic variance



(d) Stellar mass functions, 106.5 Mpc cosmic variance







However, theoretical predictions lacking at  $z=7+$   
simulations have either: **insufficient resolution or**  
**too small volumes for**  
**massive objects**

MassiveBlack I (DiMatteo et al 2011) 533 Mpc/h & 5.5 Kpc/h

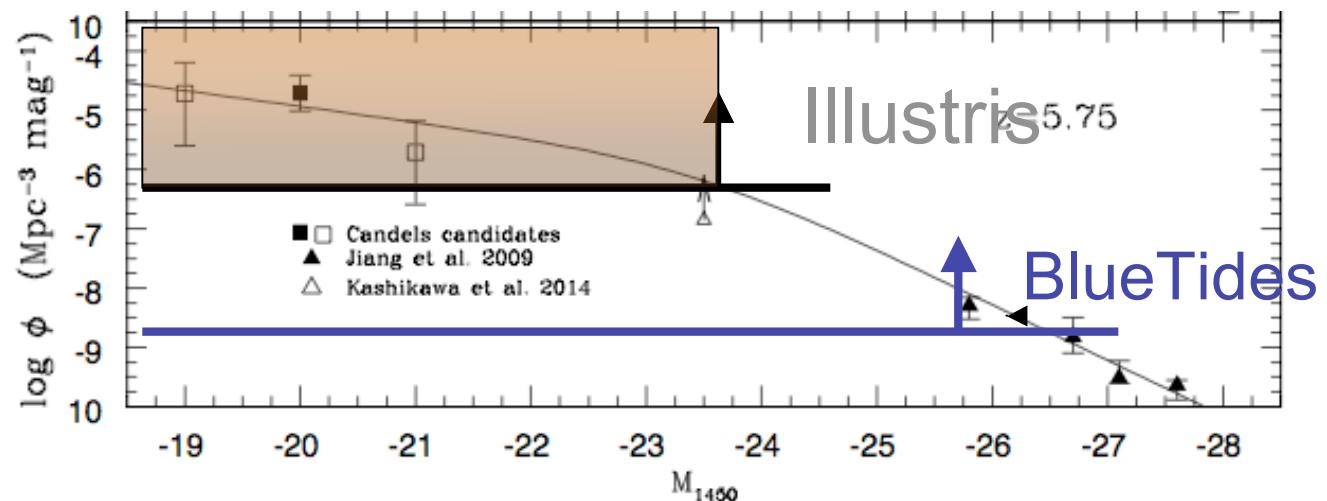
MassiveBlack II (Khandai et al 2014) 100 Mpc/h & 1.8 Kpc/h

Illustris (Vogelsberger et al 2014) : 72.5 Mpc/h & 1 Kpc/h

100Mpc/h  $\sim 0.2$  sq degrees  
at  $z=8-9$

Eagles (Crain et al 2015) : 100 Mpc/h 2.6 Kpc/h

e.g: High-z  
QSO lum. Function  
(Giallongo 15)



# BlueTides Simulation:

0.7 million cores on NCSA BlueWaters



## Goals:

- Technology Path Finder for future hydro simulation
- Predictions for high-redshift surveys

714 Mpc on the side

200 pc resolution at  $z=9$

$2 \times 7040^3$  (0.7 trillion) particles

Star formation/ AGN model compatible with Illustris

**50 times bigger volume** with highest resolution

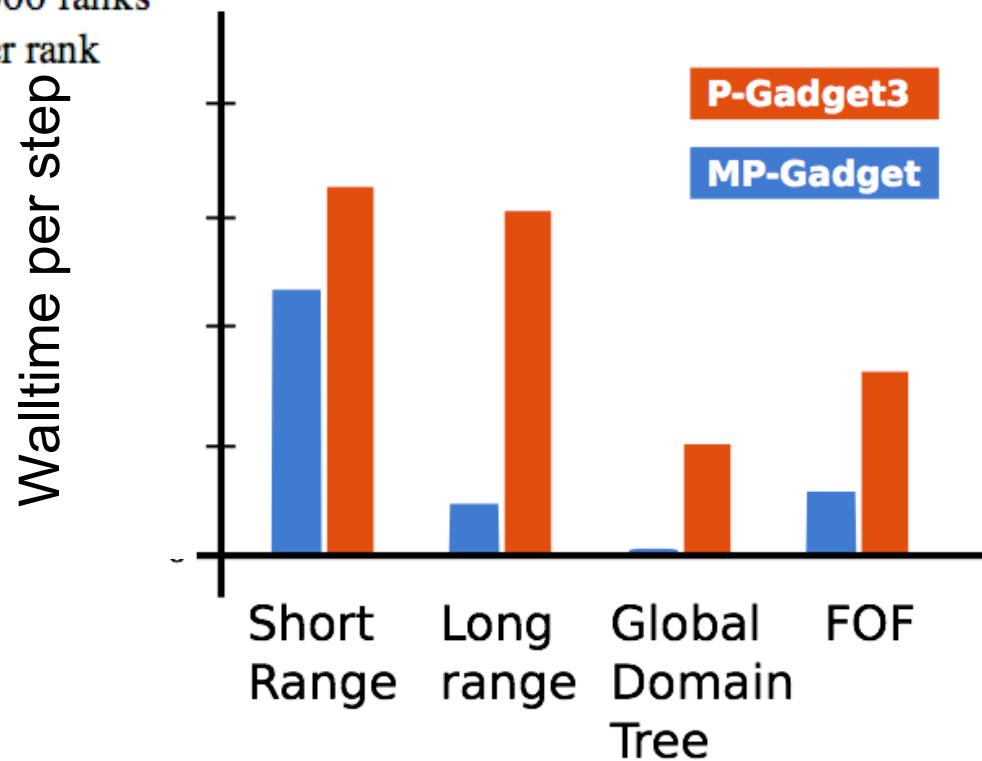
# BlueTides Simulation: Technology

## Technology Path-Finder:

- From P-Gadget to MP-Gadget
- 81000 MPI ranks, 8 OpenMP threads per rank
- Large, distributed FFT: 10,000 mesh on 81000 ranks
- Efficient thread parallelism up to 32 threads per rank

## Spinoffs and Open source contributions:

- bigfile : hierarchical snapshot format
- MP-sort : practical parallel sorting
- sharedmem : parallel data analysis
- PFFT : large-scale distributed FFT



## Long range force calculation (PM): New solver:

E.g. 8 processes:

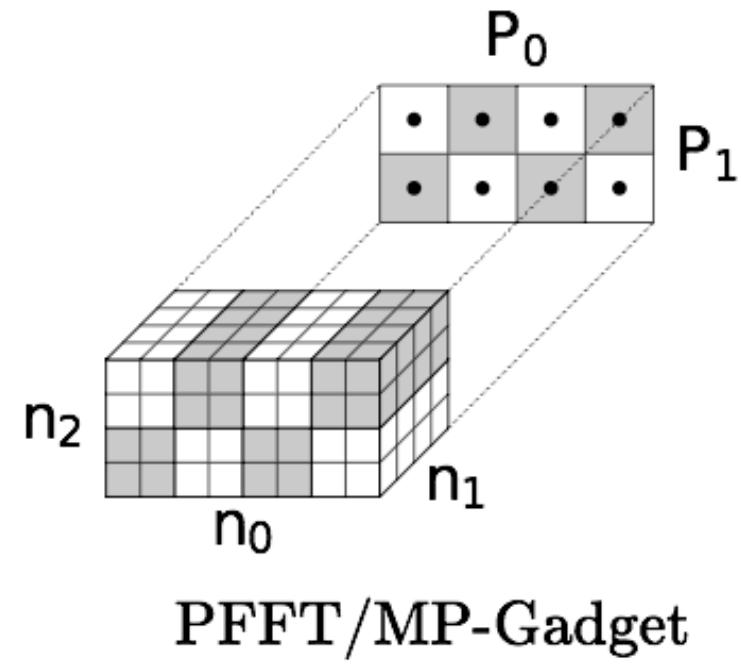
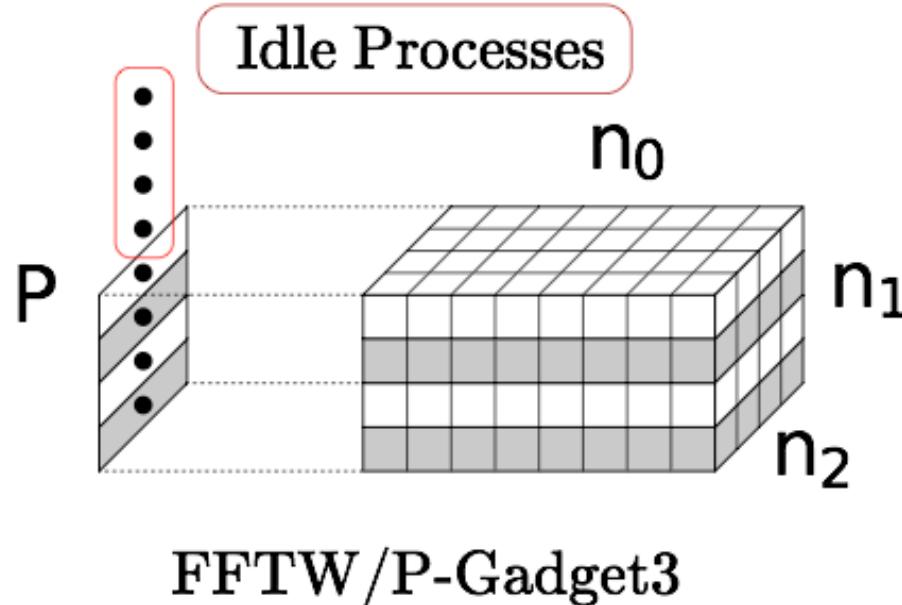


Figure from M.Pippig 2013

Blue Tides:

$N = 10000$  slabs  
on 81000 MPI ranks



Pencil beam domain  
decomposition

8 x speed-up

Open Source: Added new Array-execution interface and python binding to PFFT  
(<http://github.com/mpip/pfft>)

# BlueTides Simulation: Science

## Physics modelling in BlueTides

- Hydrodynamics (pSPH)
- Primodial cooling
- Multi-phase medium star-formation
- SN wind feedback
- H<sub>2</sub> molecule fraction
- AGN feedback
- Metal enrichment and cooling
- Non-uniform UV background calibrated from rad. Hydro sims (**Battaglia+13**)

## Science of high redshift galaxy

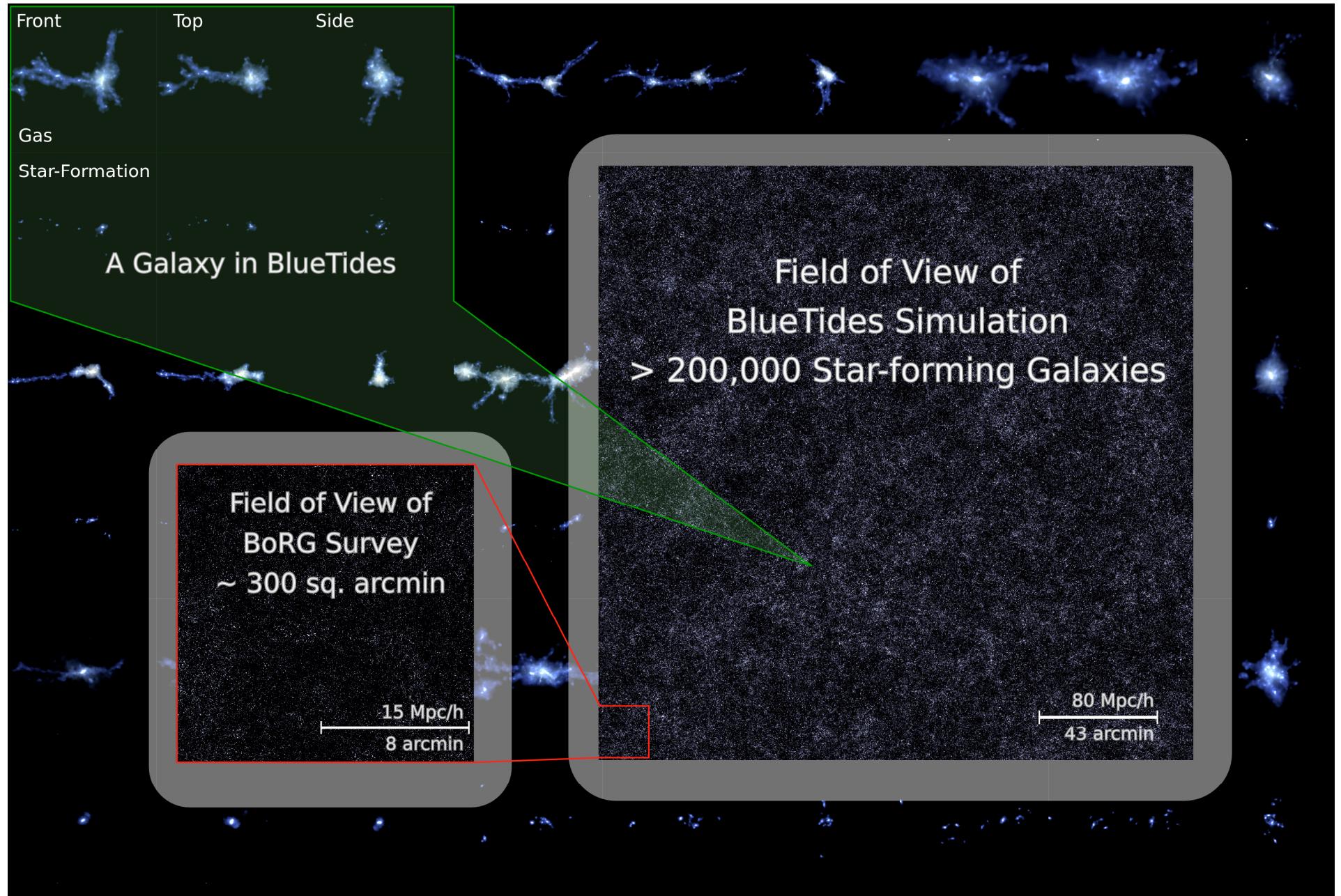
- a statistical sample of high redshift galaxies, accessible only via uniform simulations
- reionization
- morphology
- mock surveys
- high redshift AGNs
- ....

The first 500 million years ( $z=7+$ ):

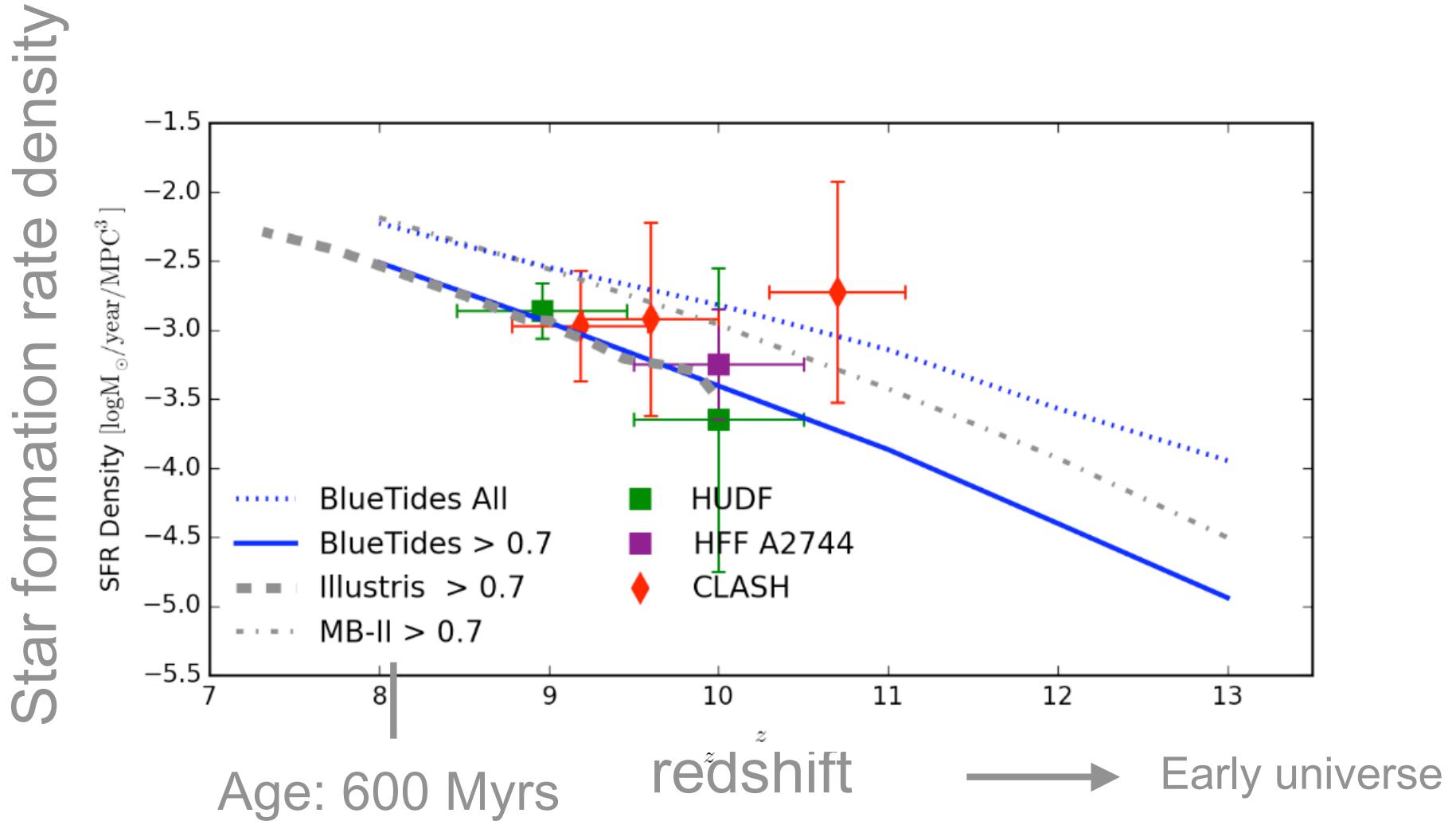
What are galaxies like?

Predictions from  
BlueTides:

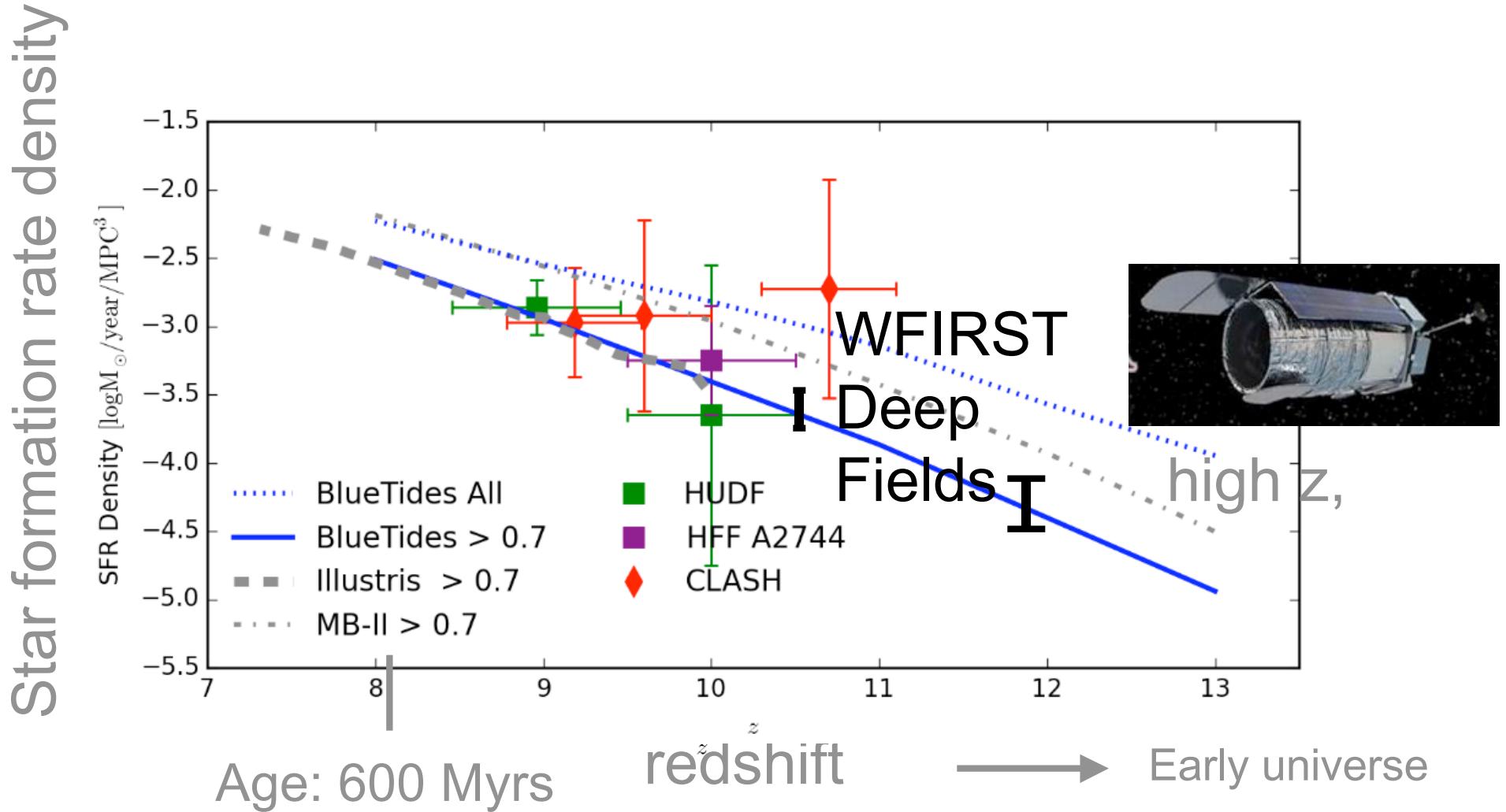
# BlueTides 400 x volume of HUDF



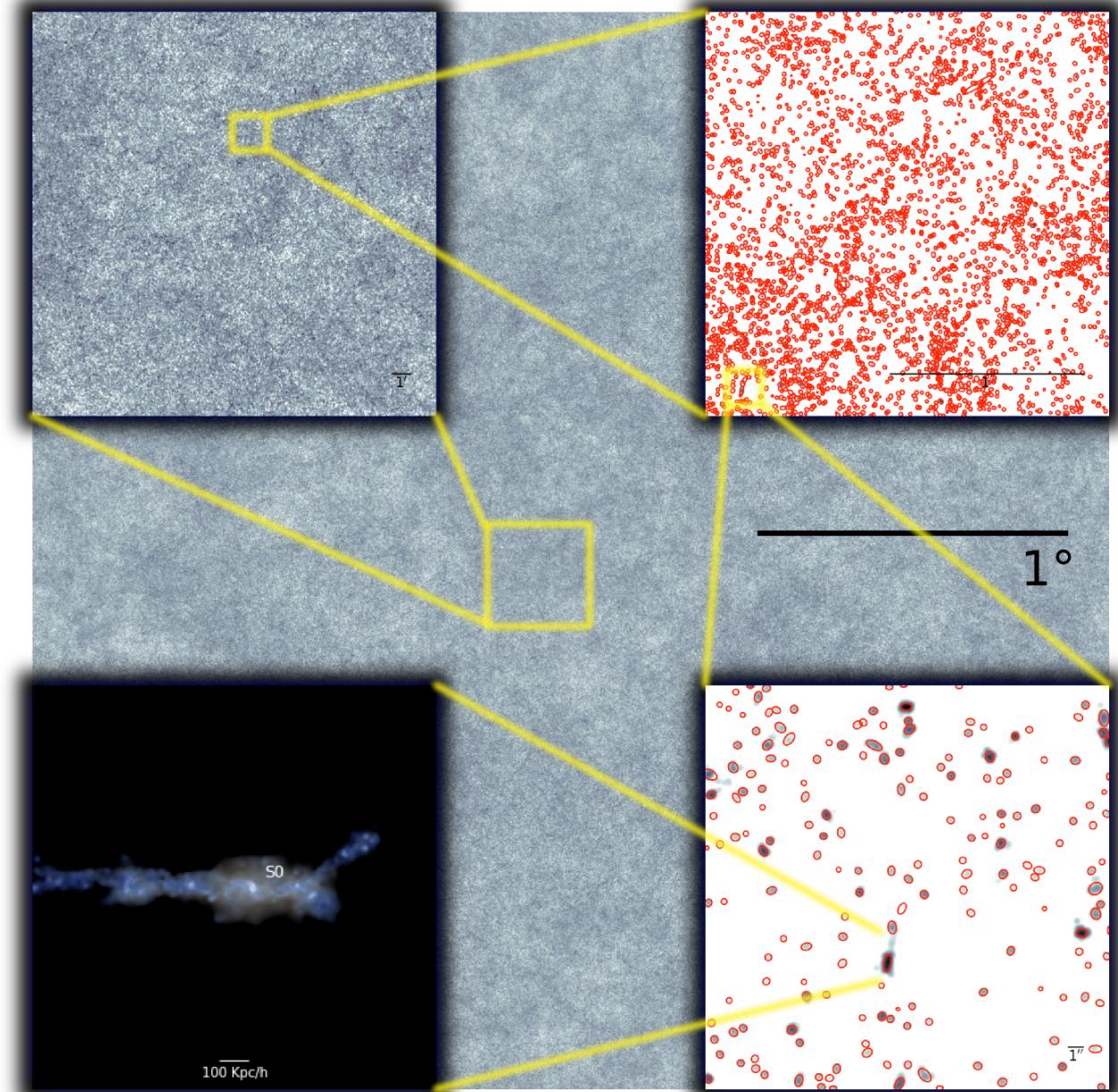
# BlueTides Simulation: Global SFRD is consistent with current observational limits.



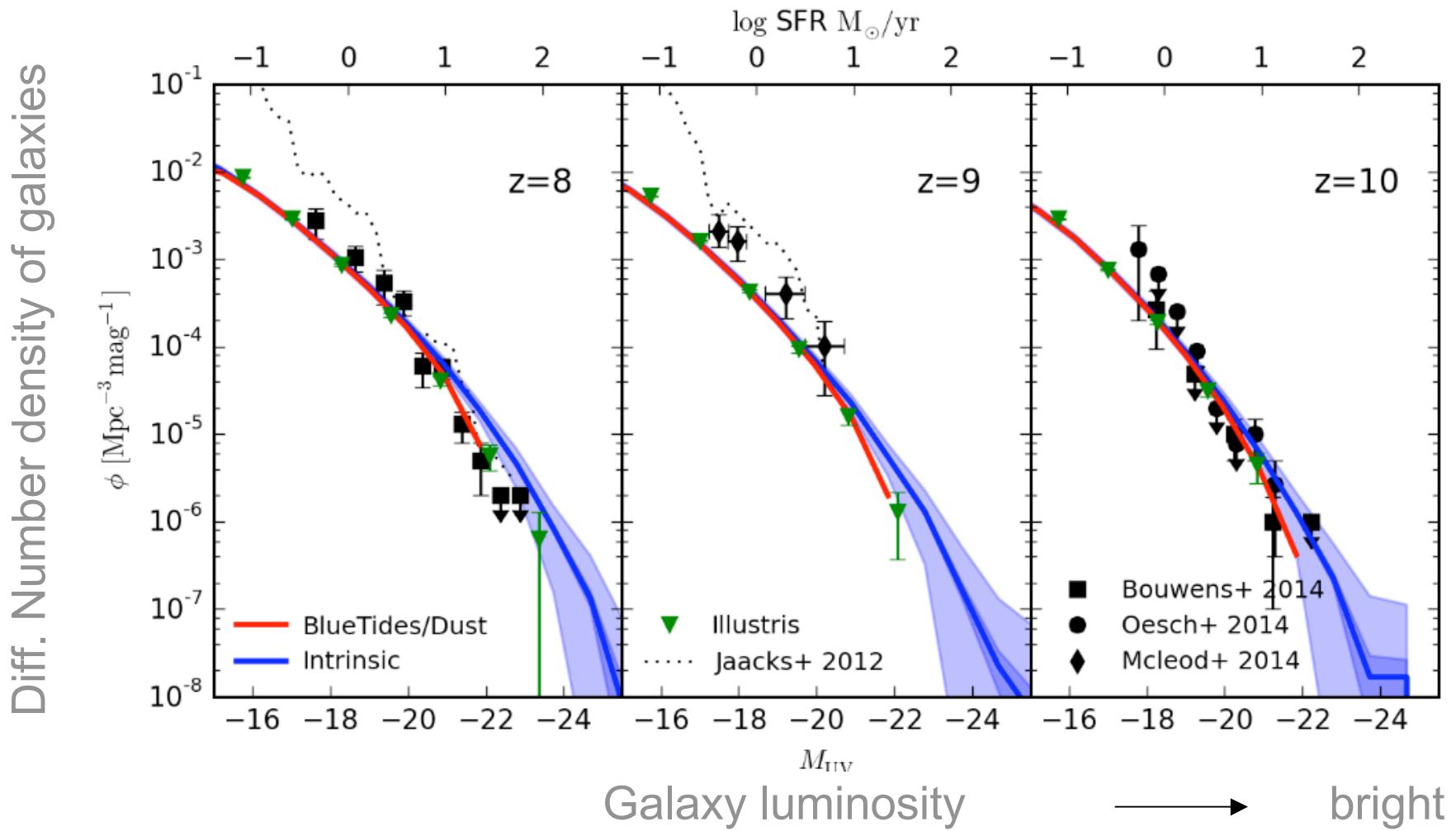
# BlueTides Simulation: Global SFRD is consistent with current observations.



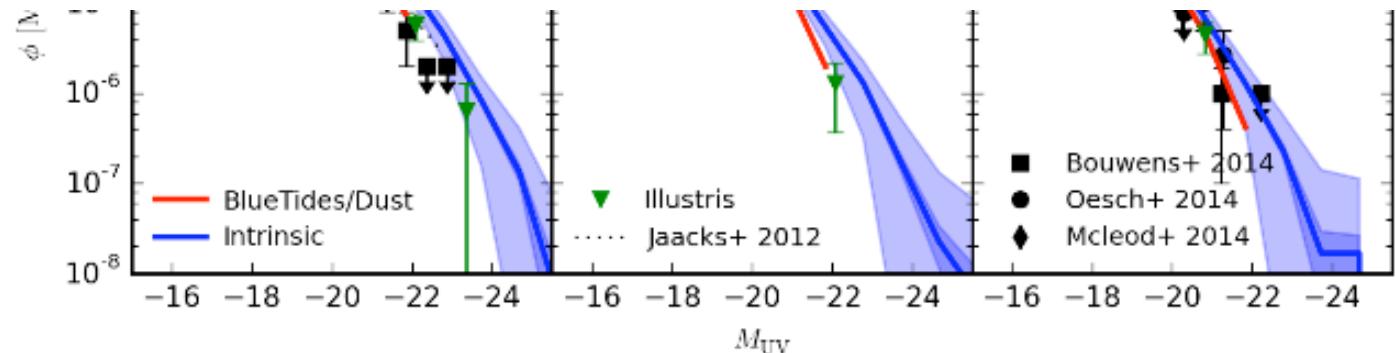
# Simulations like Observations: Create Mock Fields. Source extract detection to find galaxies



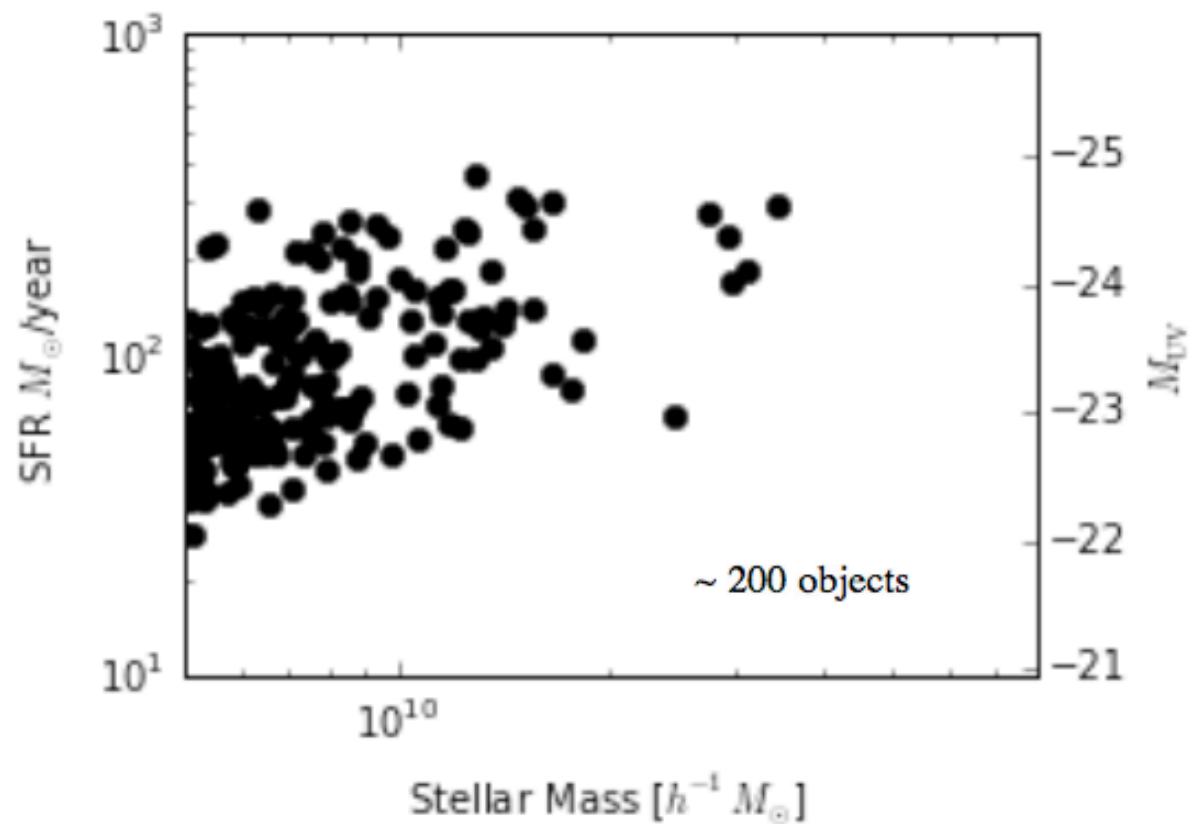
# Galaxy Luminosity Function in BlueTides consistent with Hubble Legacy Fields



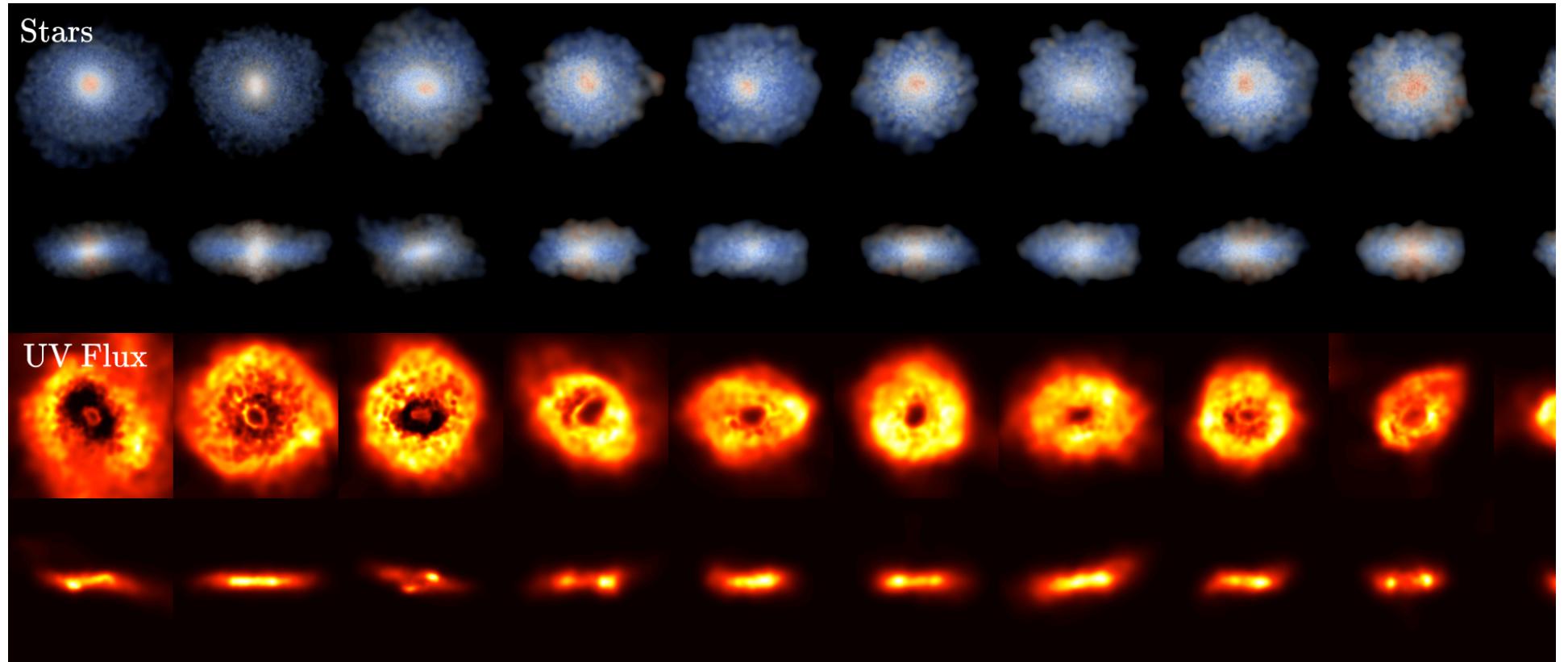
Feng et al., 2015a



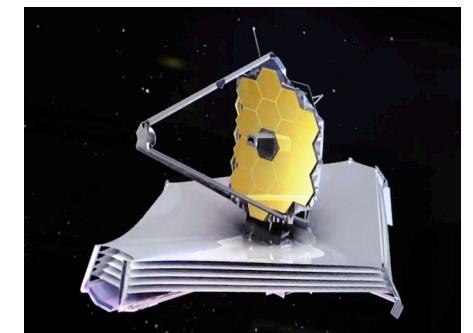
Predictions  
for the brightest  
Galaxies at  $z=8$ :



$z=8$  Milky Way (/Massive) galaxies are **disks!**

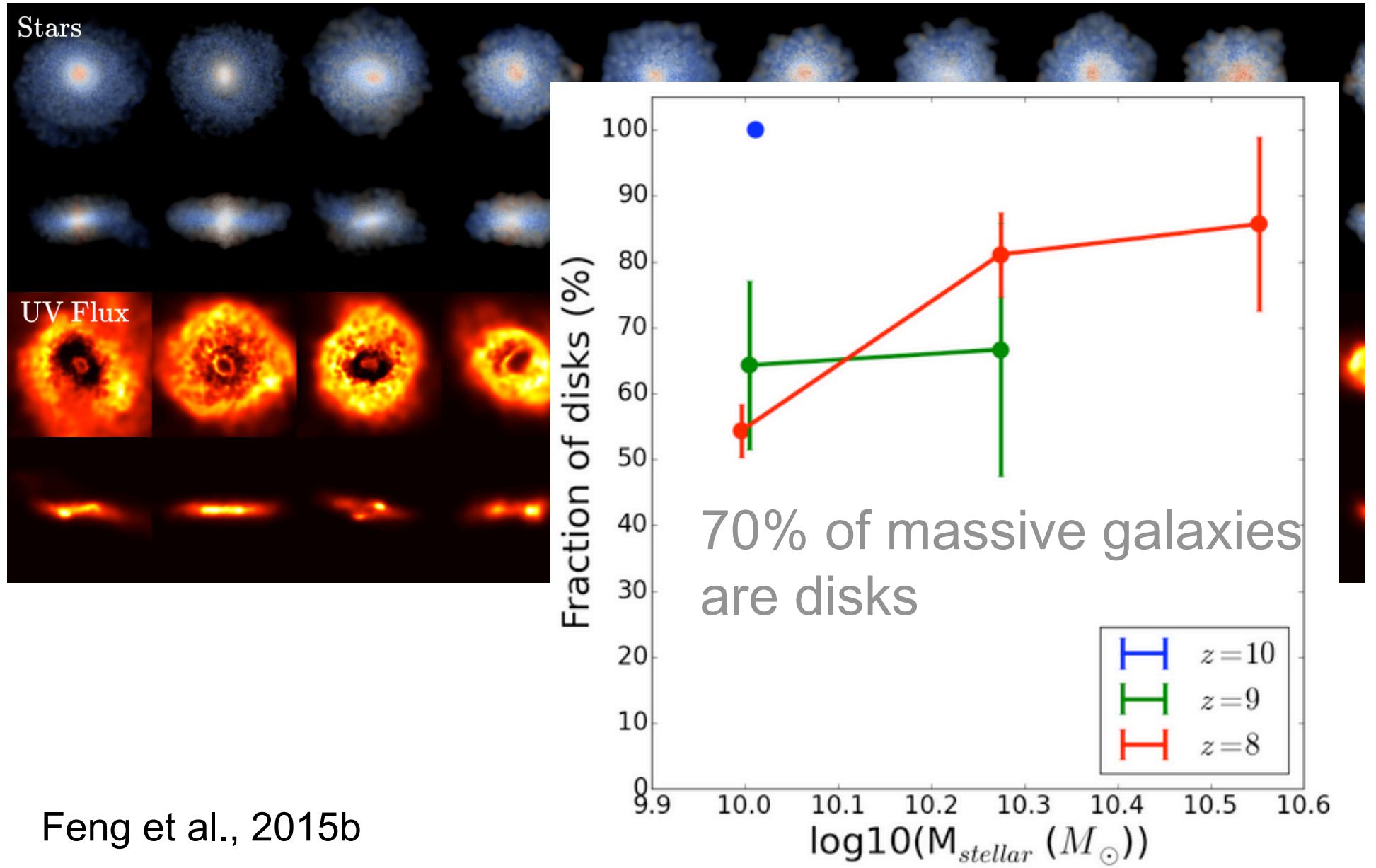


JWST

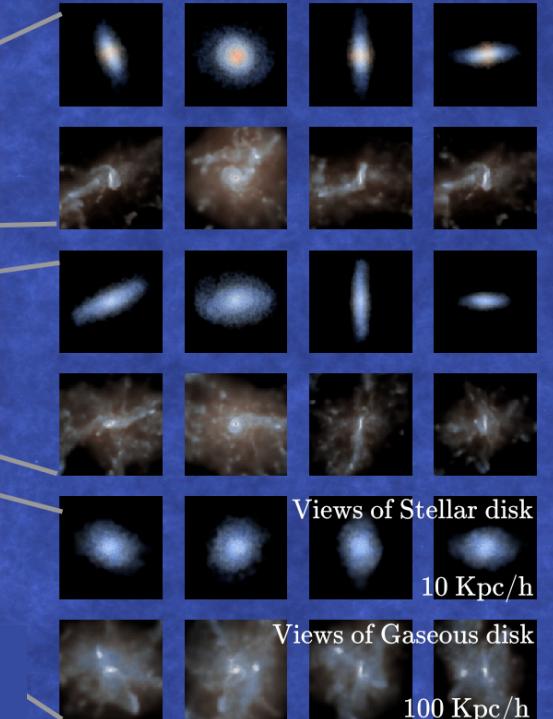
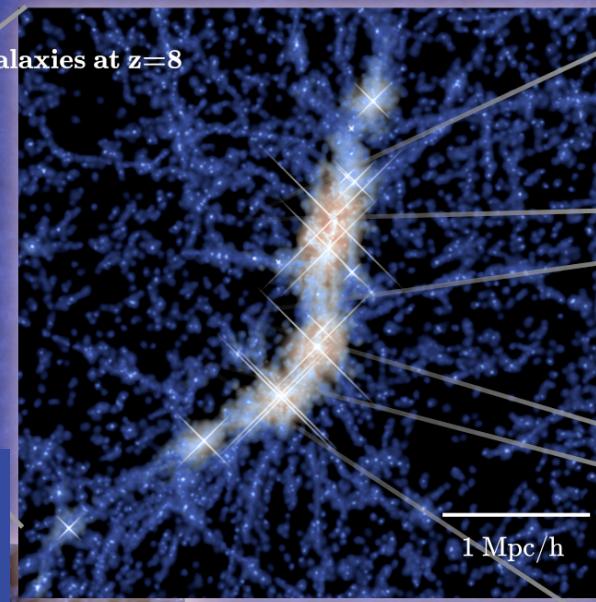
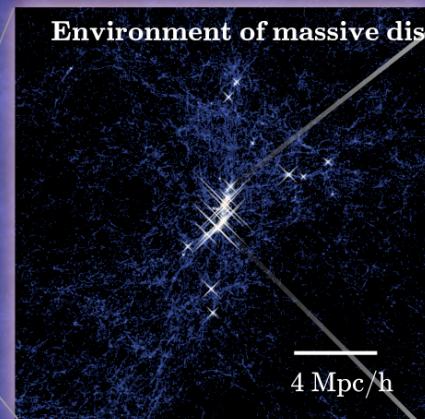


Feng et al., 2015b

# $z=8$ Milky Way (/Massive) Halos look like disks!



WFIRST  
should detect  
~ 8000 Milky Way  
mass disks  
at  $z=7-8$



The **BlueTides** Simulation

0.7 trillion particles  
0.65 million cores

**BLUE WATERS**  
SUSTAINED PETASCALE COMPUTING

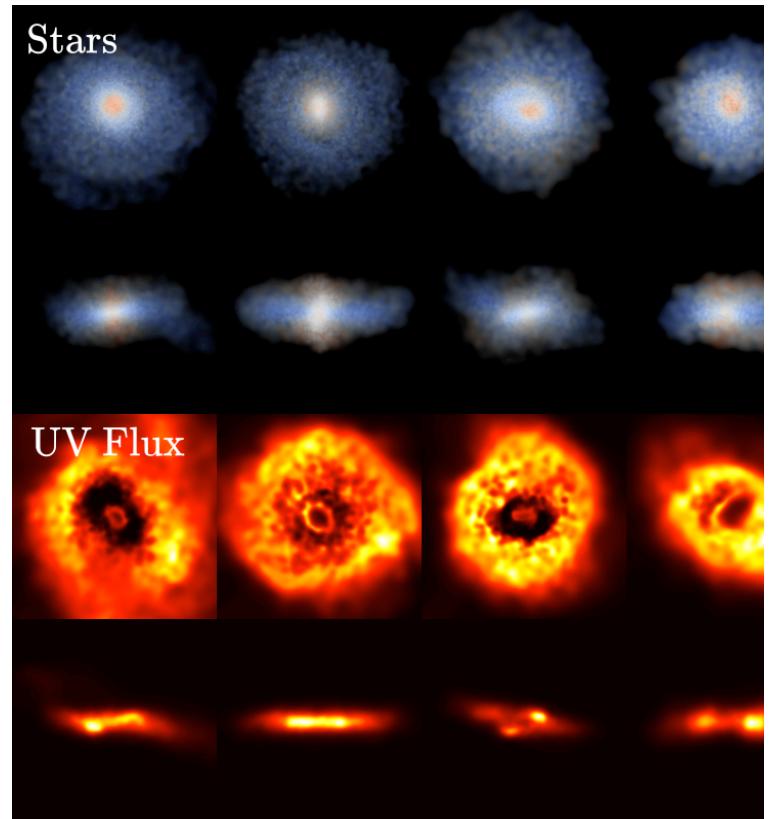
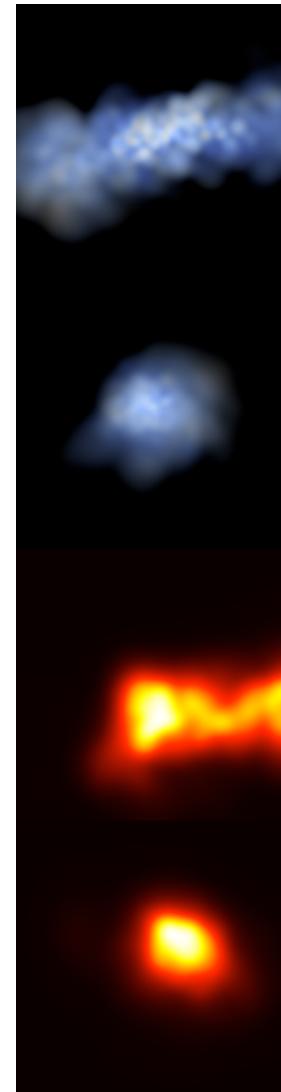
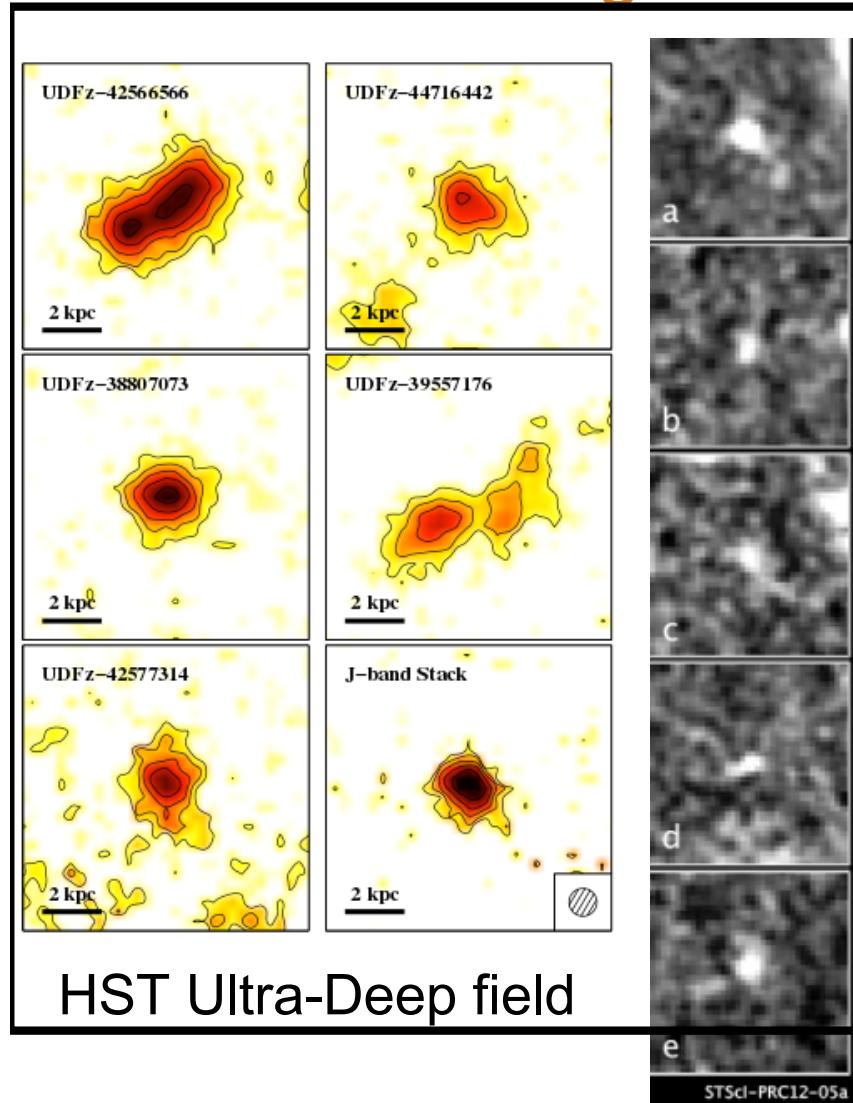


**bluebides**

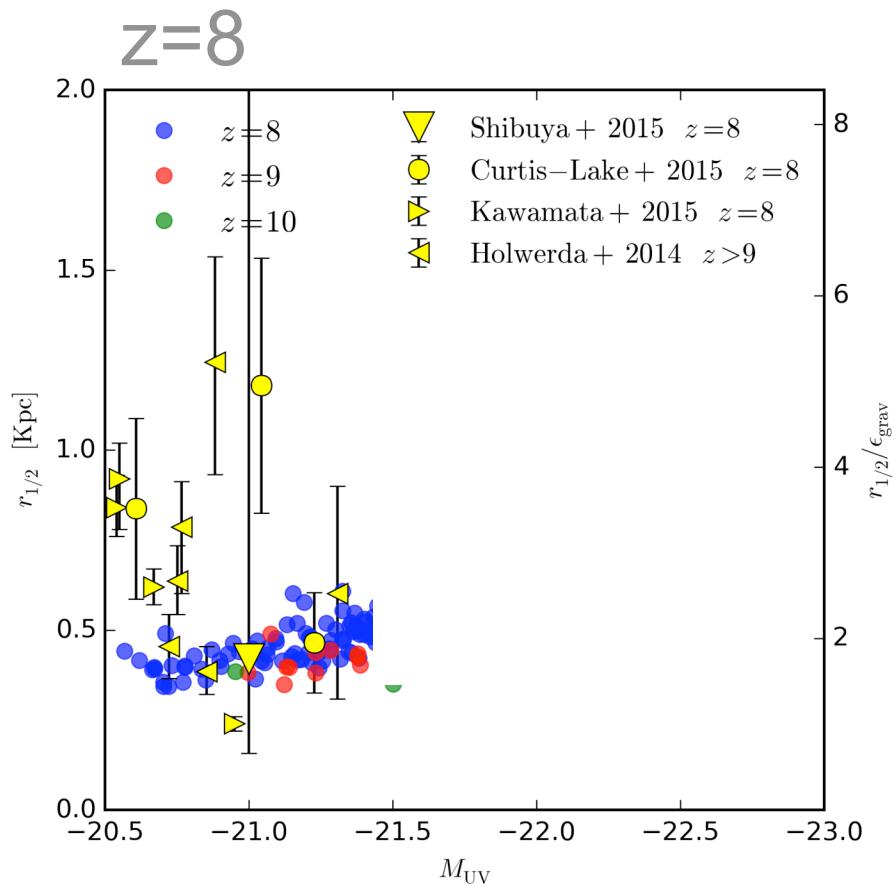
Feng et al. 2015

$z=8$  Milky Way (Massive) galaxies are **disks**!

But small faint galaxies are irregular ....

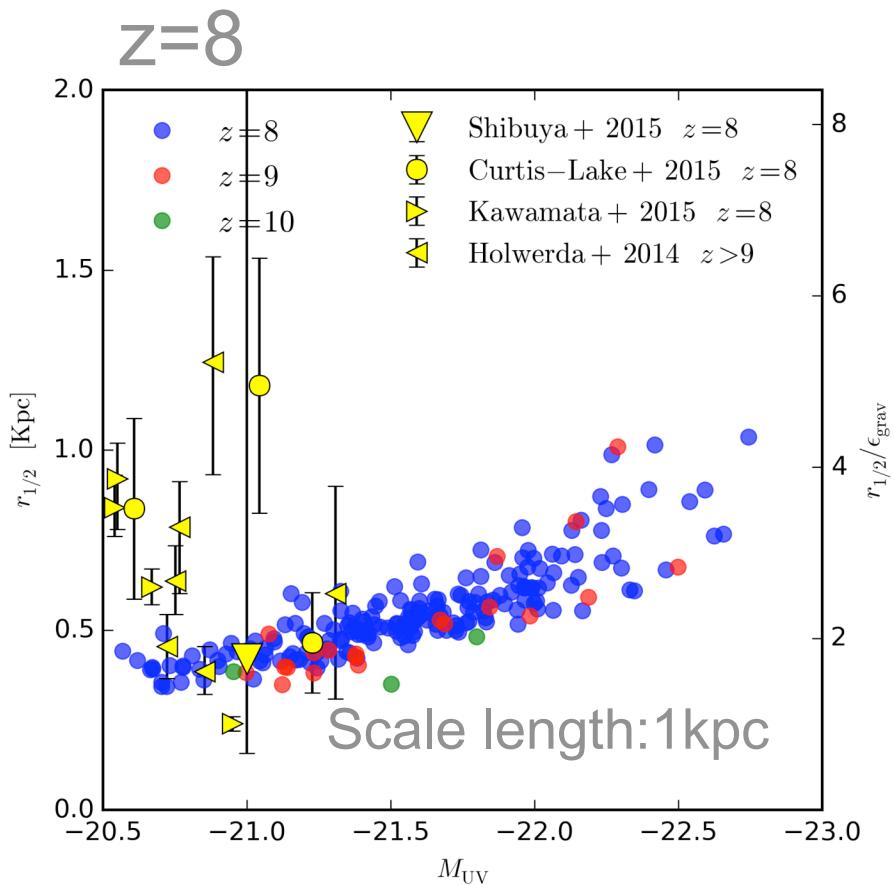


# The sizes of galaxies in BlueTides are consistent with HST observations

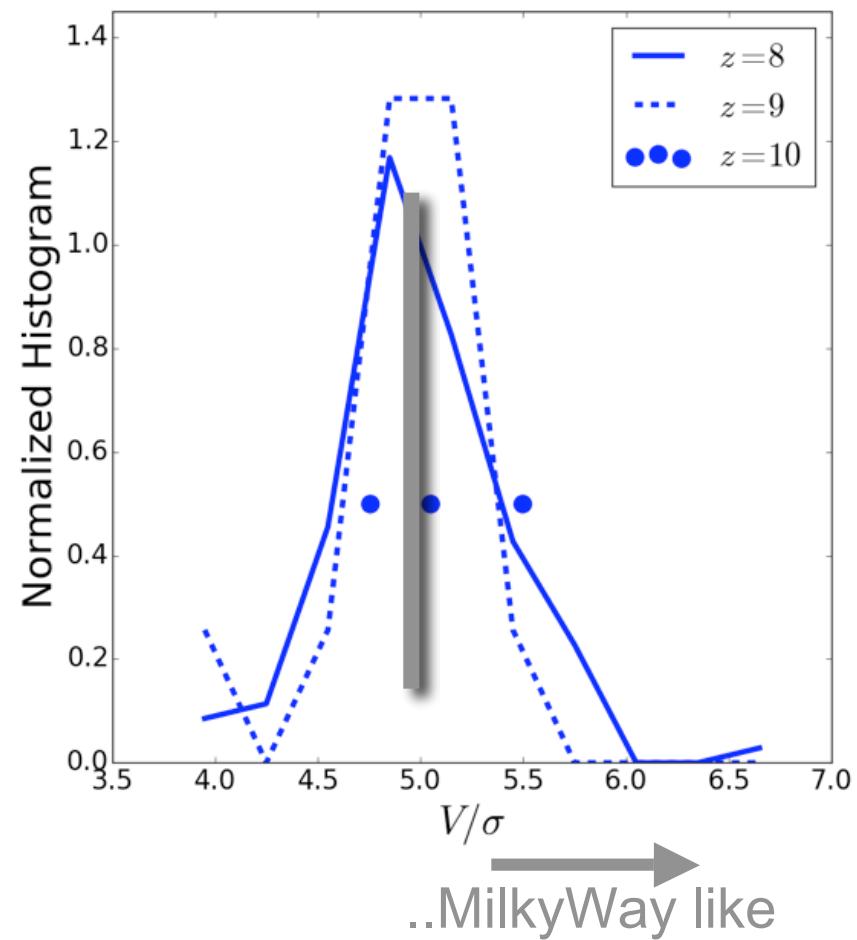
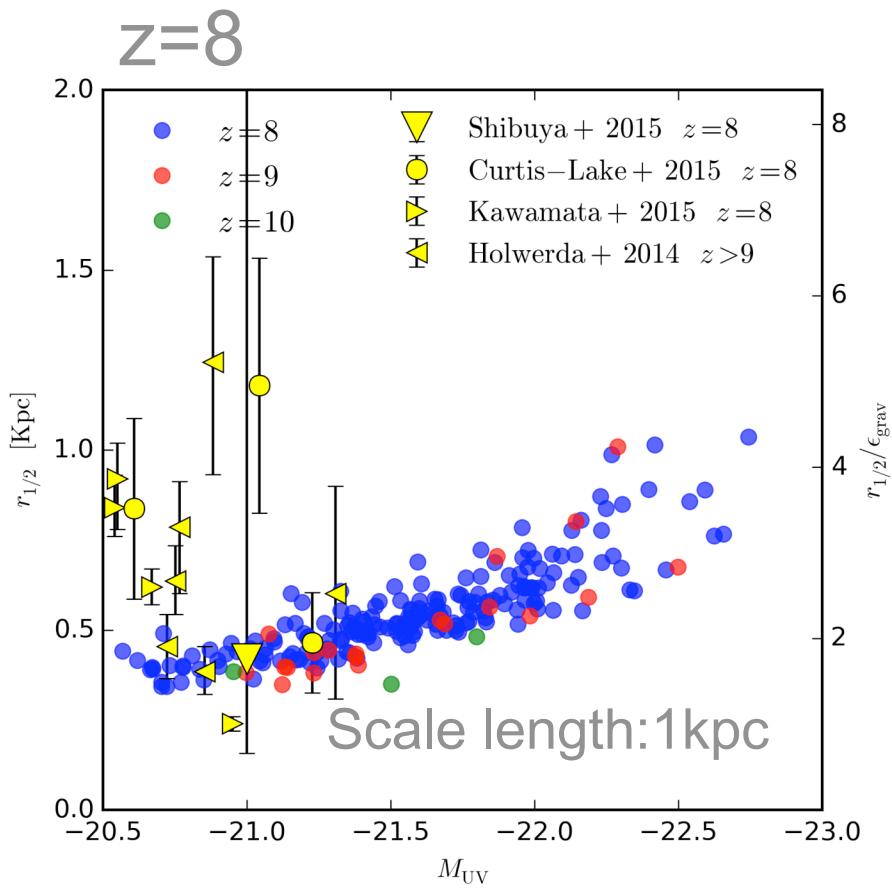


Feng et al., 2015a

The sizes of galaxies in BlueTides are consistent with HST observations --> ‘massive’ disks in bright galaxies are compact



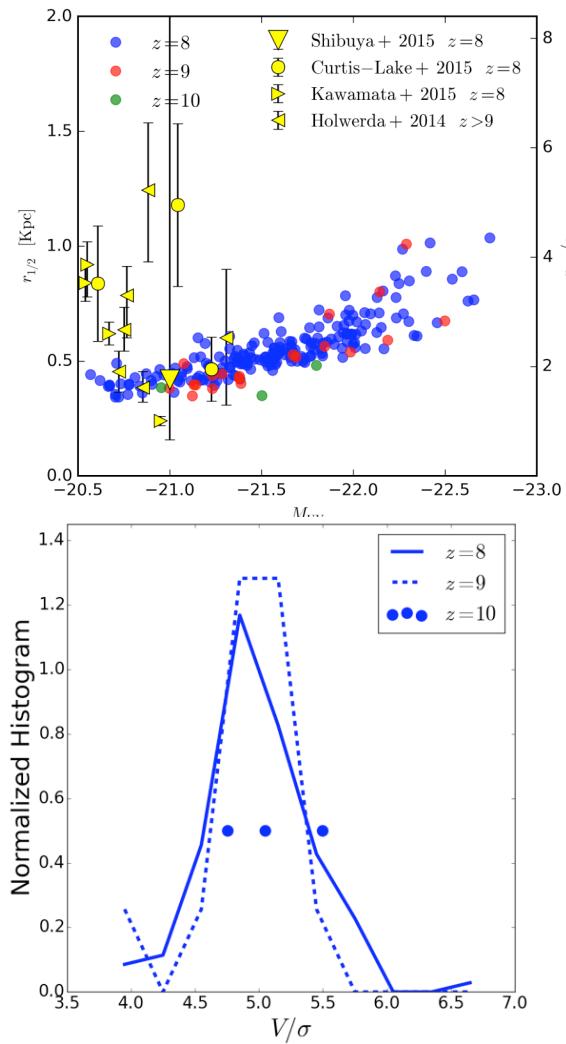
The sizes of galaxies in BlueTides are consistent with HST observations --> ‘massive’ disks in bright galaxies are compact rotationally supported



Feng et al., 2015a

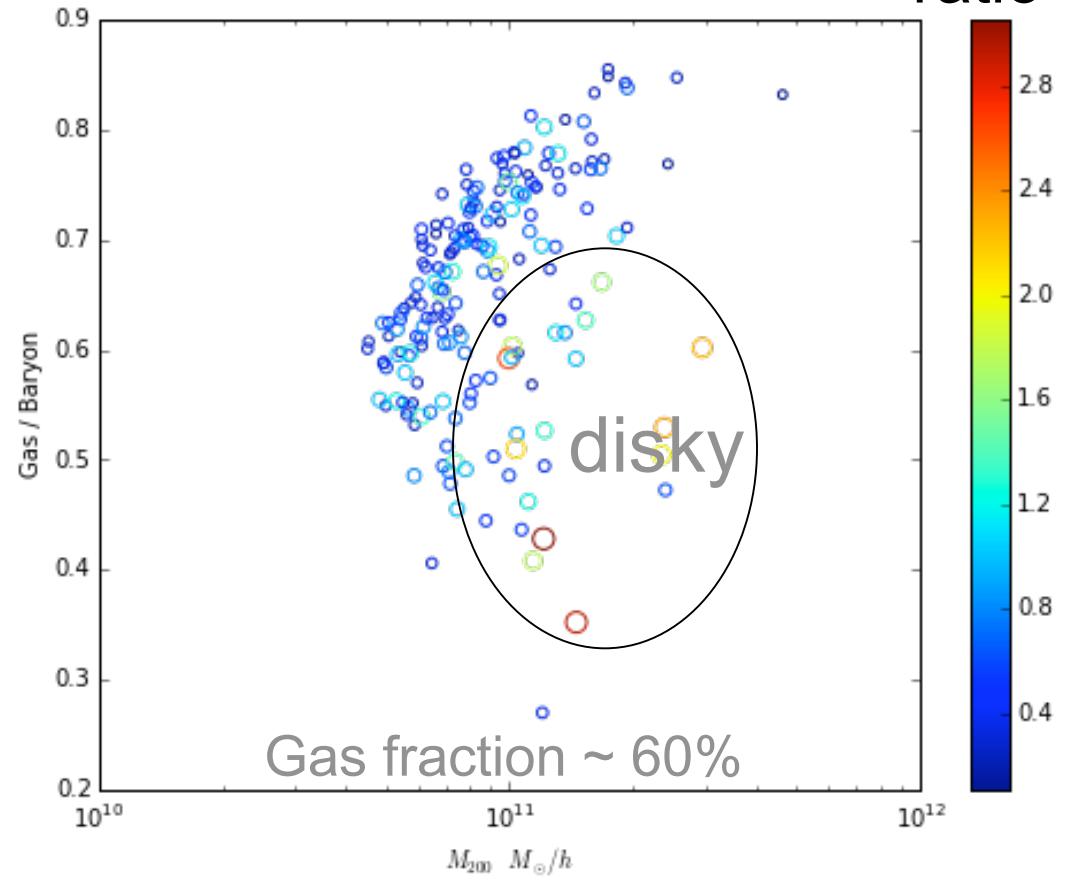
# Massive Disks at $z=8$ , more compact gas rich

$z=8$

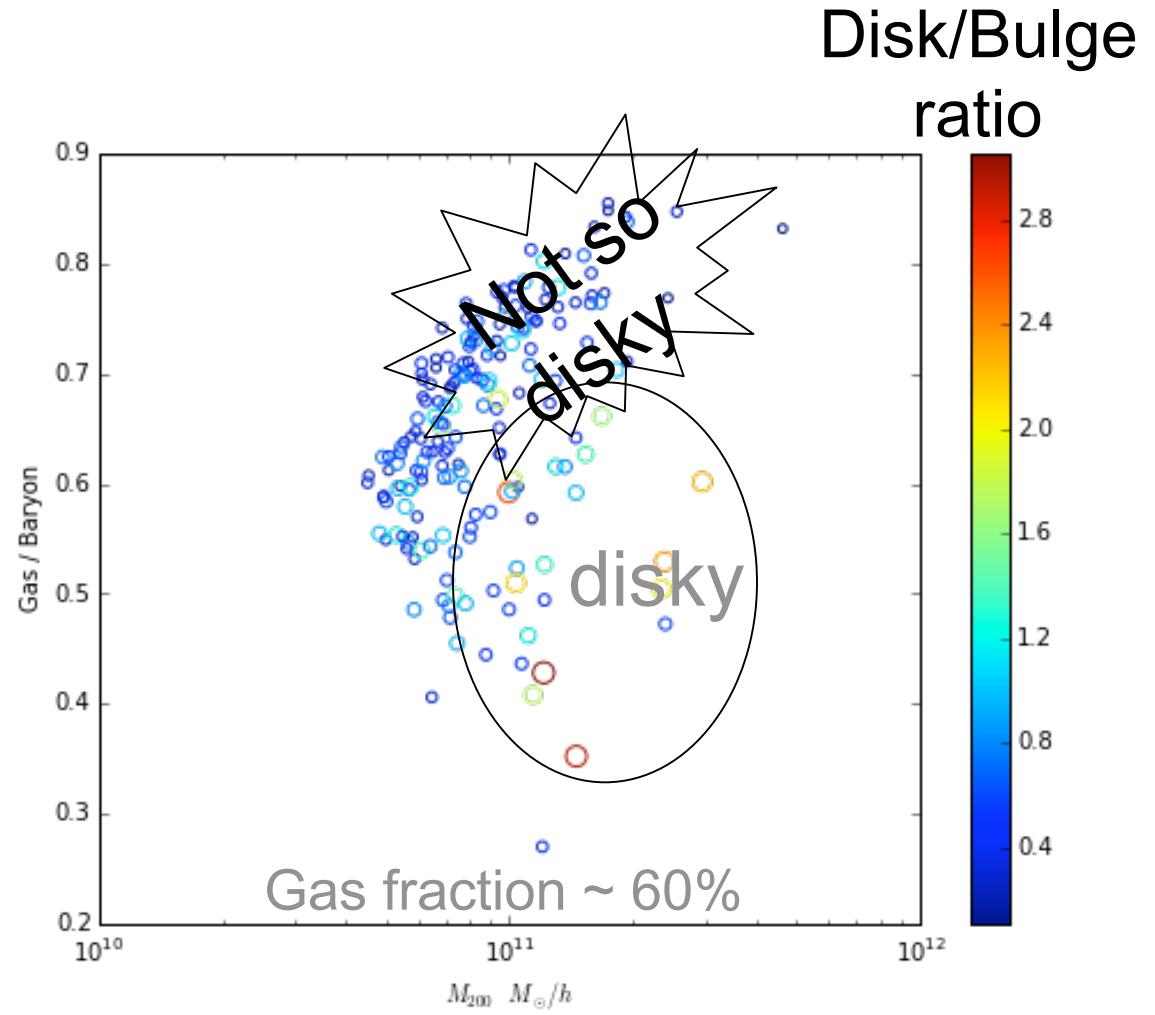
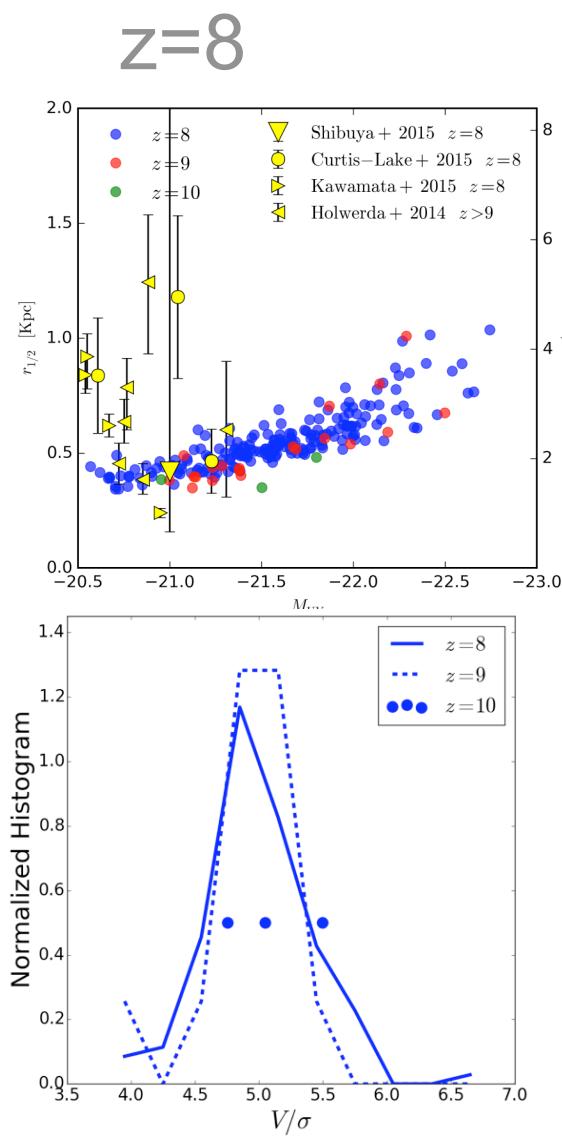


Feng et al., 2015a

Disk/Bulge  
ratio



# Gas richest systems have less dominant disk....



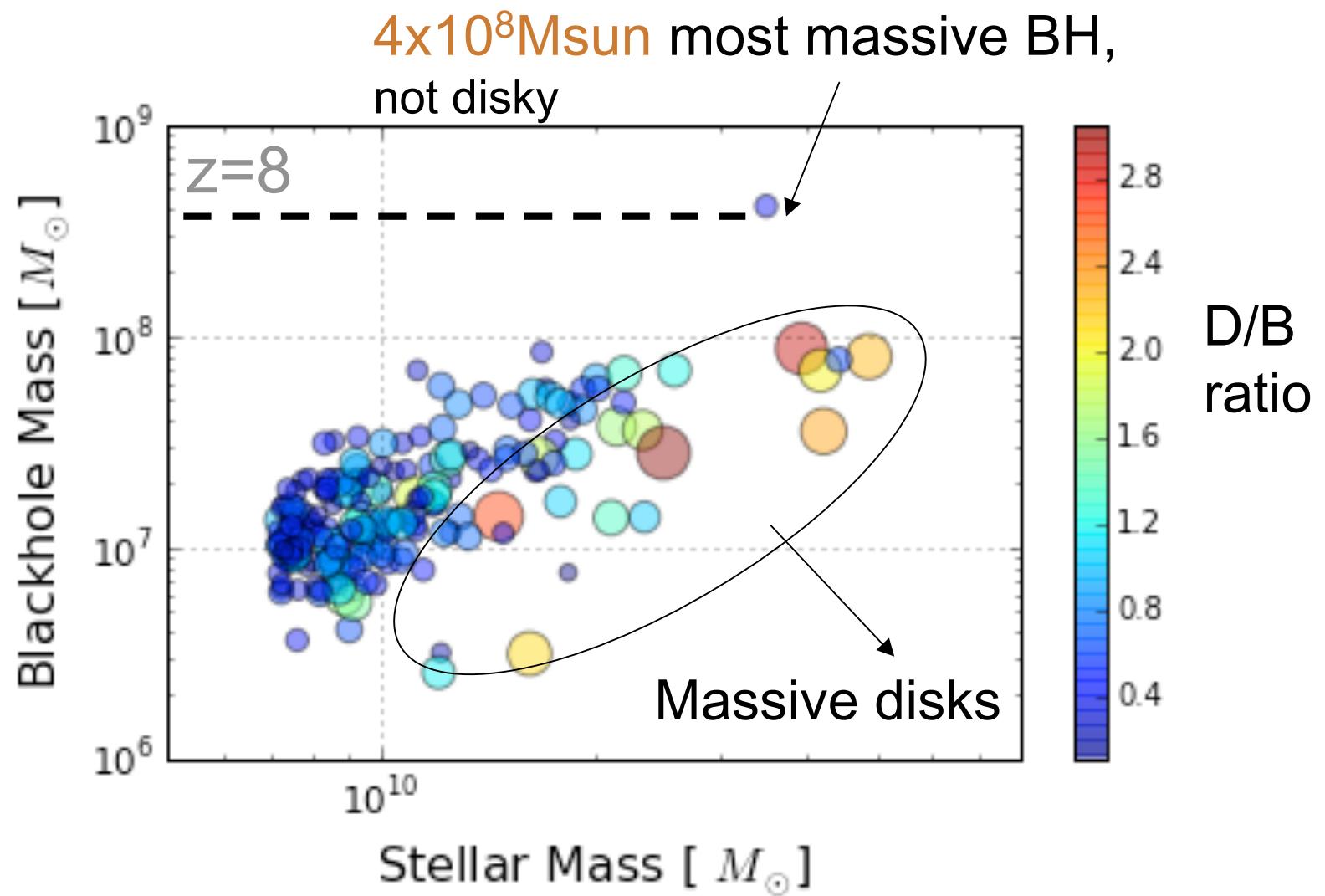
Feng et al., 2015a

The first 600 million years ( $z=7+$ ):

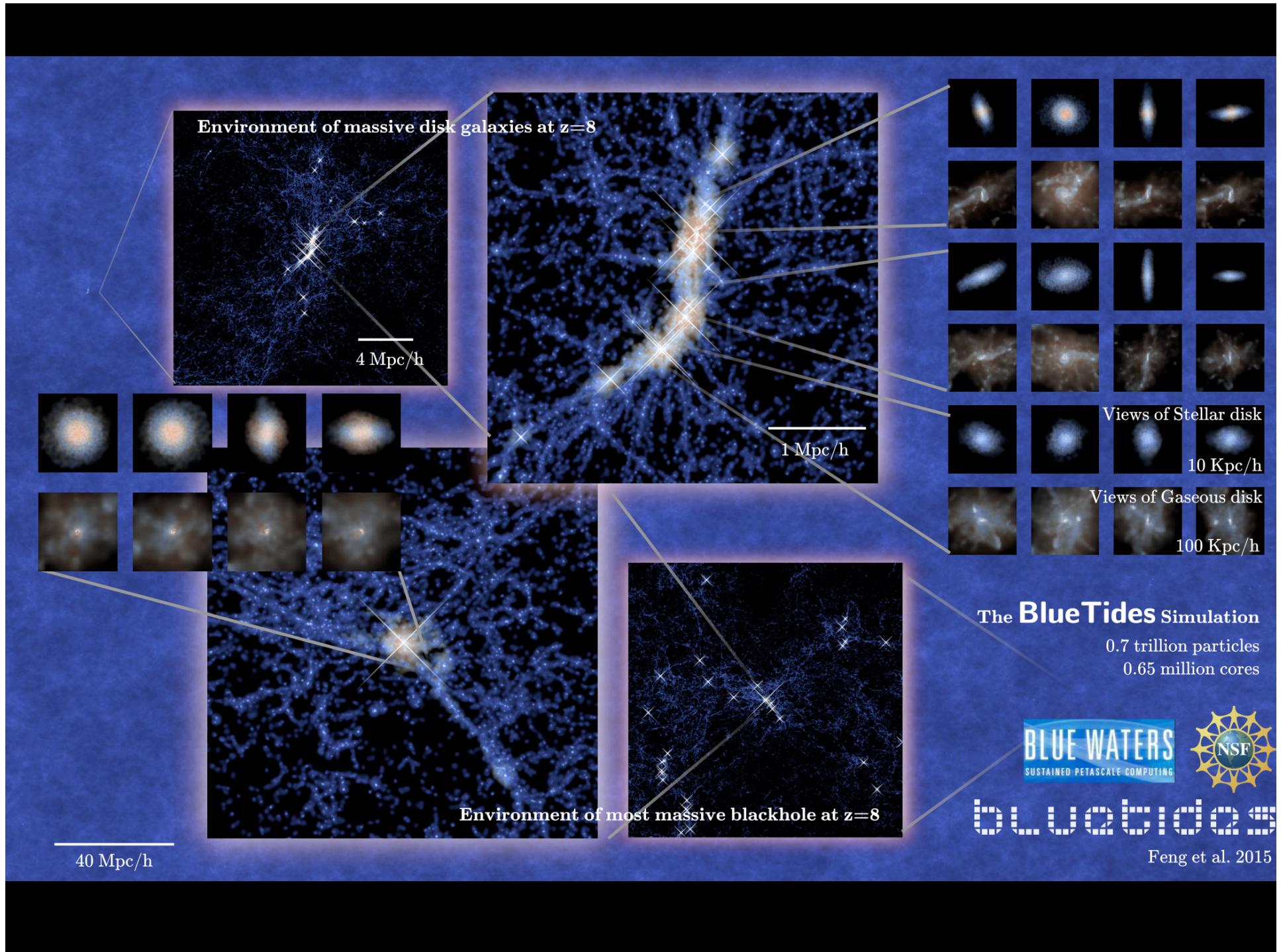
What are quasars like?

Predictions from  
BlueTides:

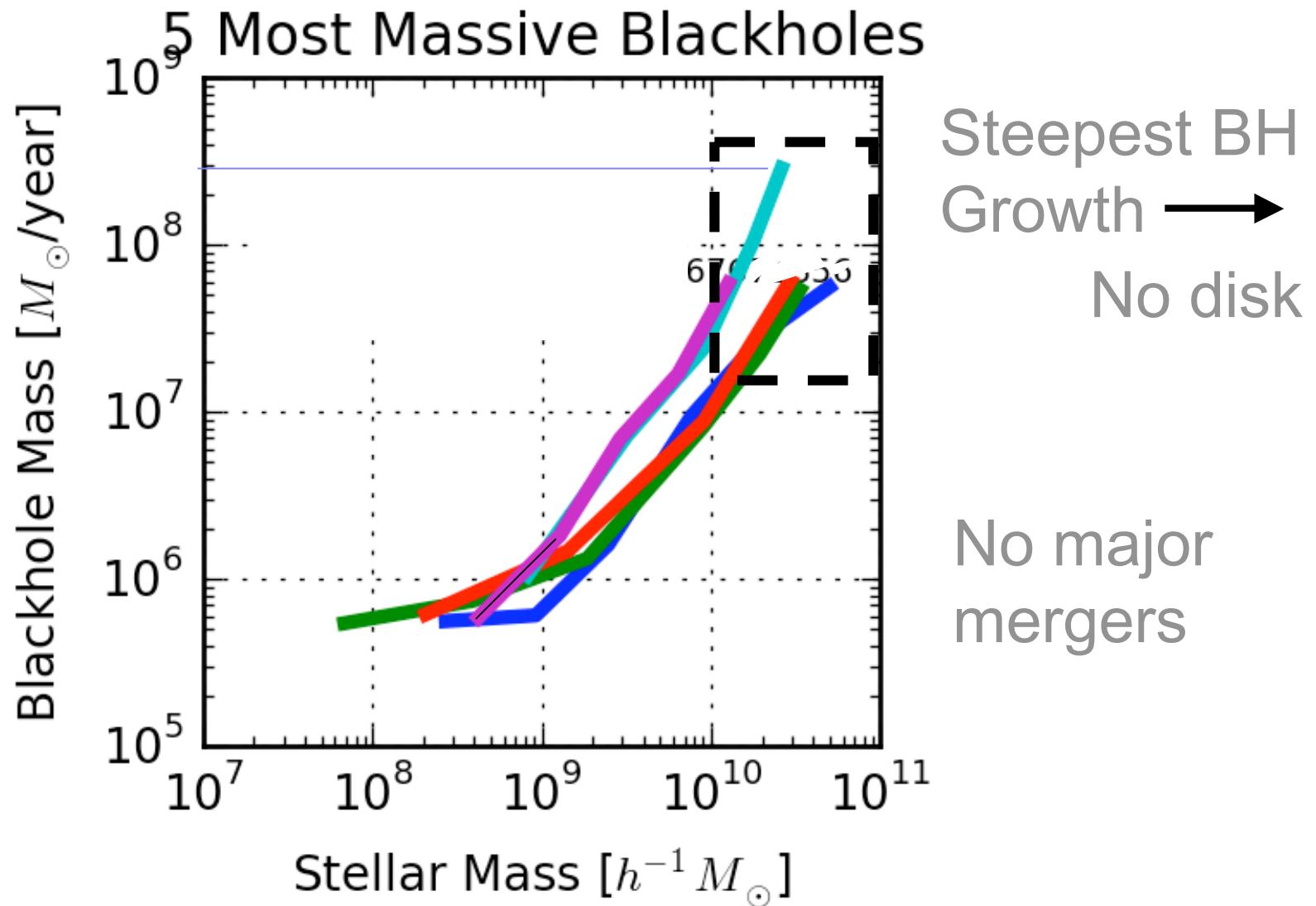
Most massive black holes at z=8, about  $10^8 M_{\odot}$



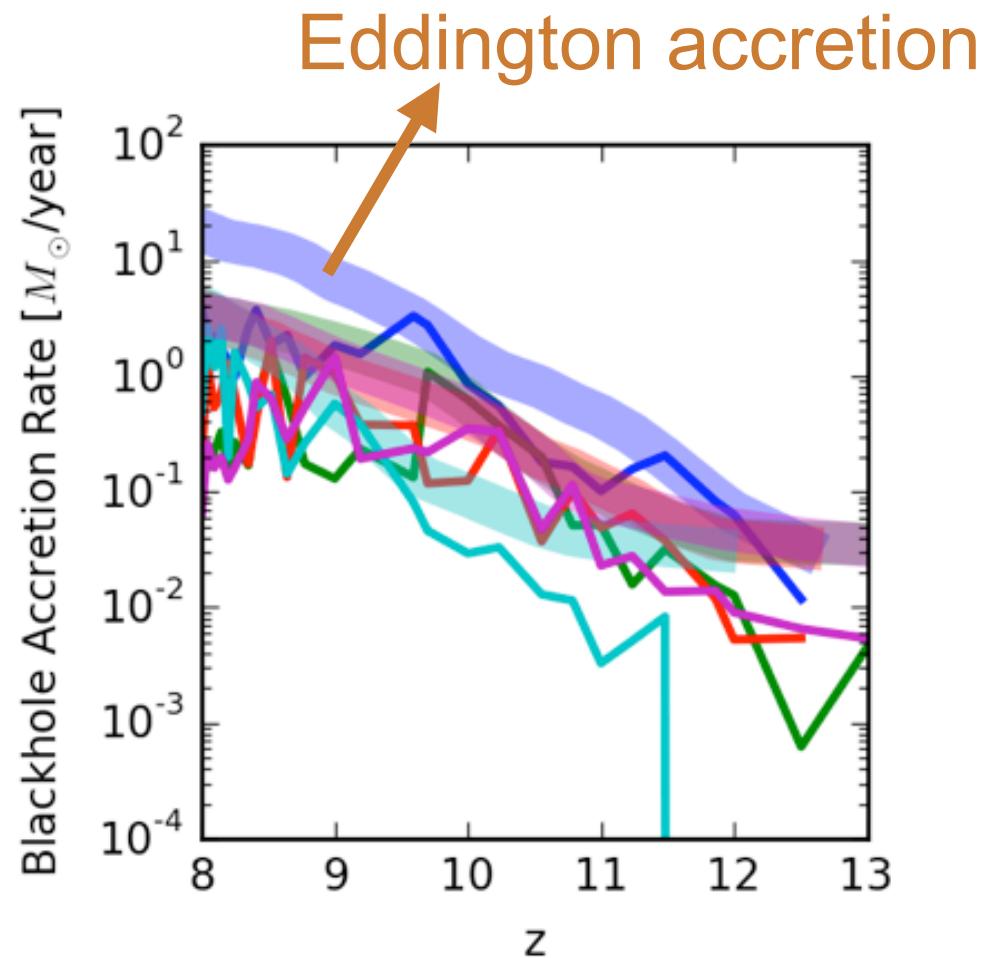
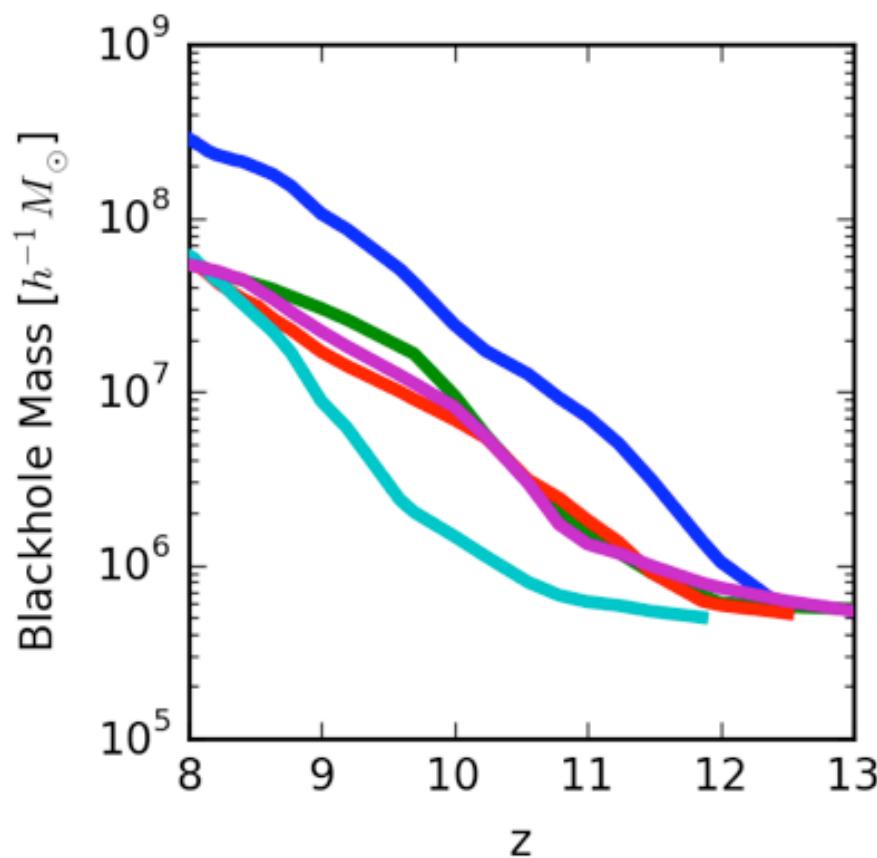
Fastest growing, massive black holes are not found in the massive disk galaxies!



# How do massive BH and galaxies grow?



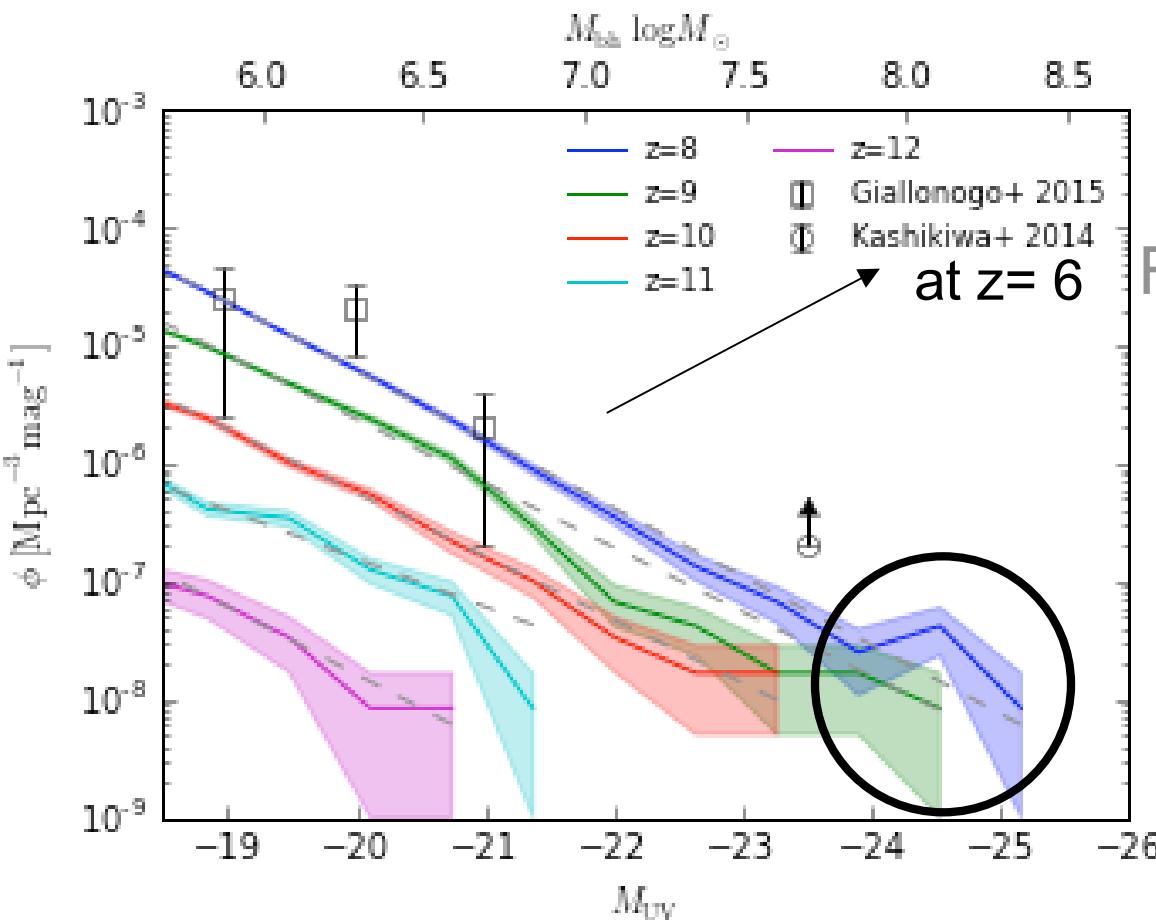
# Assembly of largest BHs at $z=8$ ,



# Black Hole/AGN Luminosity Function in BlueTides

Consistent → Predictions for first quasars

Diff. Number density of black Holes



First Quasars at  $z=8+$



WFIRST

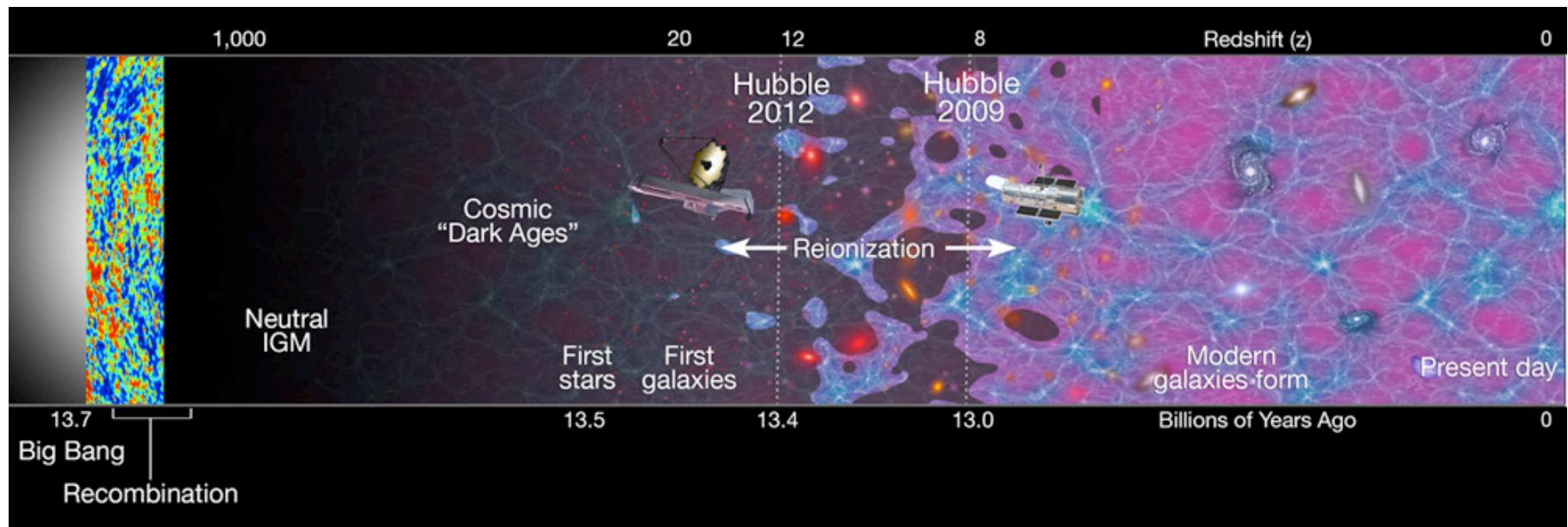


Euclid  
Deep  
Fields

Feng et al., 2015a

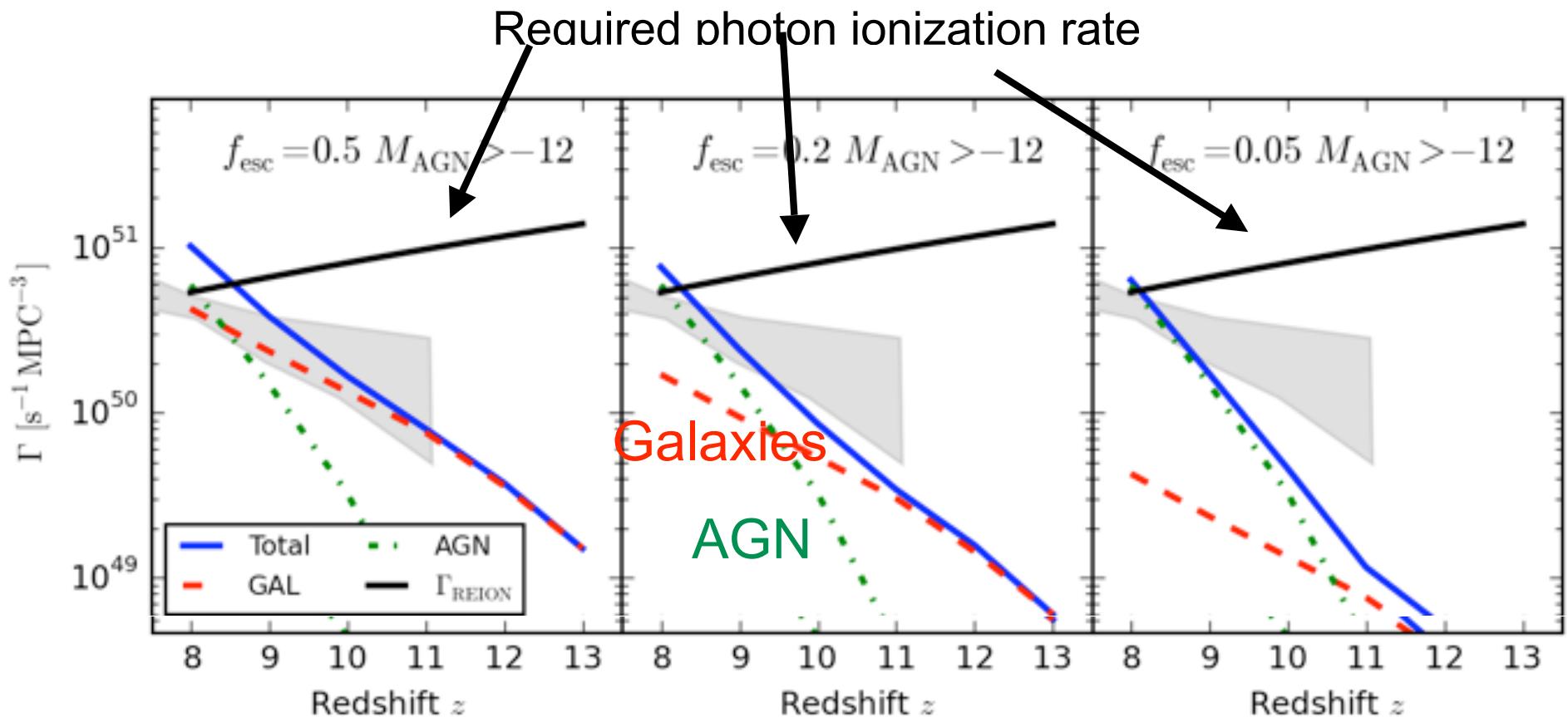
# What sources reionize the Universe?

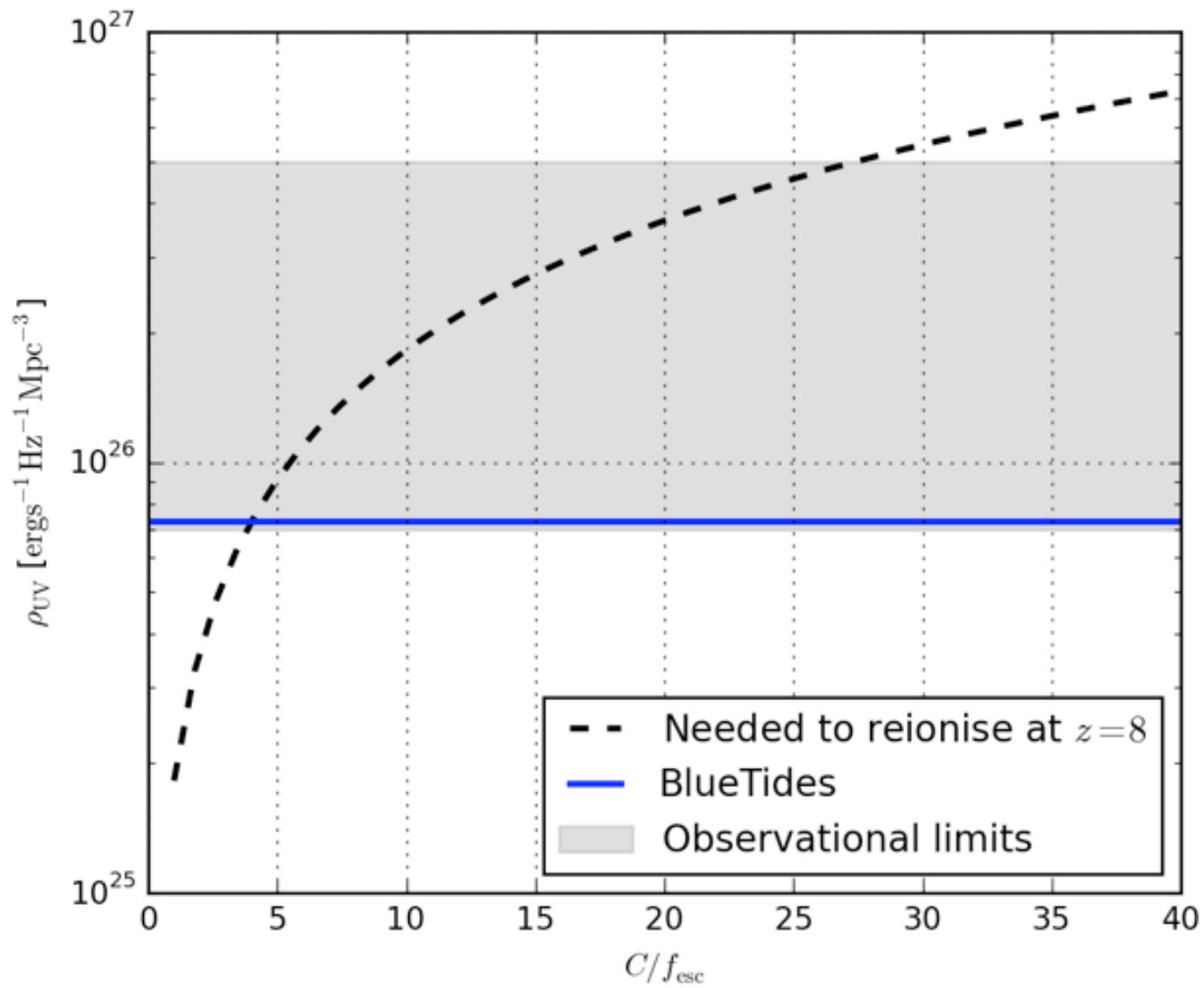
## Galaxies and AGNs in BlueTides



# BlueTides and Re-ionization history of the Universe

Galaxies can reionize the universe for high escape photon fractions. But AGNs can contribute (very?) significantly

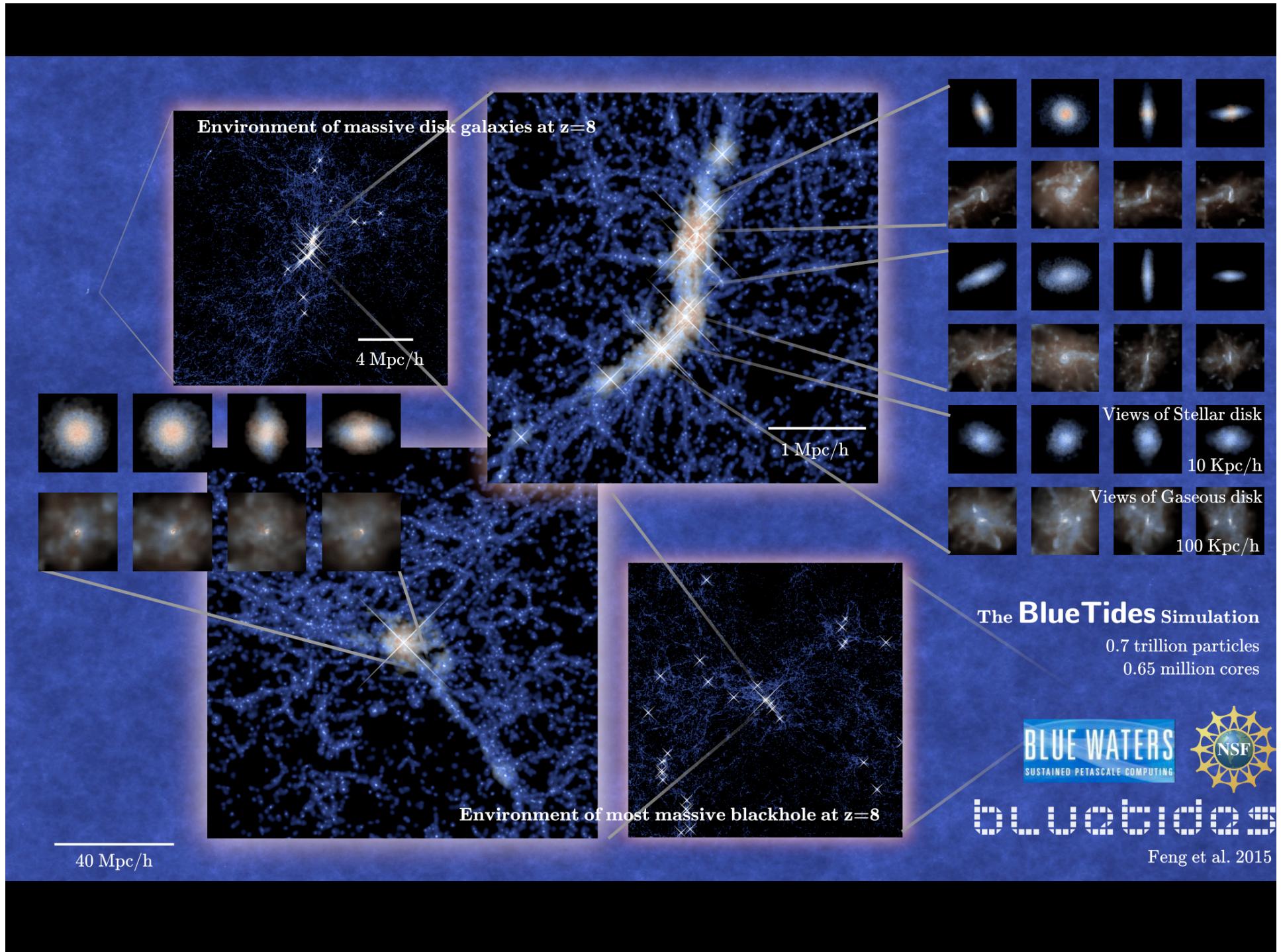




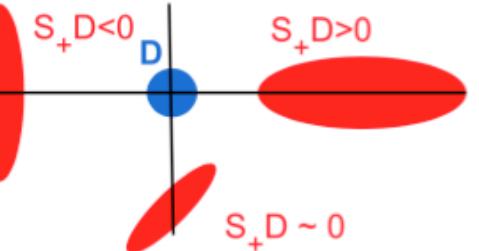
## Conclusions:

New large volume high res **BlueTides** Simulation predicts (at  $z>7$ ):

- Population of massive, compact disks --> **WFIRST**  
should detect  
thousands
- Most massive black holes,  $10^8$  Msun --> **WFIRST**  
tens of objects



# Intrinsic Alignments of Galaxies from full hydro (MBII) simulation - Effects of Baryon Physics.



## Projected-shape density ( $w_{\delta+}$ ) correlation

- $M > 10^{12} h^{-1} M_{\odot}$
- For stellar matter, larger ellipticities increase  $w_{\delta+}$  (Tenneti et al. 2015, arXiv:1505.03124)

